

Solar System Alerts in Year 1 of LSST

The Effects of Template Generation

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of EDINBURGH

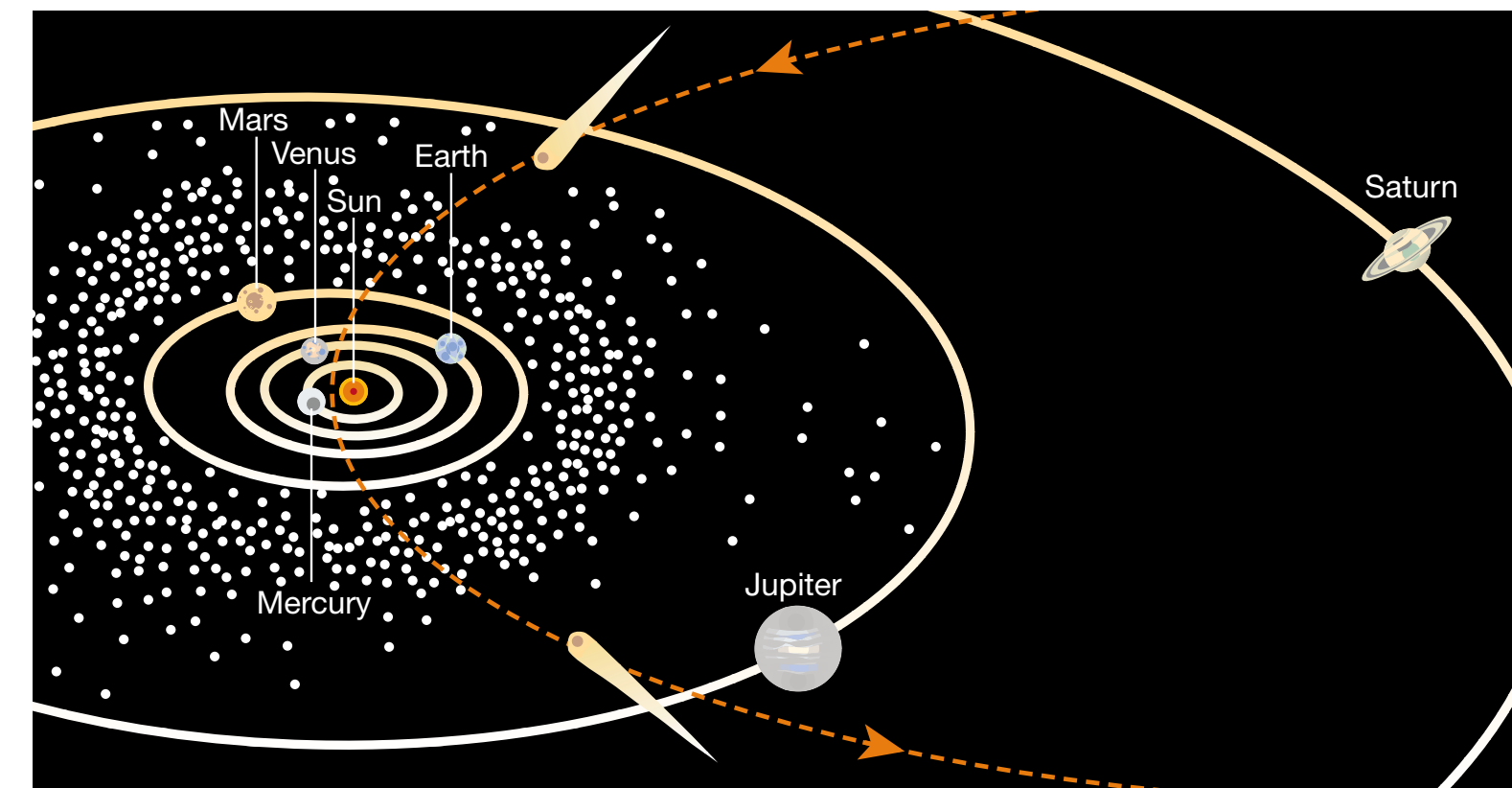


Rubin LSST Survey Strategy Optimization Focus Issue:
<https://doi.org/10.3847/1538-4365/adc307>



Solar System Science with LSST

- Depth and discovery - order of magnitude increase in known objects
- Number of observations - enables phase/light curve analysis
- Multiple filters *ugrizy* - colour characterisation
- Rapid "real-time" alerts (60s) - followup opportunities



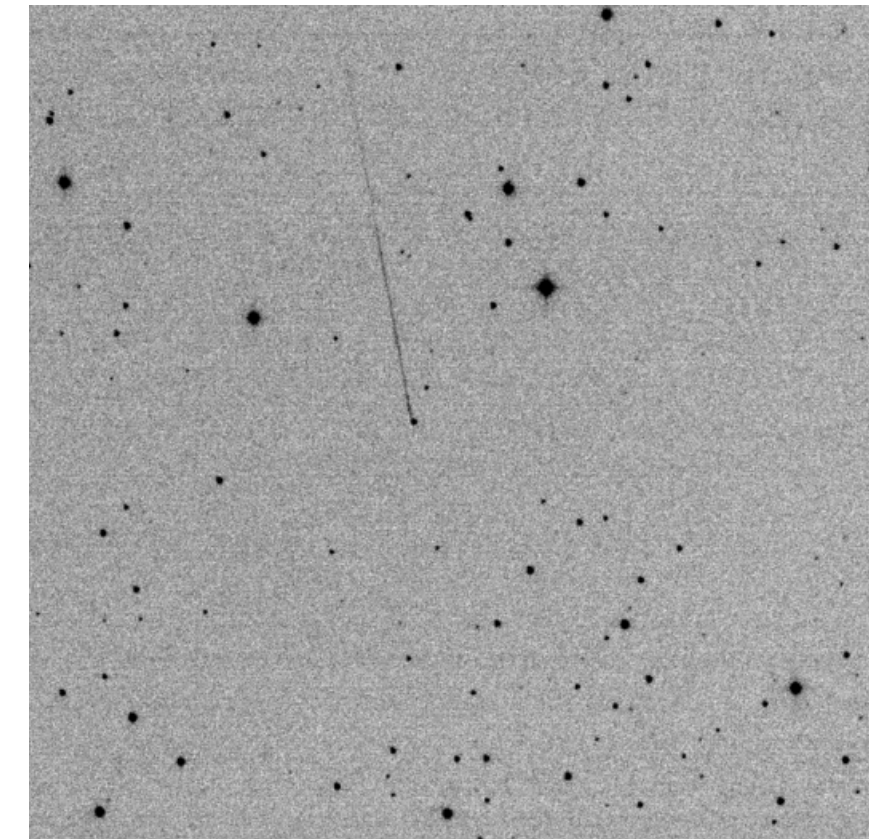
Traditional view
of asteroid and
comets
(Yeomans 2000)

	Currently Known	LSST Discoveries	Typical number of observations
Near Earth Objects (NEOs)	~20,000	200,000	(D>250m) 60
Main Belt Asteroids (MBAs)	~650,000	6,000,000	(D>500m) 200
Jupiter Trojans	~7000	280,000	(D>2km) 300
TransNeptunian Objects (TNOs) + Scattered Disk Objects (SDOs)	~3000	40,000	(D>200km) 450
Comets	~3000	10,000	?
Interstellar Objects (ISOs)	2	10	?

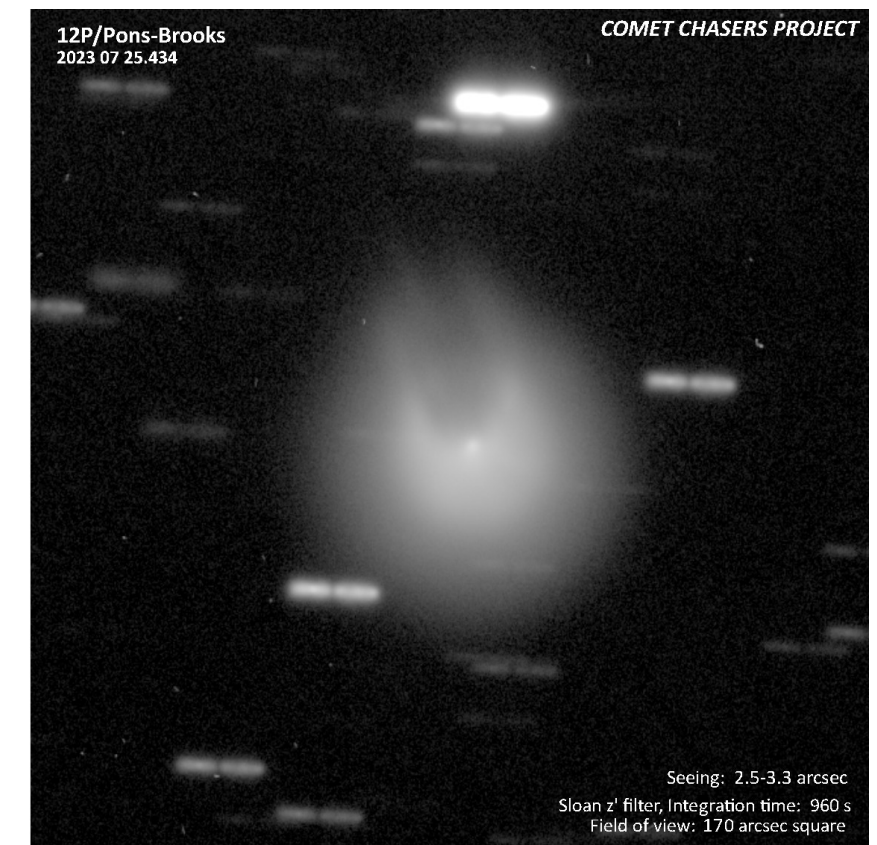
Rapid Solar System Science

Certain Solar System Objects (SSOs) will benefit from rapid discovery and followup:

- Interstellar Objects (ISOs) - limited observation window
- Near-Earth Objects (NEOs) - impact hazard and characterisation of small SSOs
- Active Objects - transient periods of activity/outburst

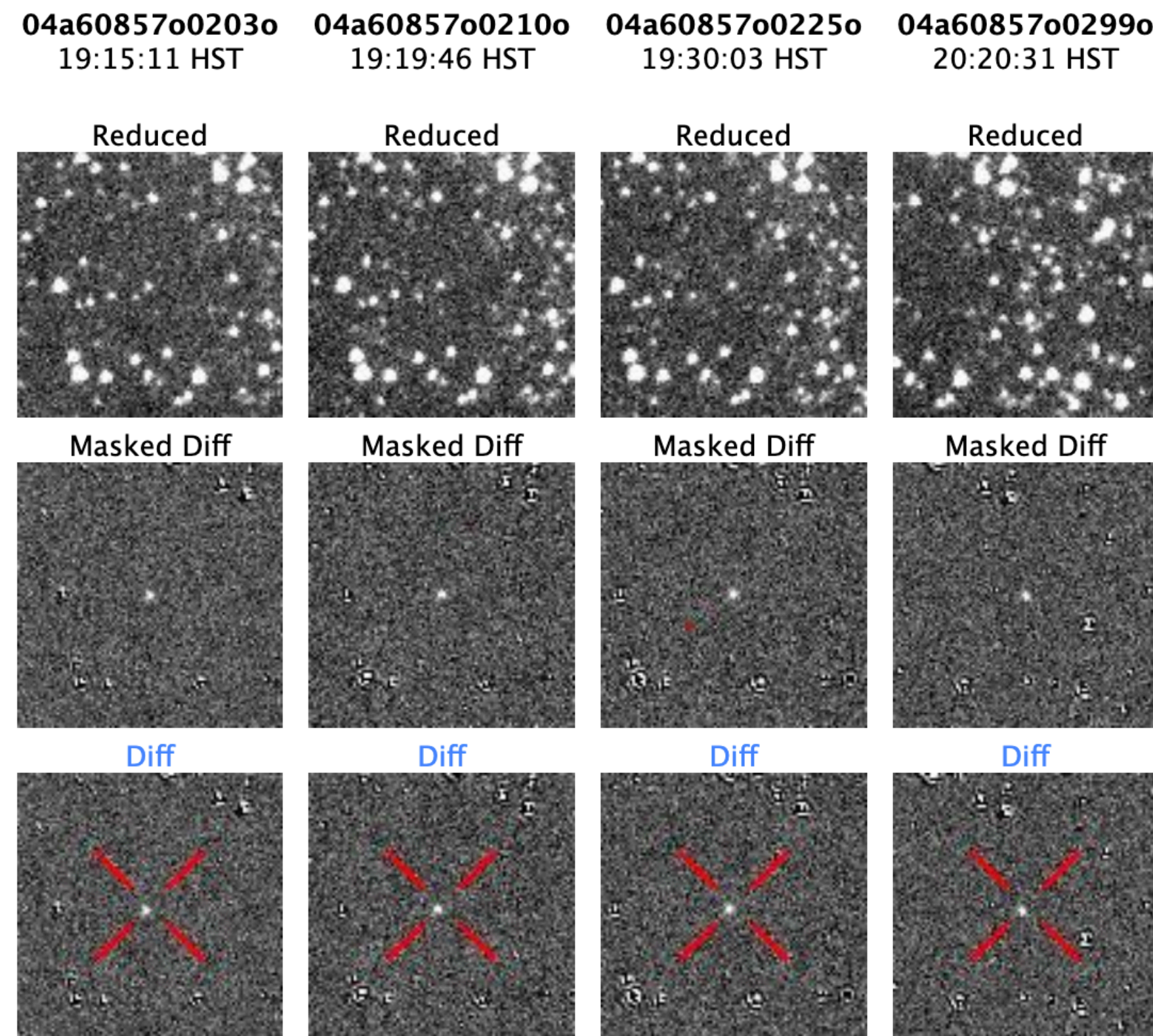


Final detection of impacting NEO 2024 BX1
L. Buzzi, G. V. Schiaparelli Observatory

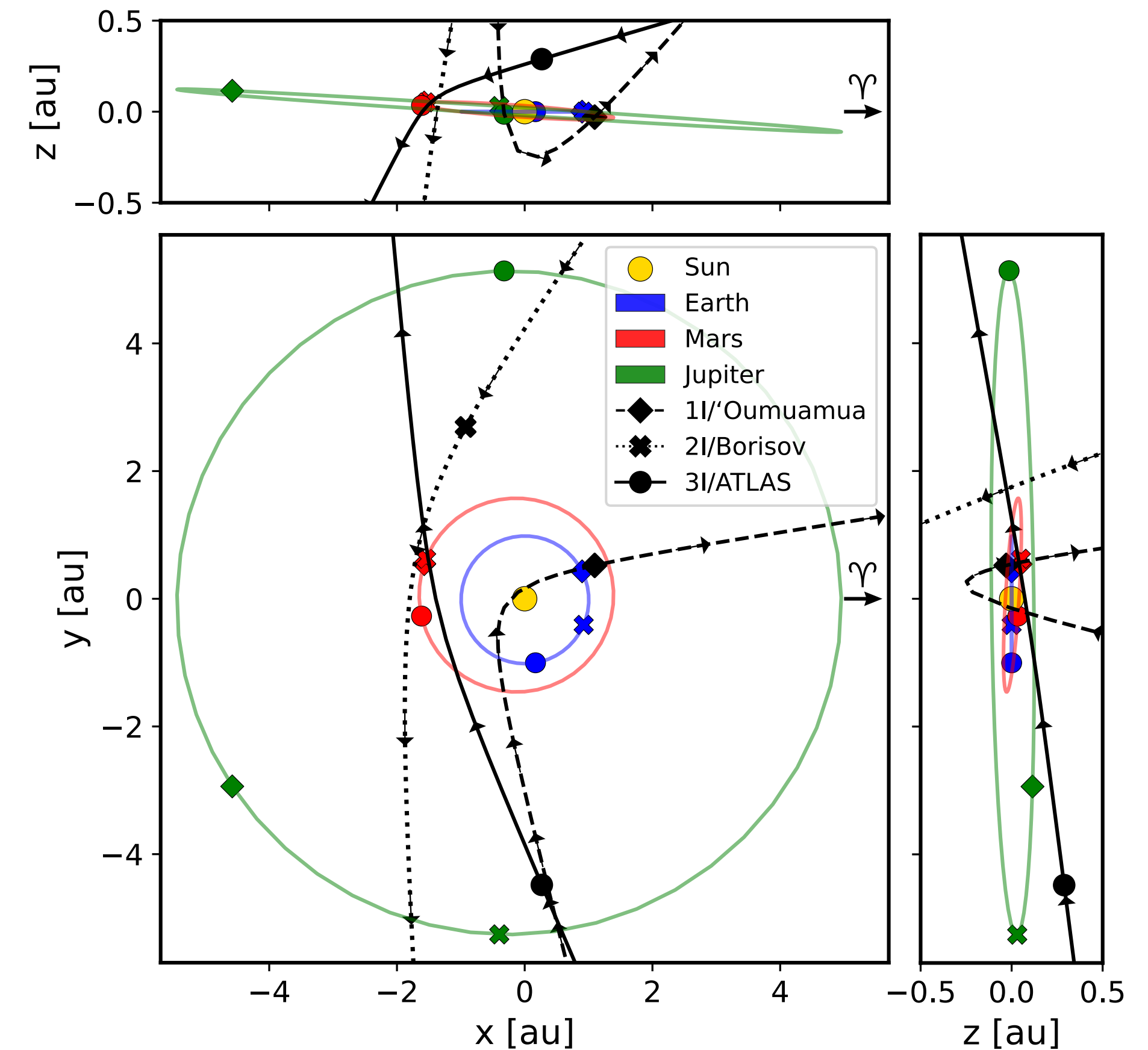


Comet 12P outburst
Comet Chasers/LCO

3I/ATLAS = C/2025 N1 (ATLAS)



ATLAS discovery cutouts



Seligman et al. 2025

See also Dorsey et al. 2025 for LSST ISO predictions

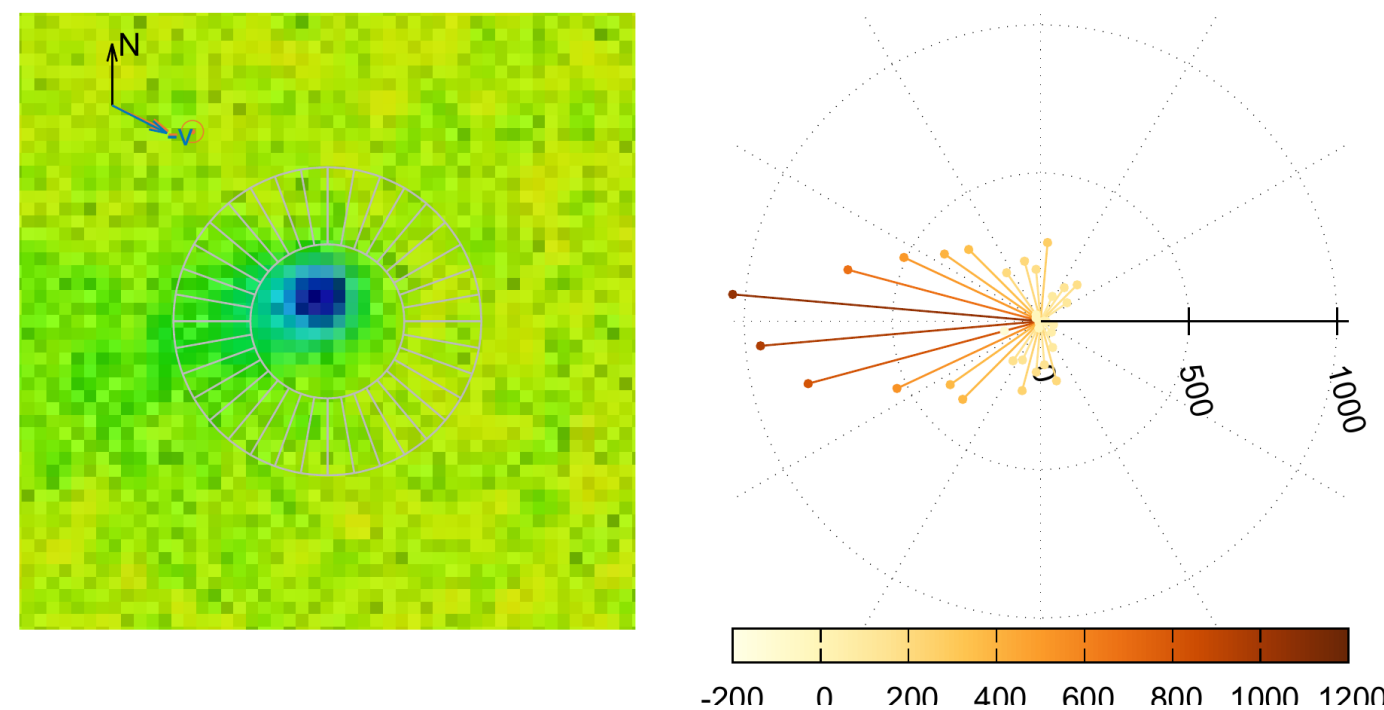
SSO Activity

Adler GitHub:

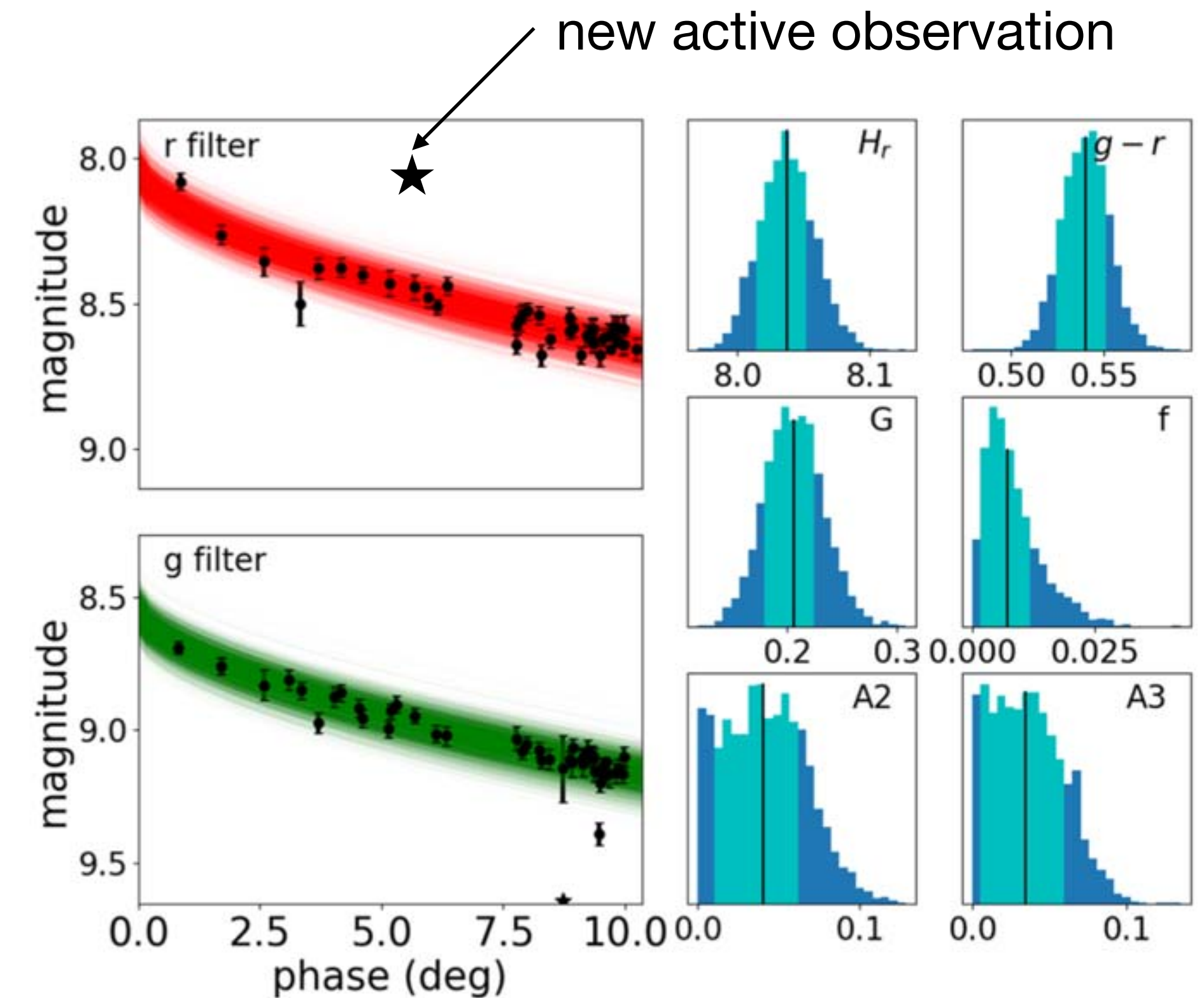
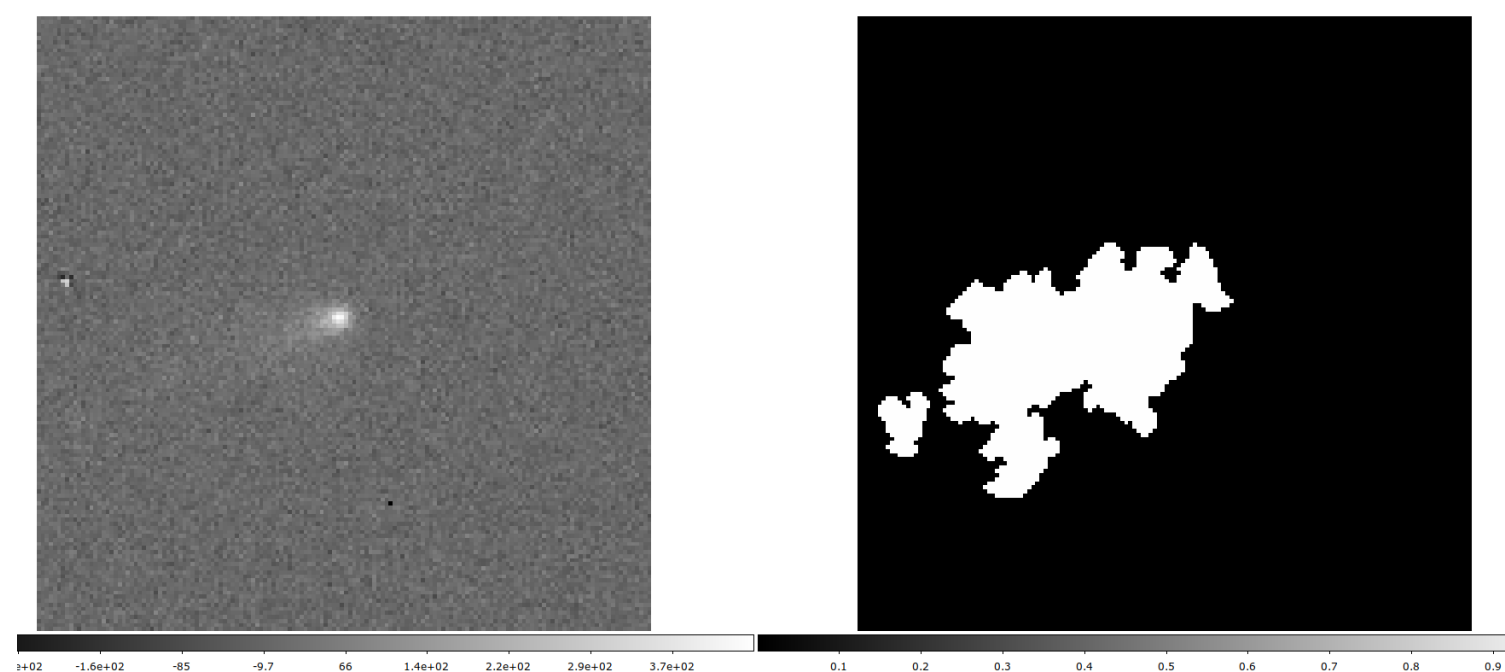


- LSST alerts - Extendedness, trailed source, predicted mag
- **LSST:UK Adler** - additional routines for phase curve photometry and cutout analysis
- Modular set of tools, to be run locally/Rubin Science Platform/alert broker (Lasair)

Wedge
Photometry



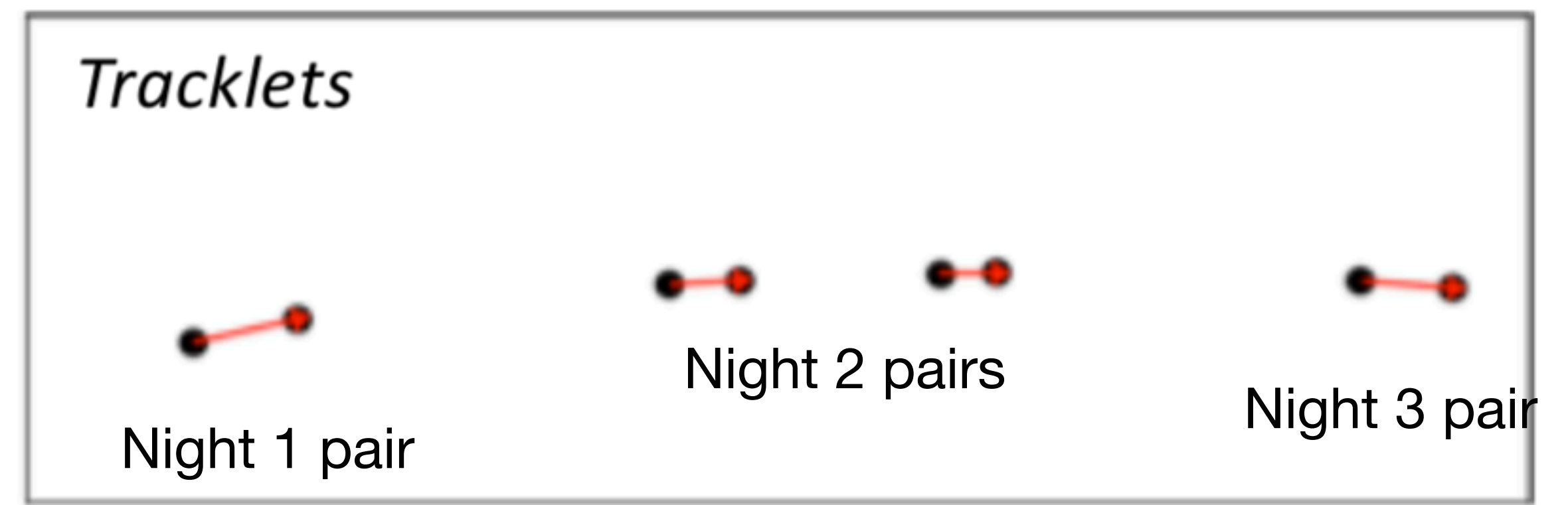
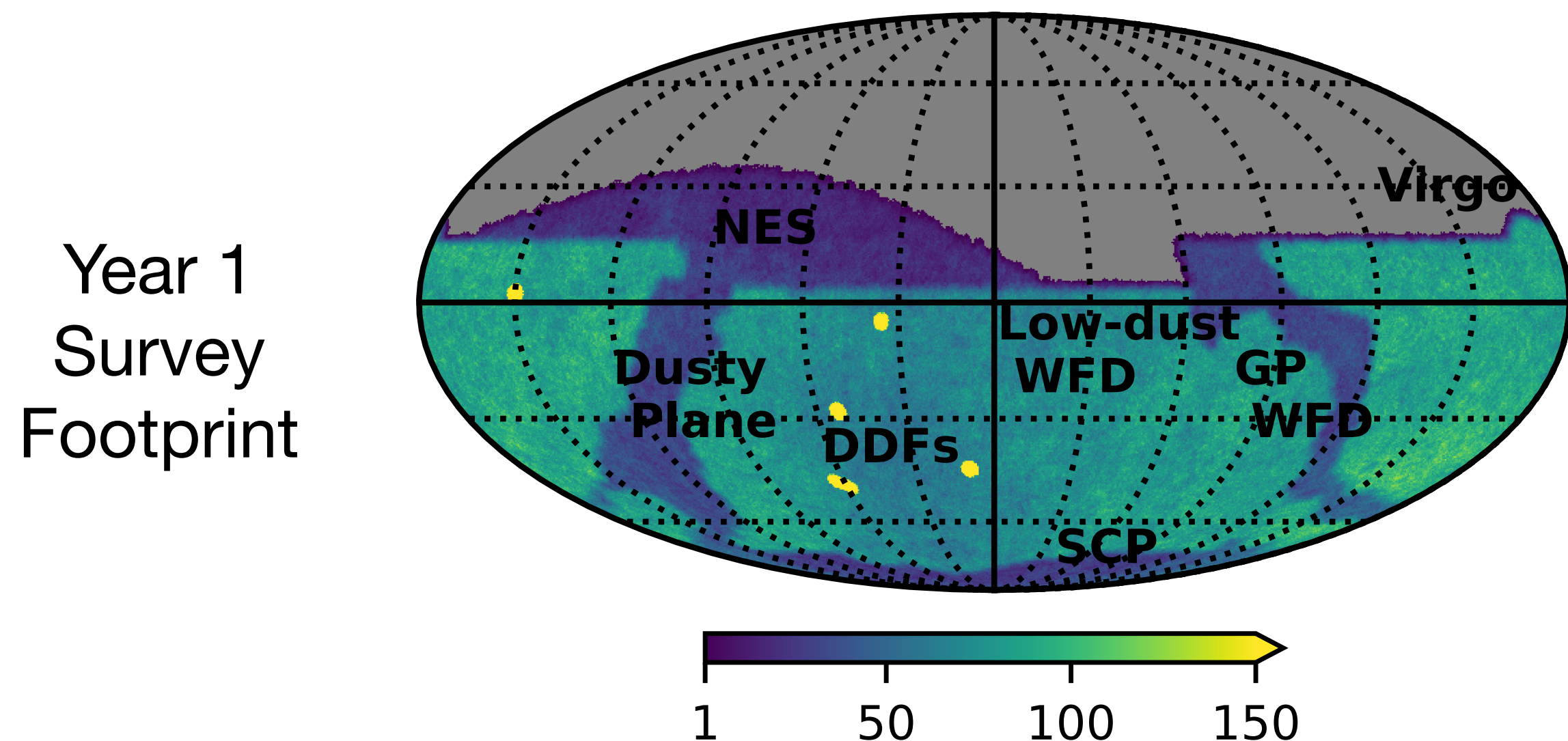
Noisechisel
(Gnuastro)



Rotational likelihood phase curve model
(Schemel & Brown 2021),
see also sHG1G2 (Carry et al. 2024)

LSST Alerts

- The Rubin Prompt Processing pipeline requires a template image of a visit sky area (in the same filter) to perform difference imaging and generate alerts
- Solar System Processing builds discovery tracklets from alerts - without alerts there can be no SSO discoveries
- Subsequent data releases will reprocess all images - alerts are not lost, they just aren't real-time



Mario Jurić

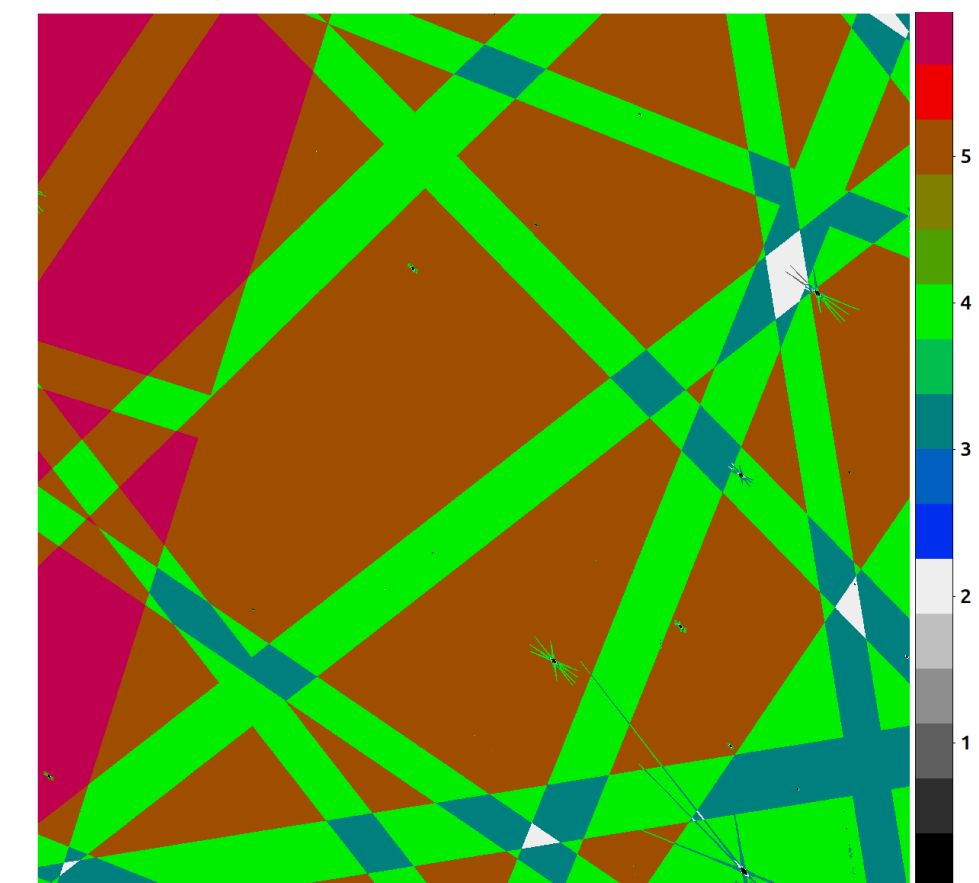
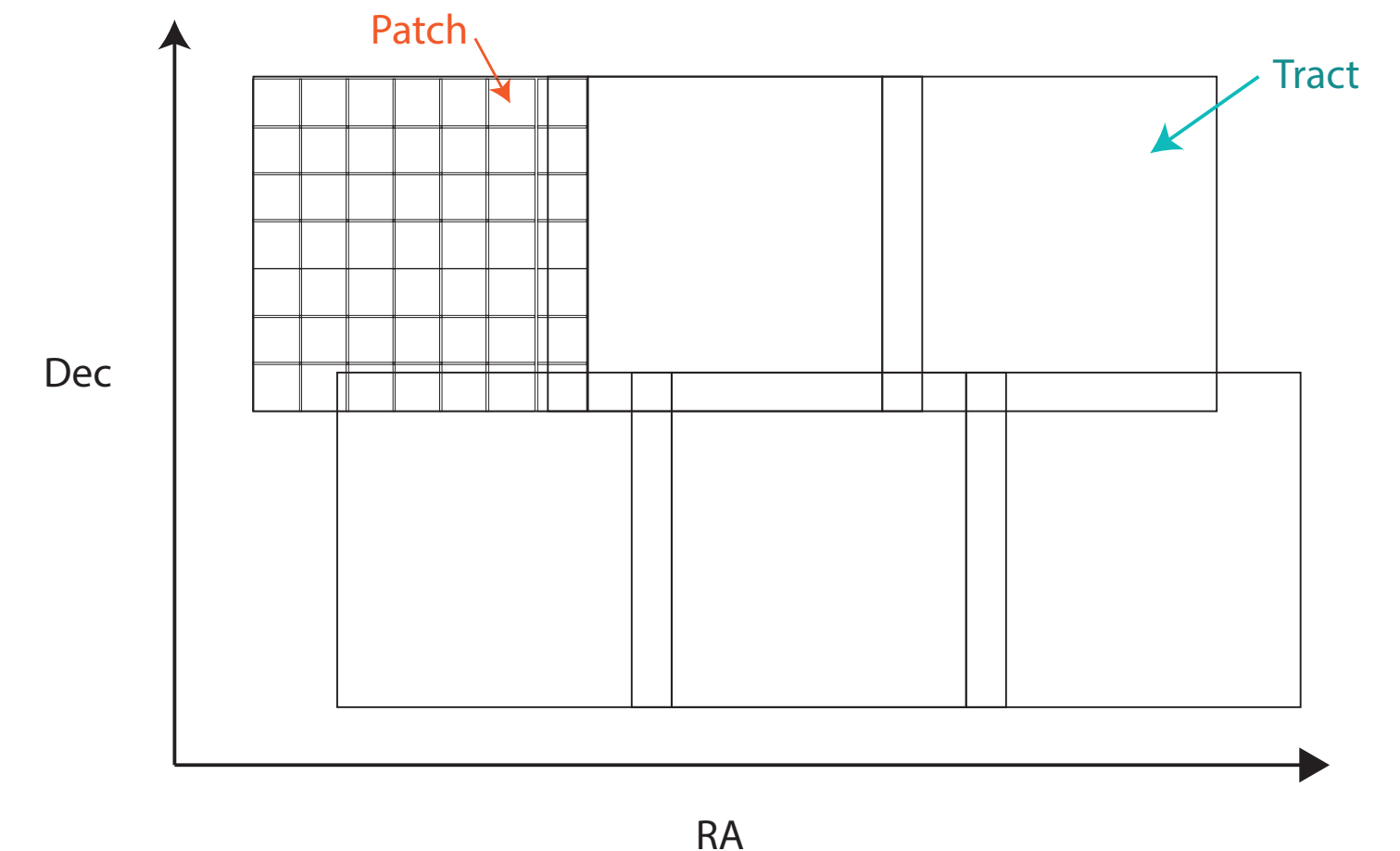
Simulating template generation

- Consider templates for one_snap_v4.0, nside = 256 healpixels (~14 arcmin scale, comparable to patch size) using Rubin Metric Analysis Framework (MAF; Jones et al. 2014)
- Pick timescale for template generation (every $\Delta t = 3, 7, 14, 28$ days)
- Consider all visits (in given filter) up to template generation night $t_n = t_0 + \Delta t * n$ (year 1 only)
- A healpixel has a template on night t_{n-1} if ≥ 4 visits (of sufficient quality) were available and will generate alerts in the period $t_{n-1} < t < t_n$
- Consider the year 1 visit database in cumulative chunks ($t \leq t_n$) and record metrics for healpixels with templates

Template Image Requirements

seeing/min(seeing) < seeing ratio = 2

(max(fiveSigmaDepth) - fiveSigmaDepth) < fiveSigmaDepth range = 0.5

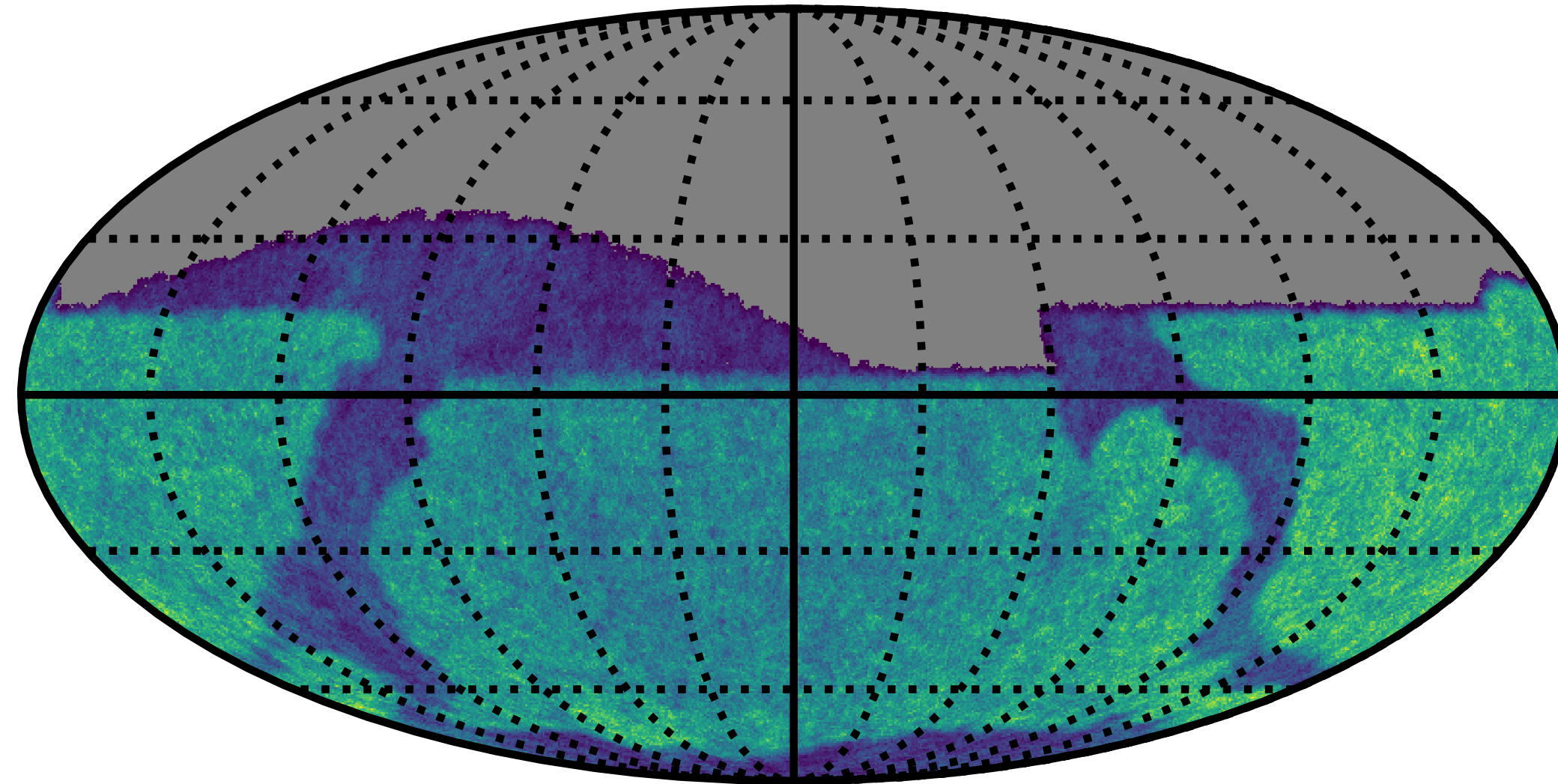


Coverage of a random patch
(John Parejko)

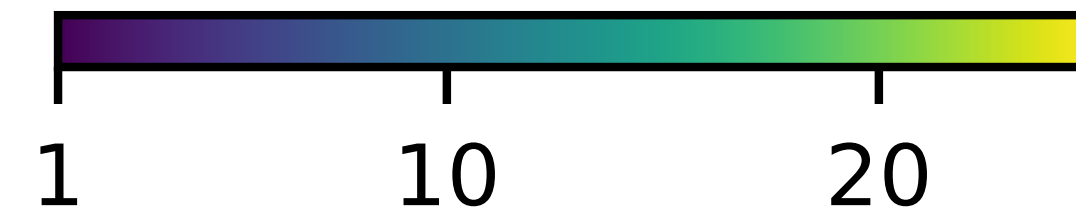
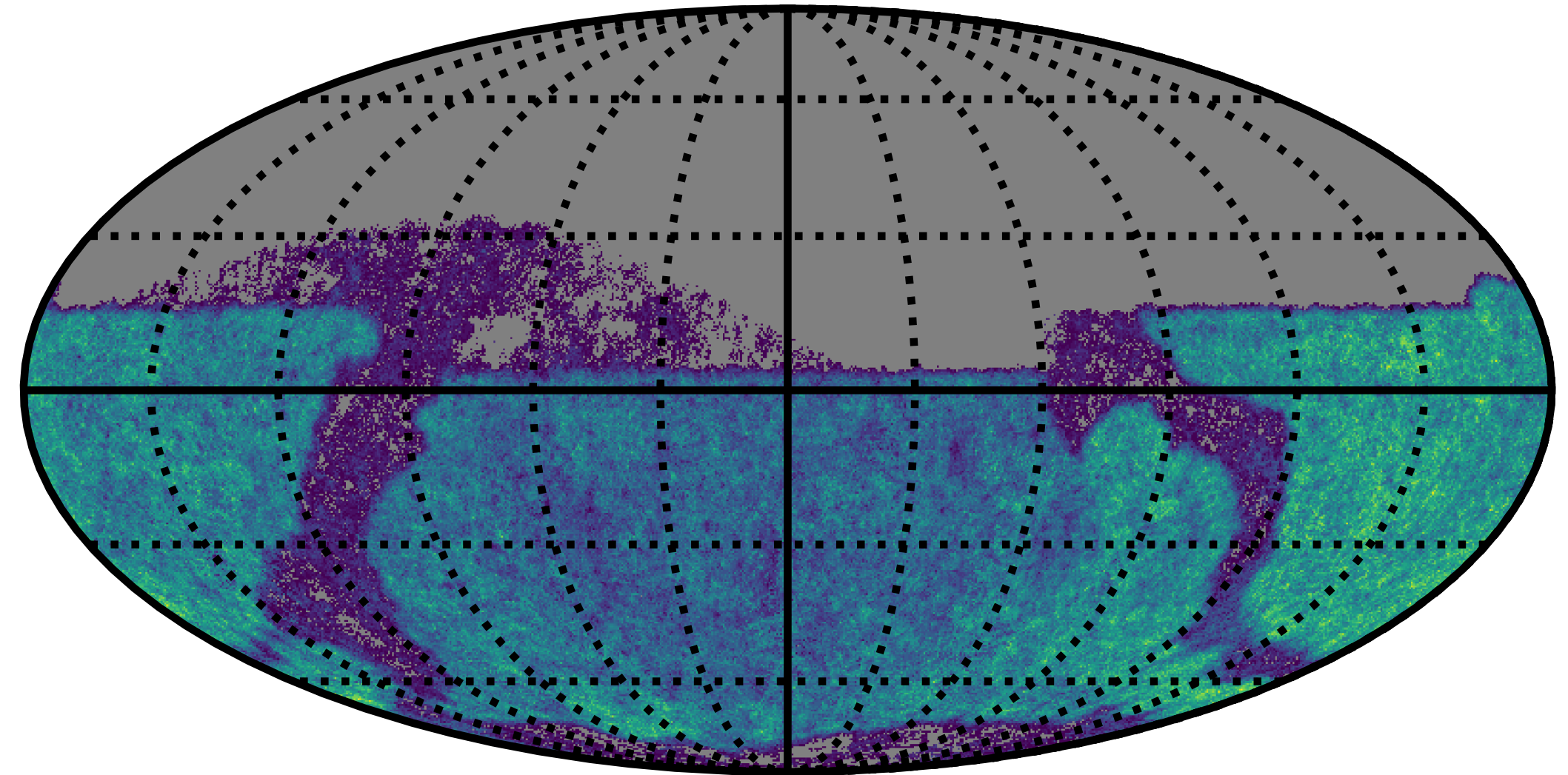
Thank you to Lynne Jones and Peter Yoachim for the notebooks to start this work!

Template Coverage

N visits, filter = r



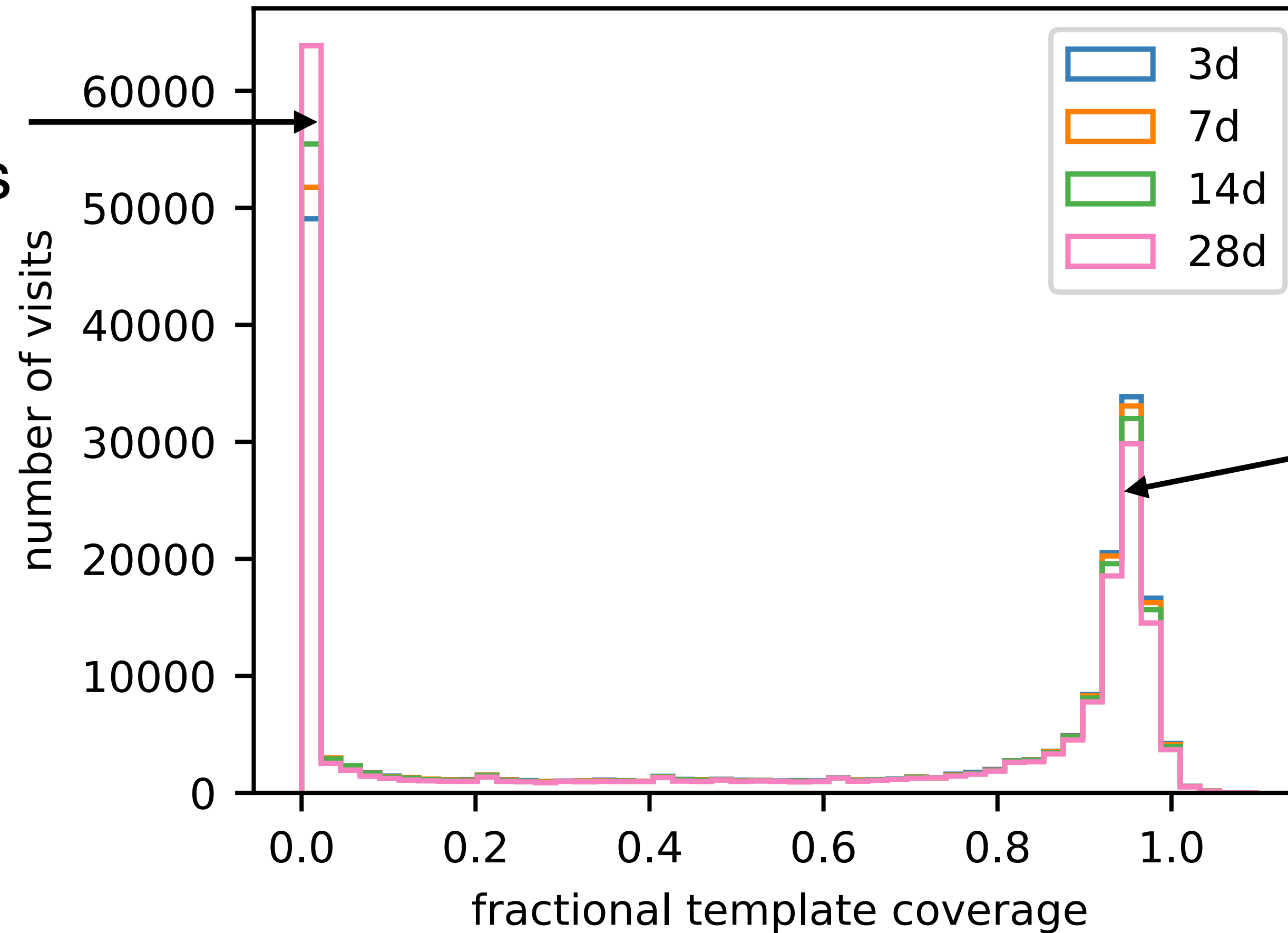
N visits, $\Delta t = 7\text{d}$, filter = r



Template generation reduces number of visits producing alerts, with patchier coverage (Robinson et al. 2025)

Template Coverage

Visits likely to be
used for templates

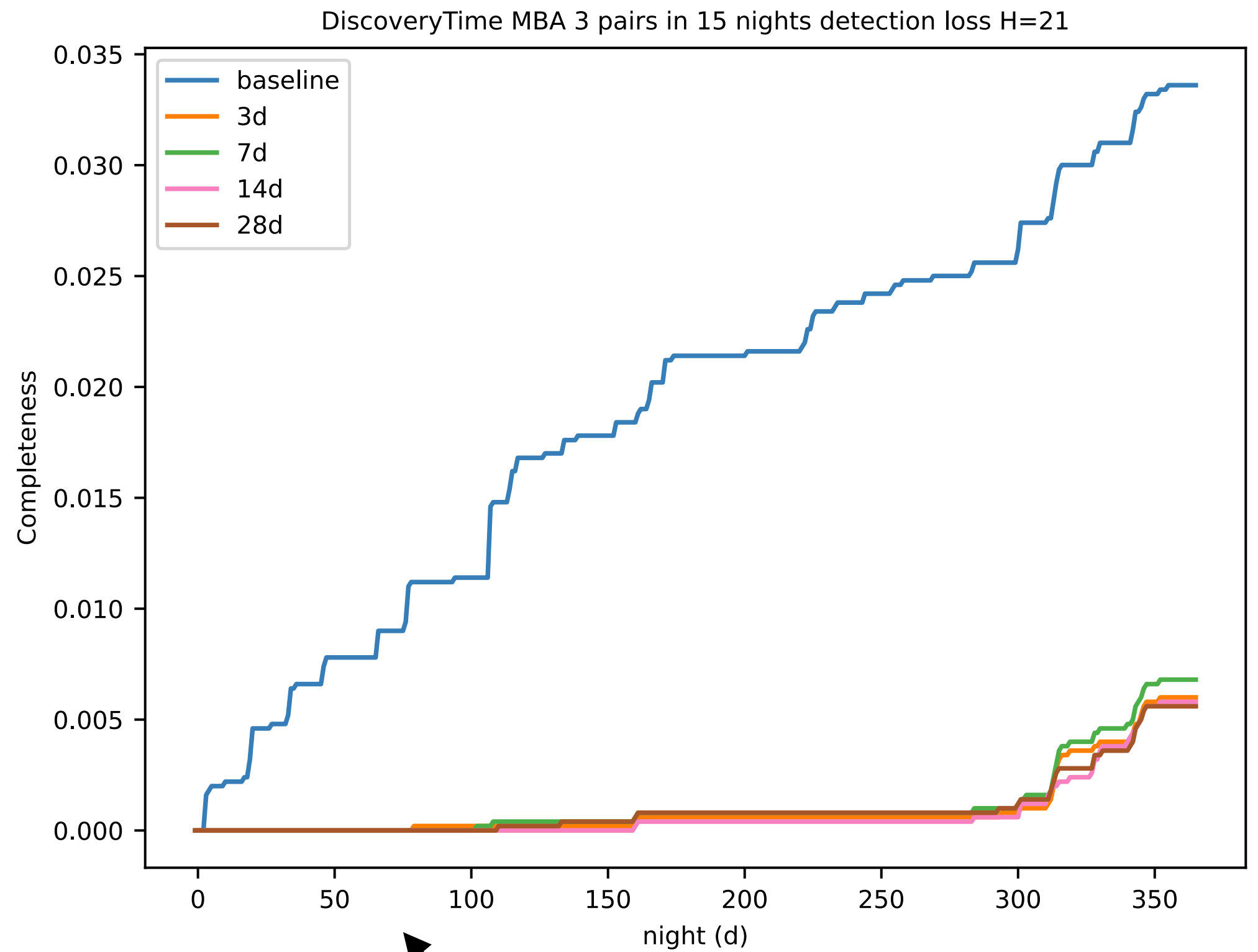
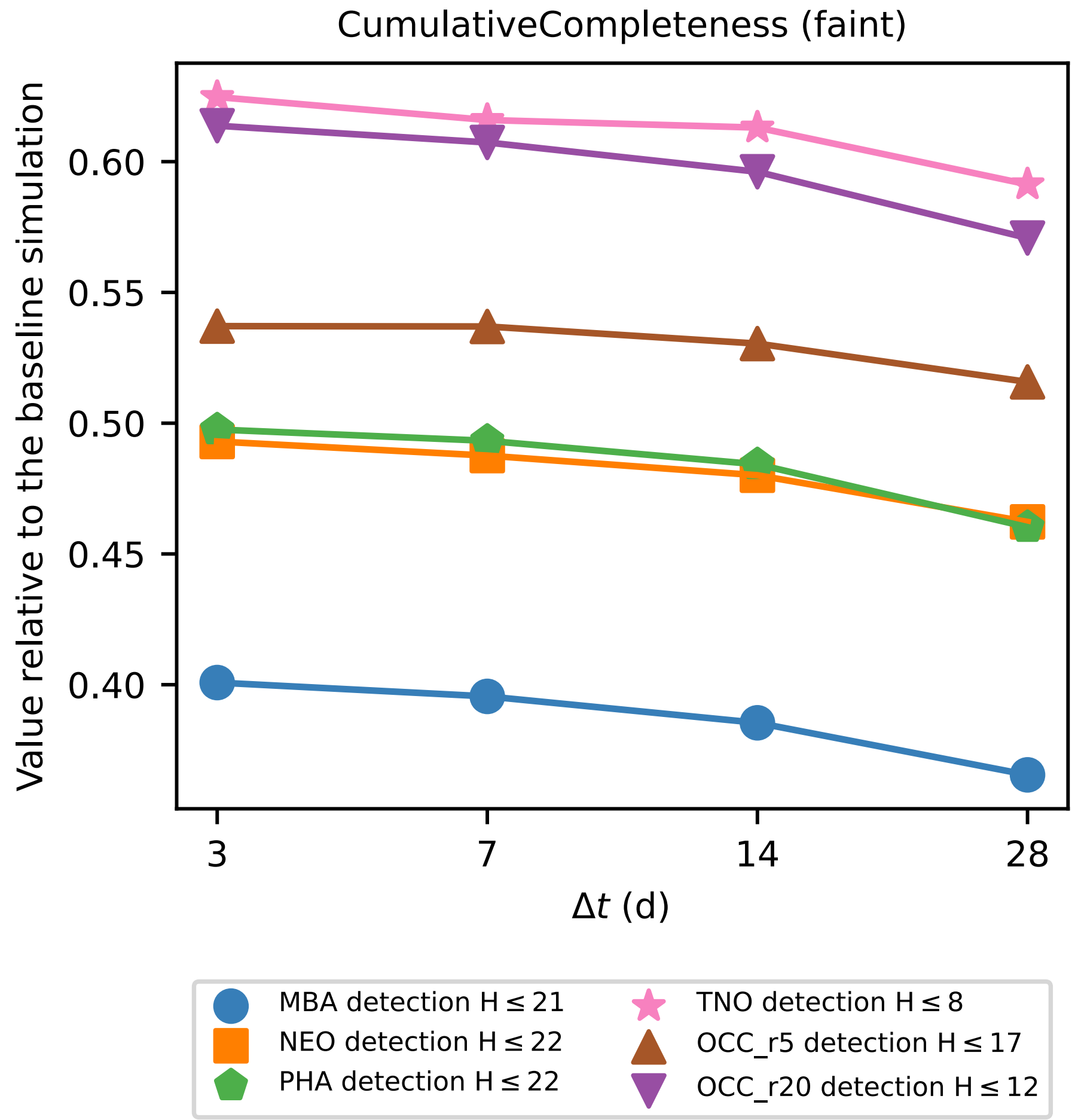


Visits likely to
produce alerts

Visits (~183 healpixels) generally have 0% or >90%
template coverage

SSO Discoveries

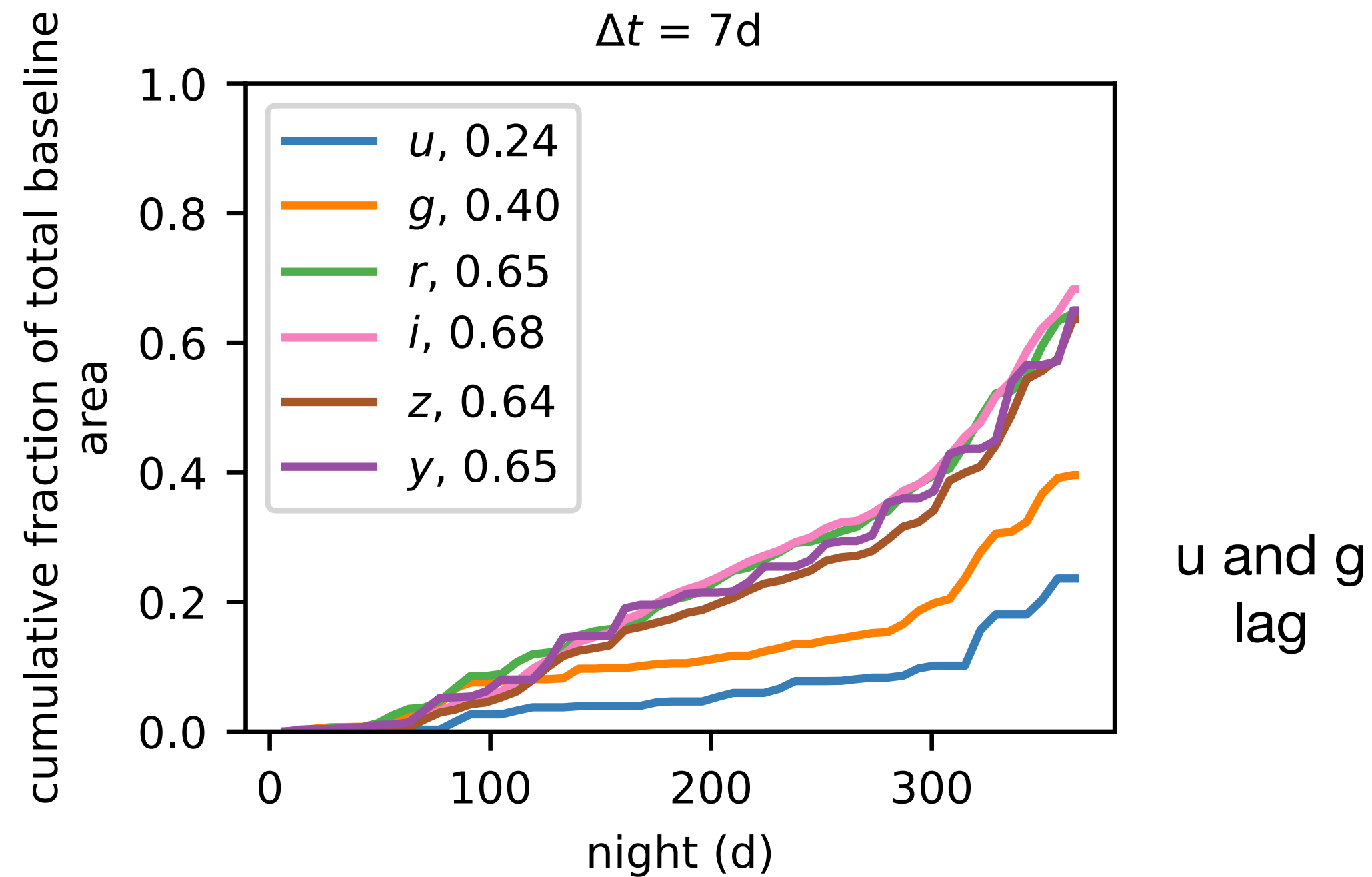
Drop in discoveries depends on population:
Main Belt Asteroids (MBAs)
Near-Earth Objects (NEOs)
Potentially Hazardous Asteroids (PHAs)
Trans-Neptunian Objects (TNOs)
Oort Cloud Comets (OCC)



~50 day time delay

*results calculated for visits with
>90% template coverage only

Problem Filters/Areas

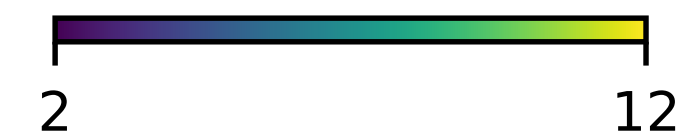
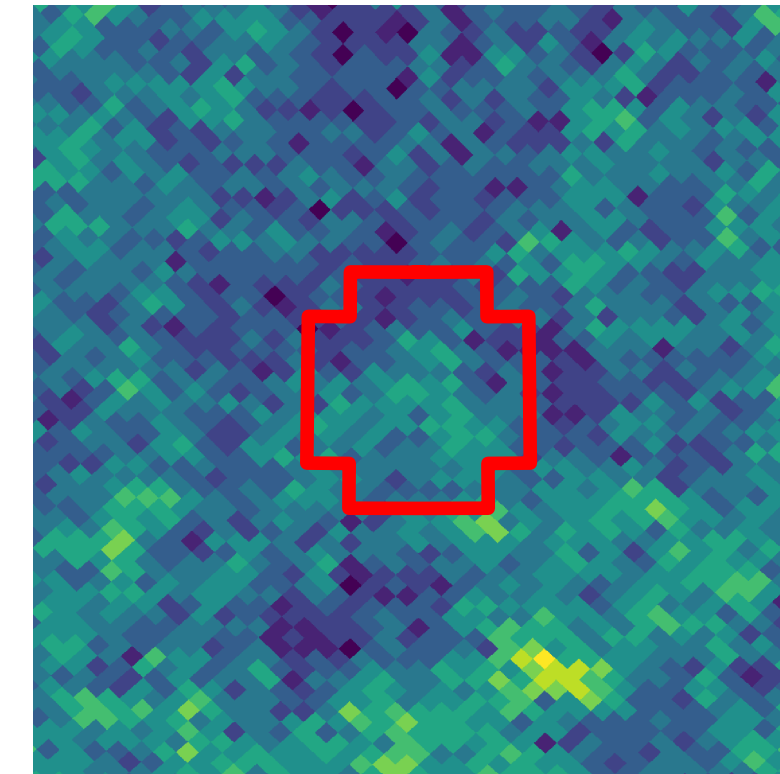


Filters or sky areas with low coverage in one_snap v4.0 are impacted more strongly by template generation. Important for:

- colour measurements
- spatially dependent populations

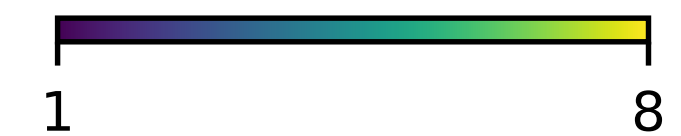
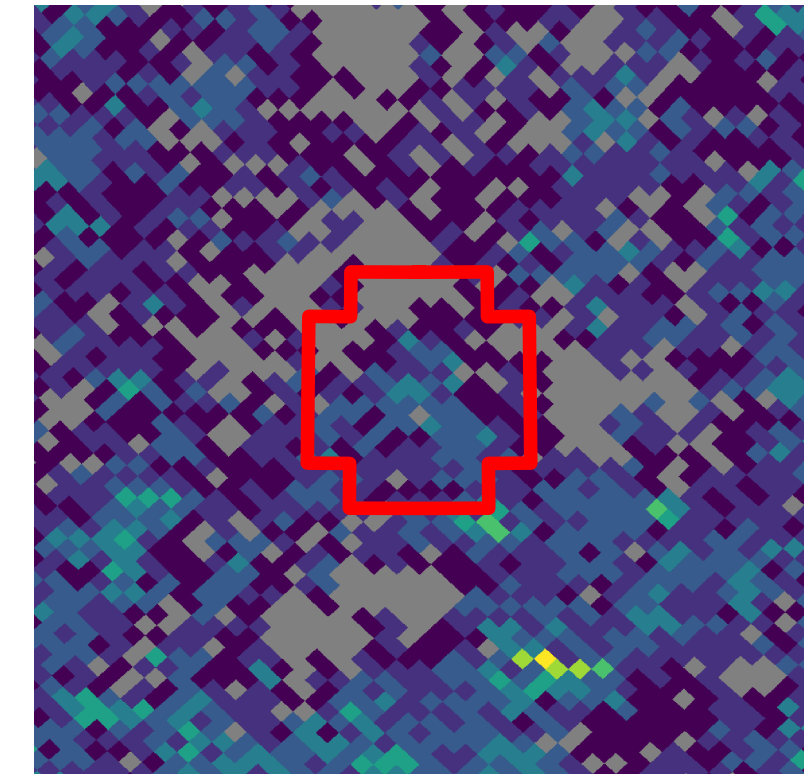
1

N visits, baseline



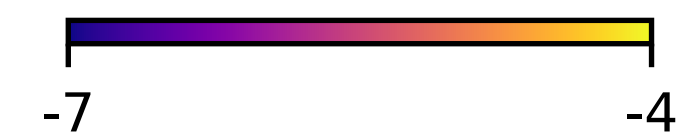
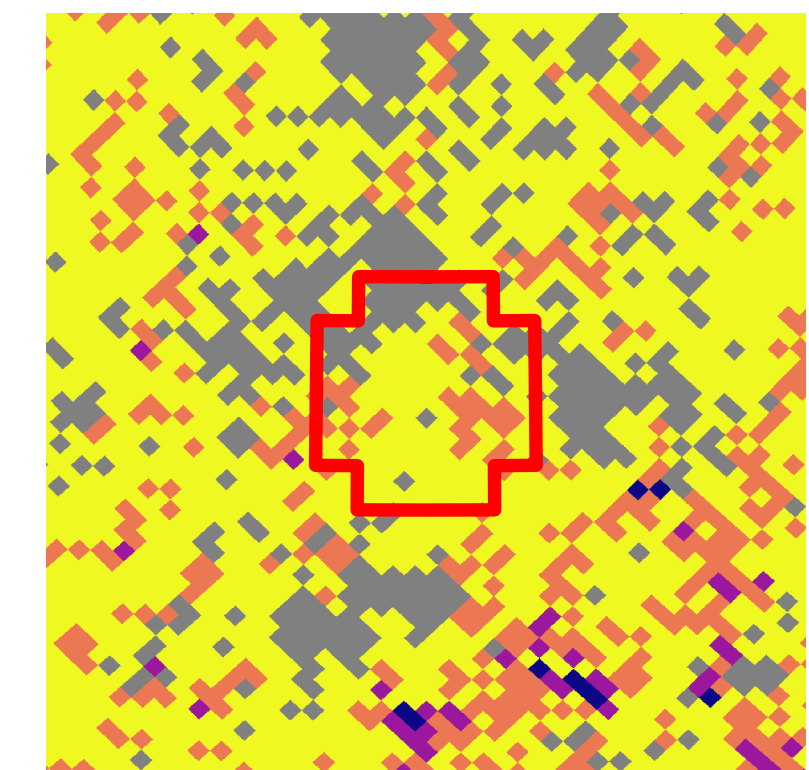
2

N visits, $\Delta t = 7d$



3

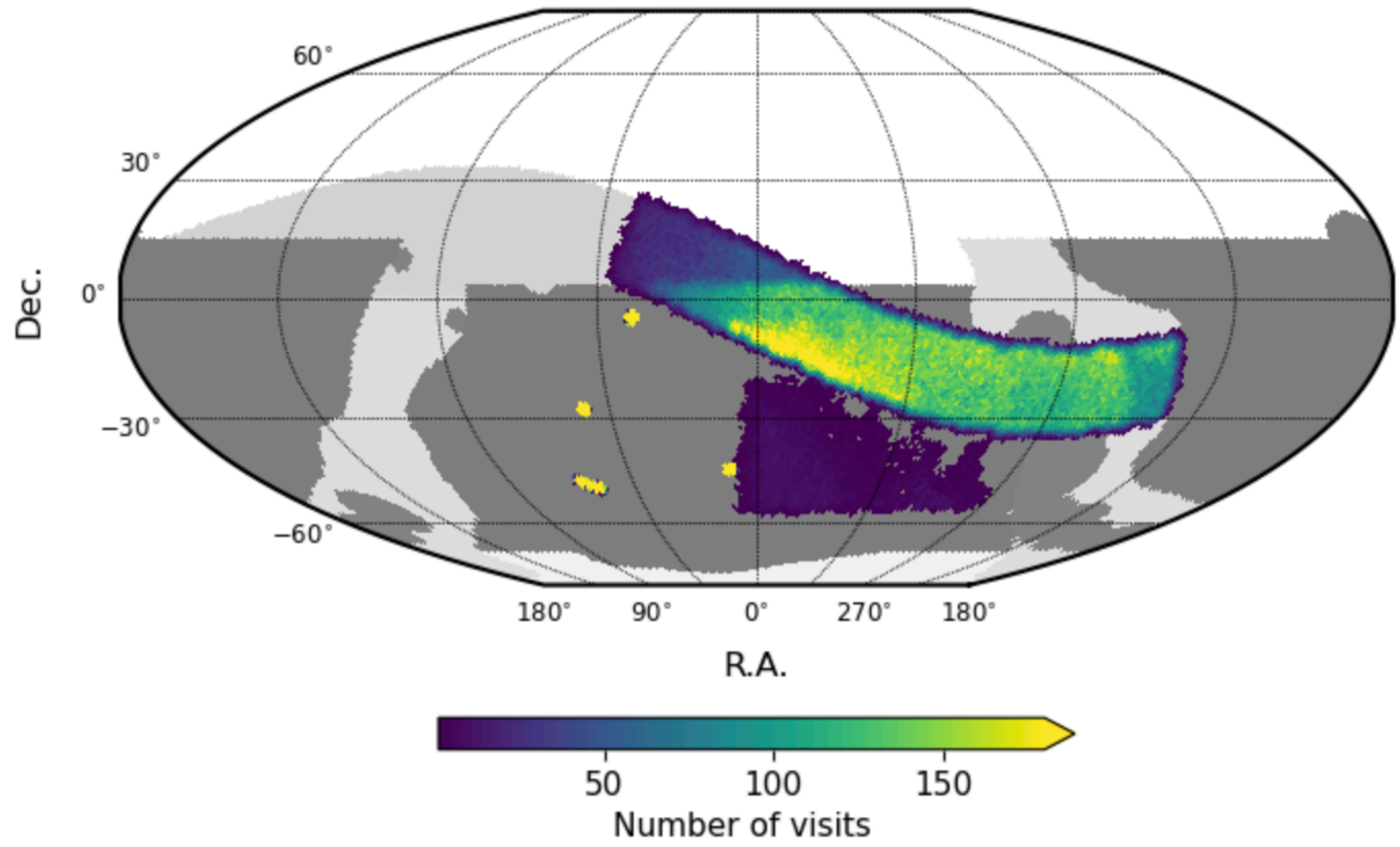
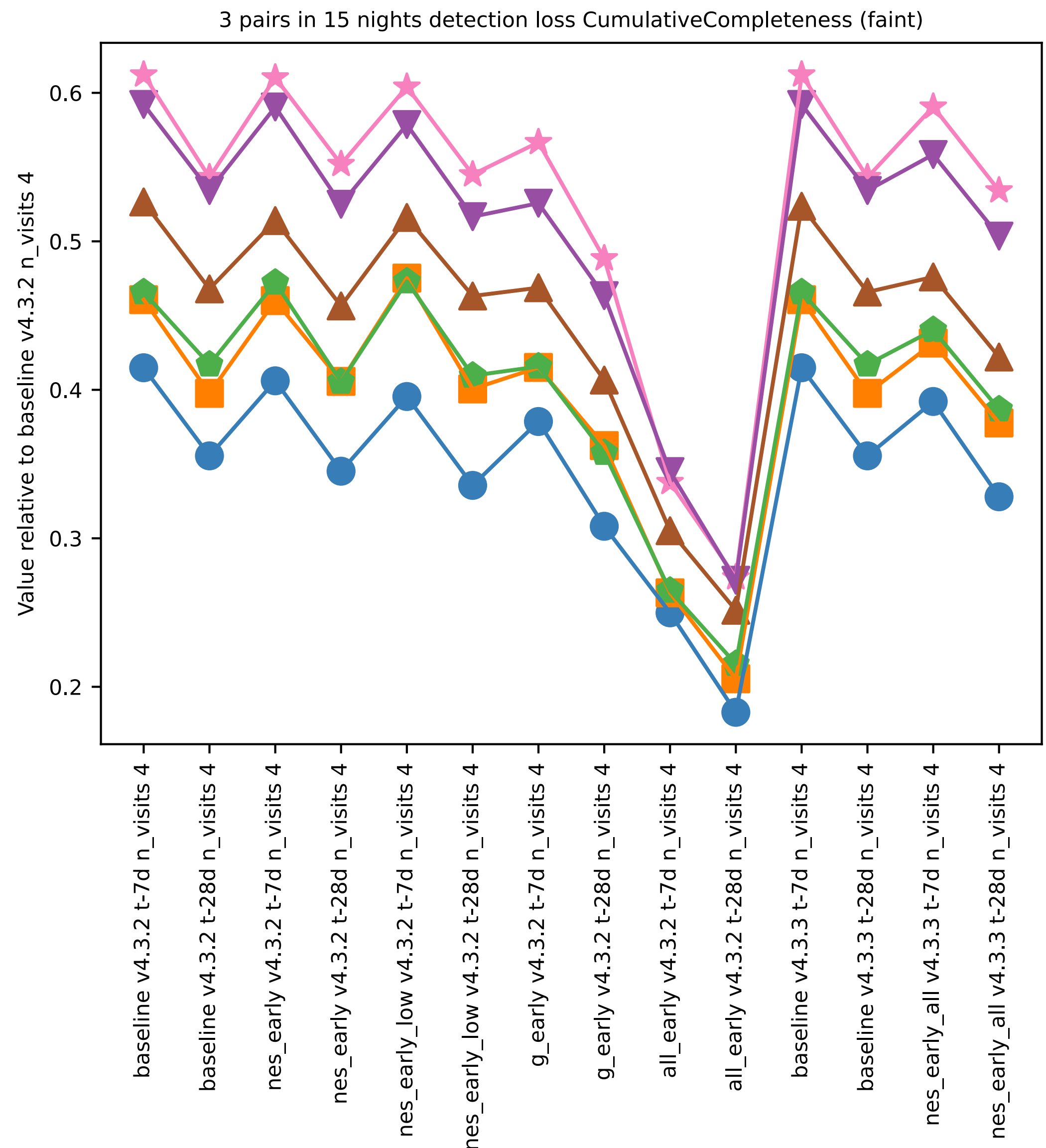
ΔN visits, $\Delta t = 7d$



Sample North Ecliptic Spur (NES) region in the r filter:

1. one_snap v4.0
2. template generation ($\Delta t = 7d$)
3. Difference between 2 and 1

New Baselines and Science Validation



new baselines
on community
forum:



SV info:
SITCOMTN-005

Conclusions

- **Template generation is relevant to any science case that requires real-time alerts in year 1**
- Template generation reduces the sky area capable of producing alerts - SSO discovery & characterisation in year 1 is greatly impacted (up to 50% loss)
- Generating templates as soon as sufficient images are available maximises alert production
- Possible tweaks to year 1 strategy? Assign more time to low coverage areas/filters? We have assumed no commissioning data for templates
- **The new year 1 cadence simulations need to be assessed for impact on LSST science goals**
- More info at [#lsstc.slack.com](https://lsstc.slack.com) #incremental-template-generation and Focus Issue paper



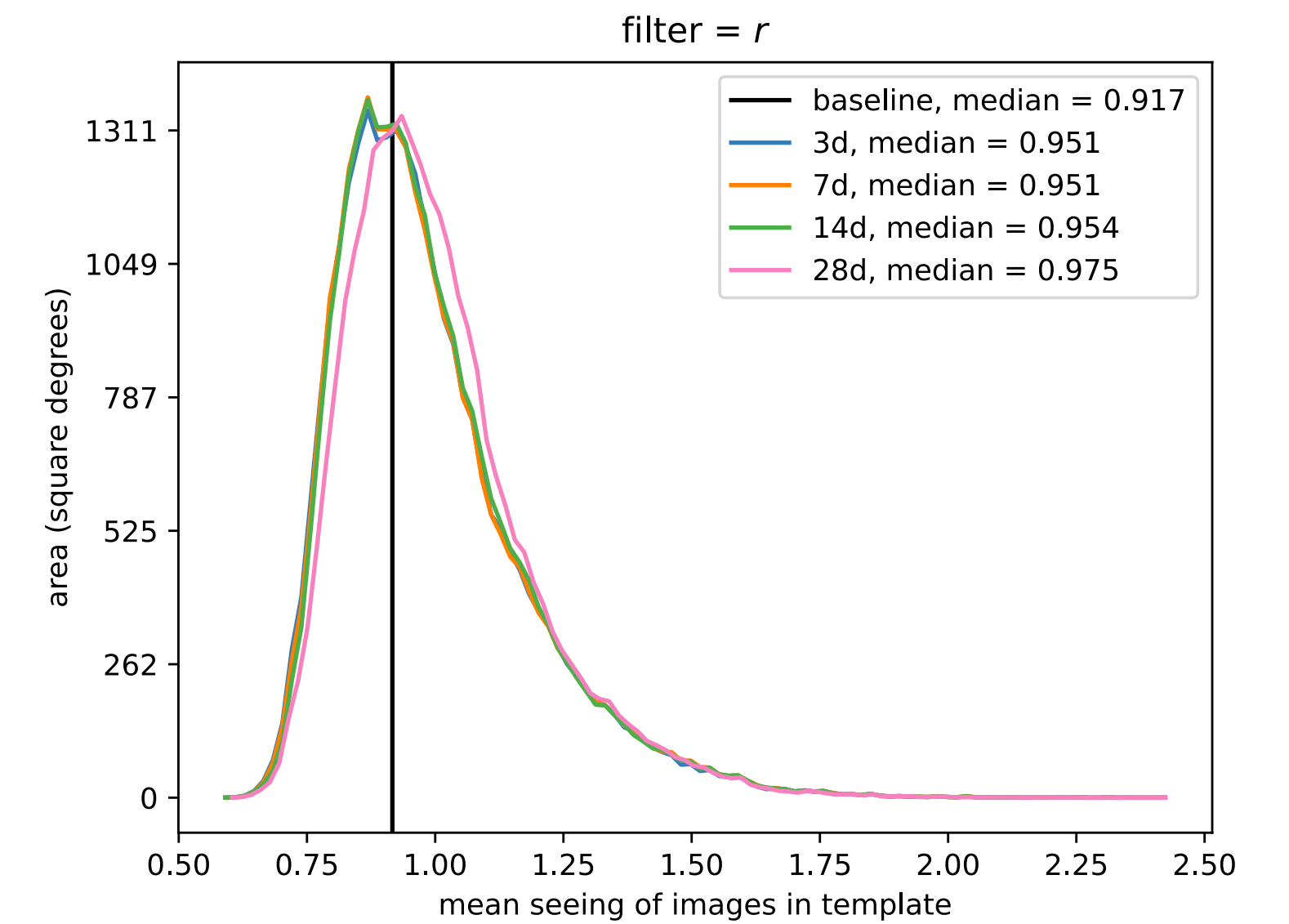
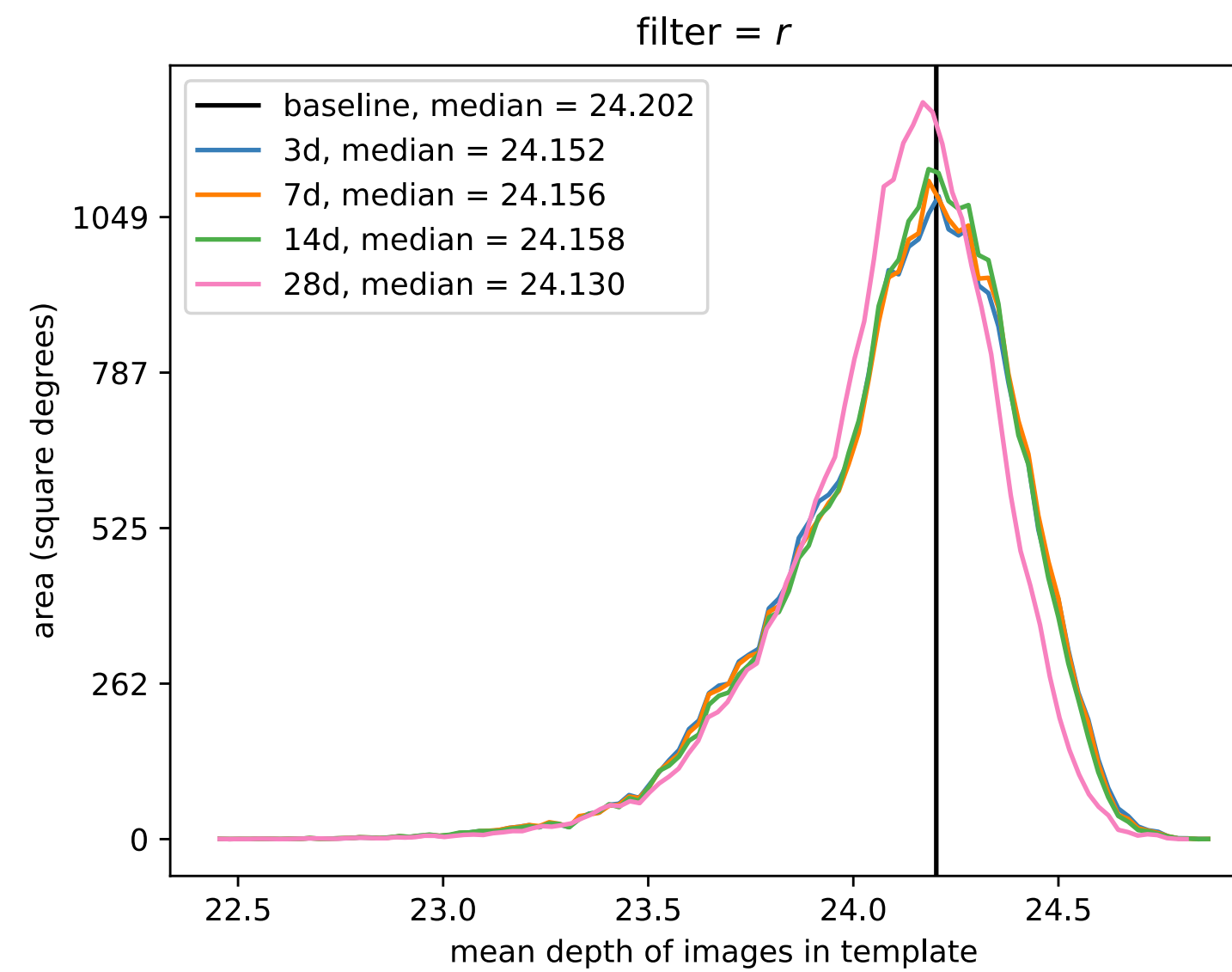
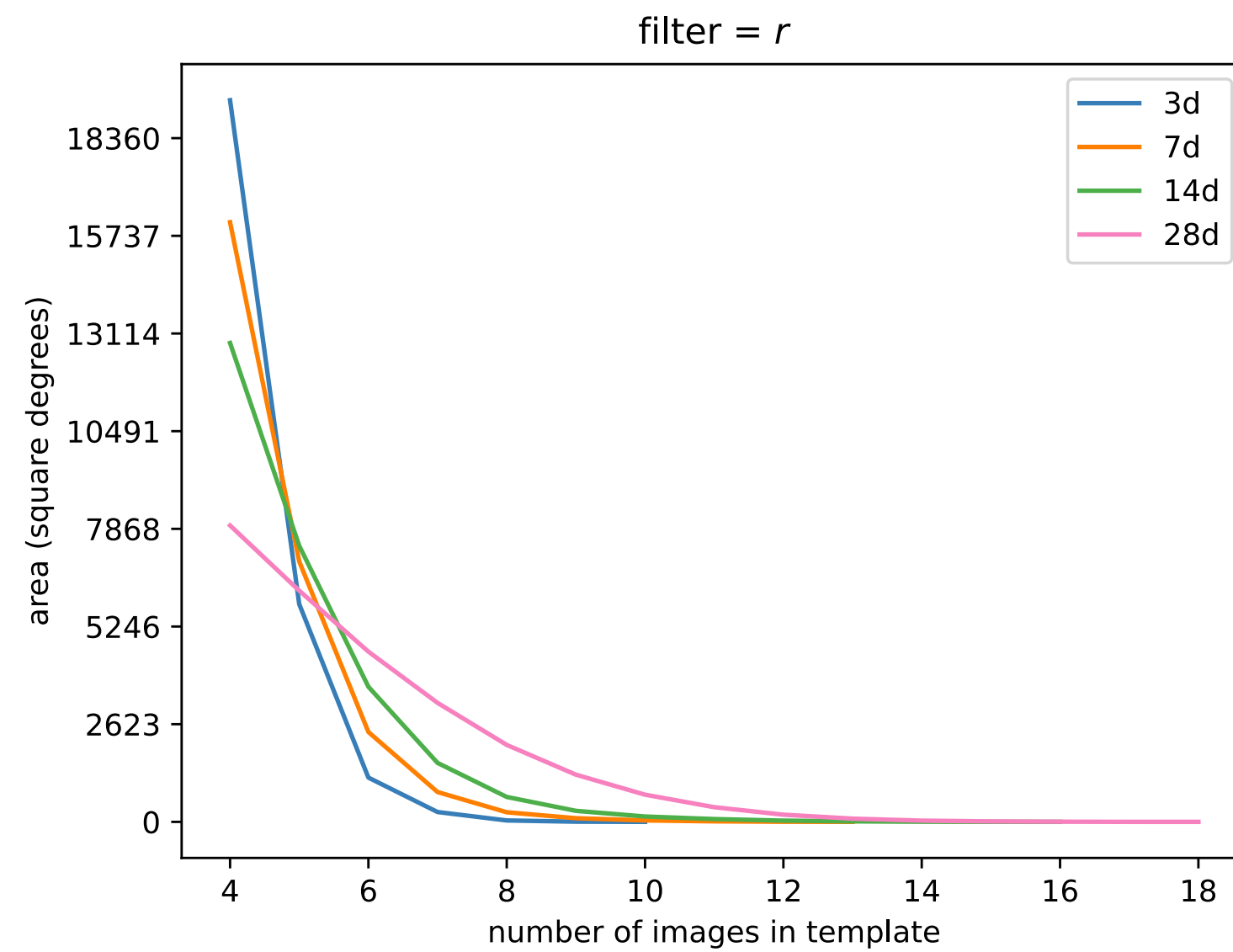
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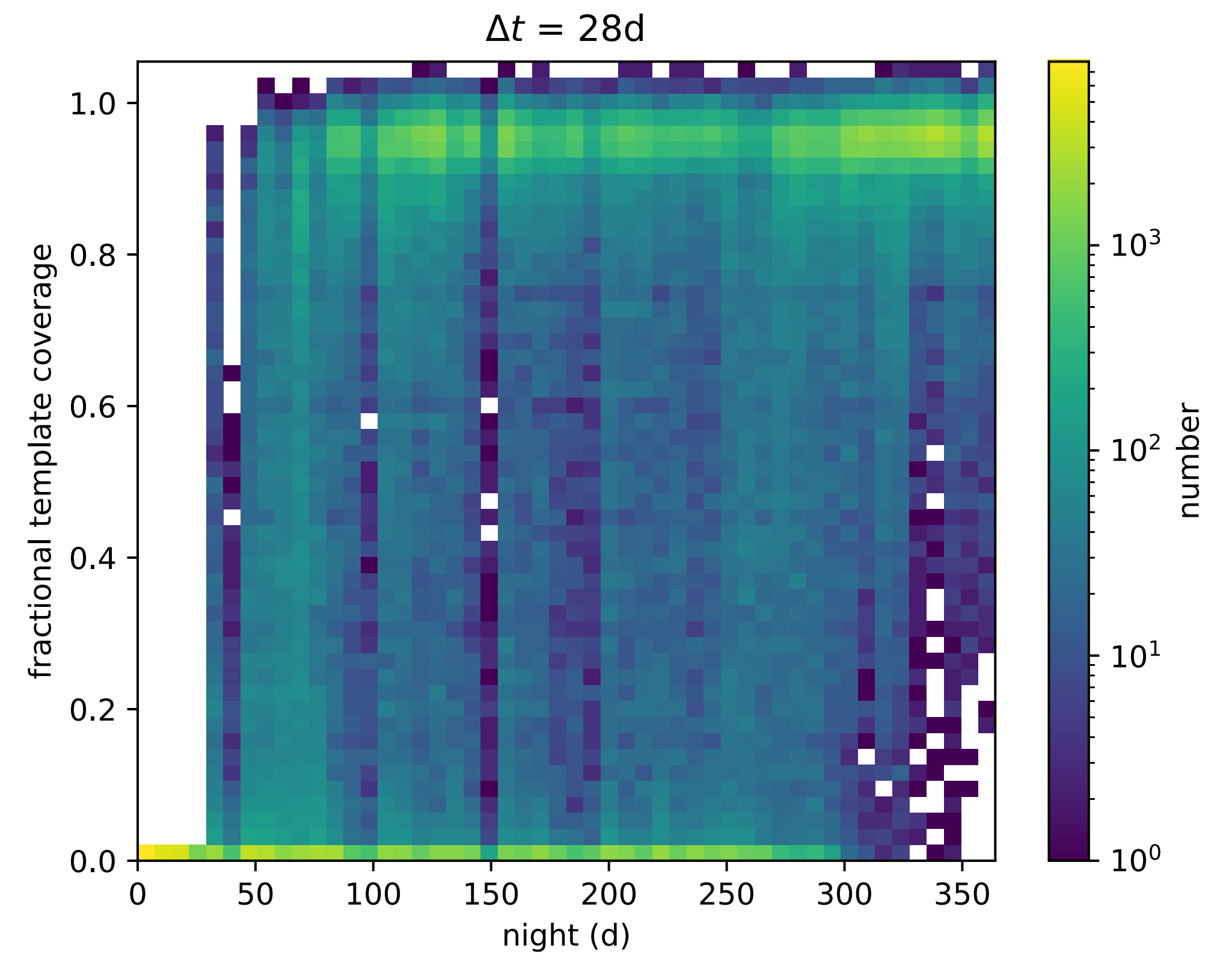
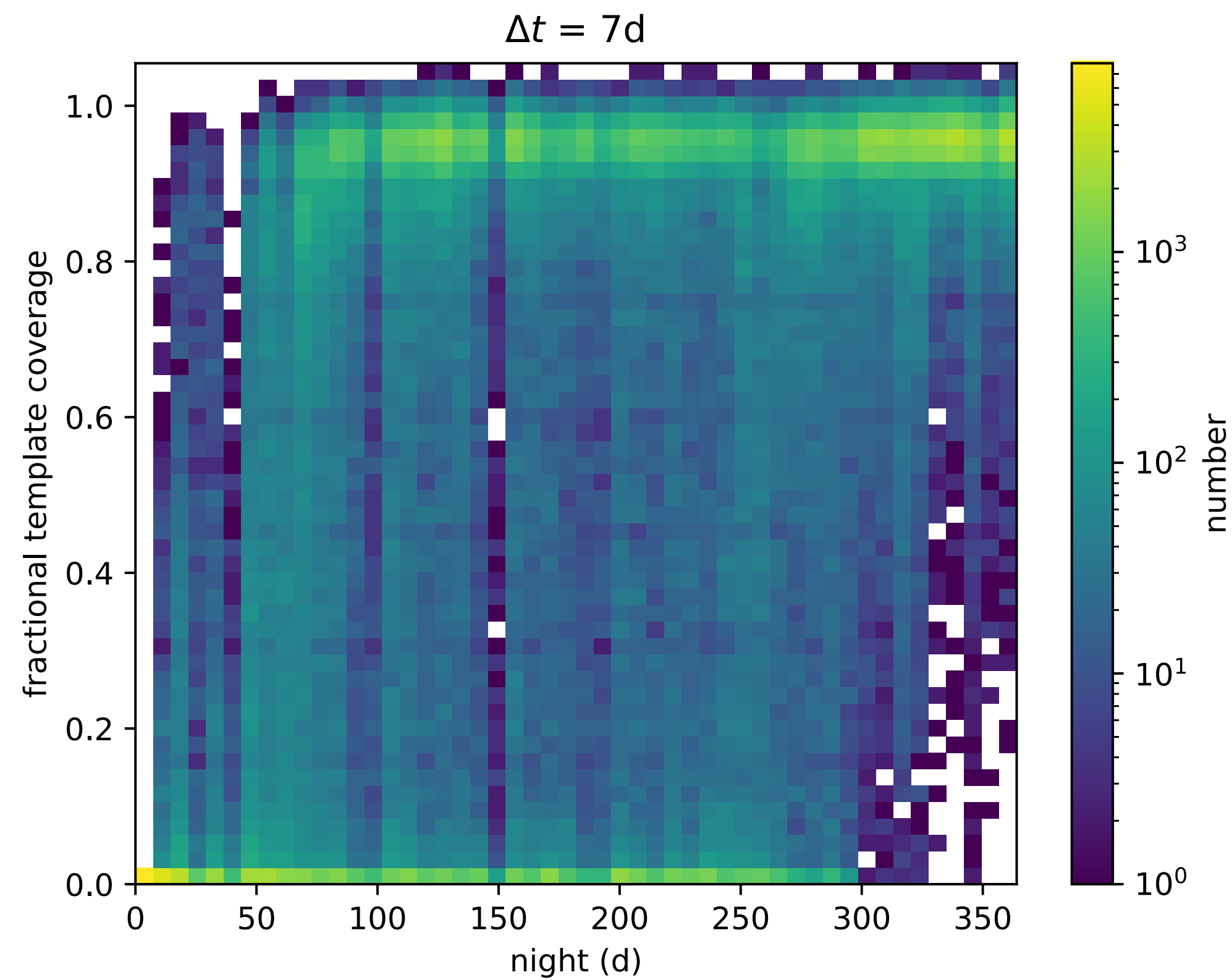
Extra Slides

(baseline v3.4 but similar enough to one_snap v4.0)

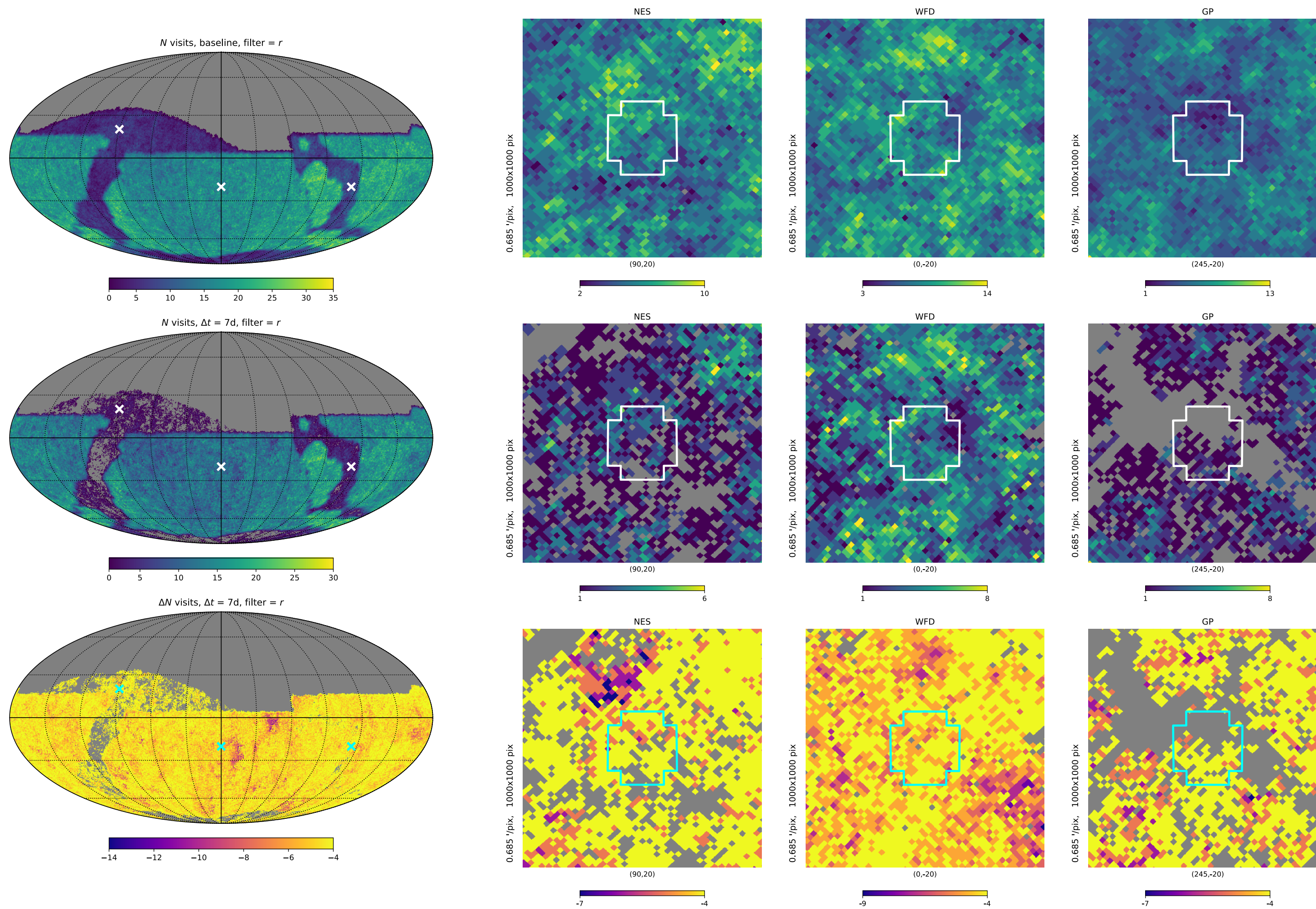
Template image quality



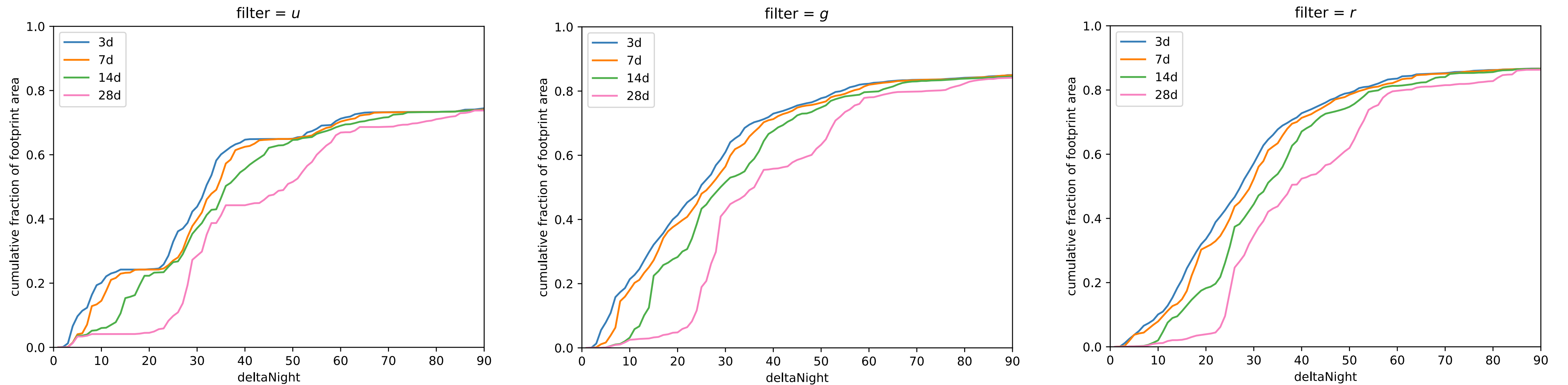
Visit template coverage



Low coverage areas



deltaNight



Number of nights between first visit to healpixel and its template generation

SSO Colours

