

Fewer Companions in the Crowd

The Low Close-Binary Fraction in Globular Clusters from *Gaia* RVS

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In collaboration with Vasily Belokurov

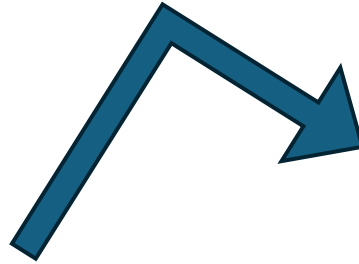
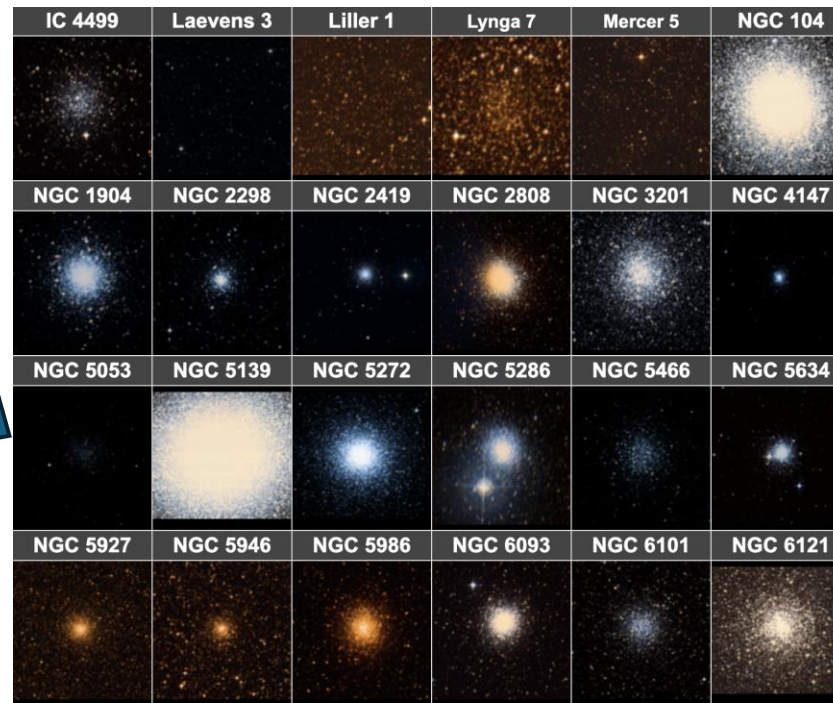
Motivation



Motivation

GC dynamics

Binary stars

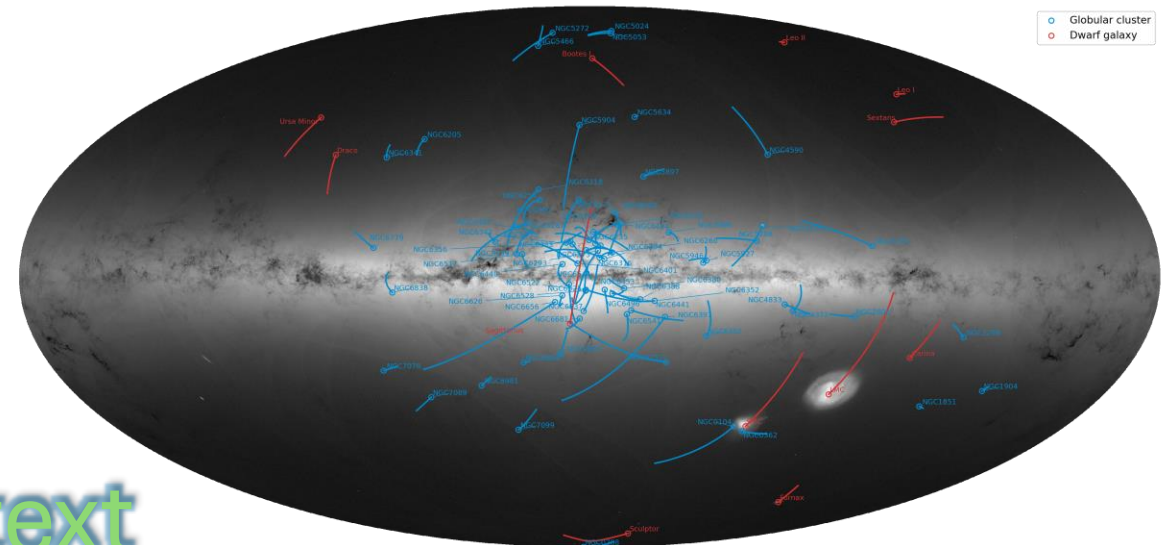
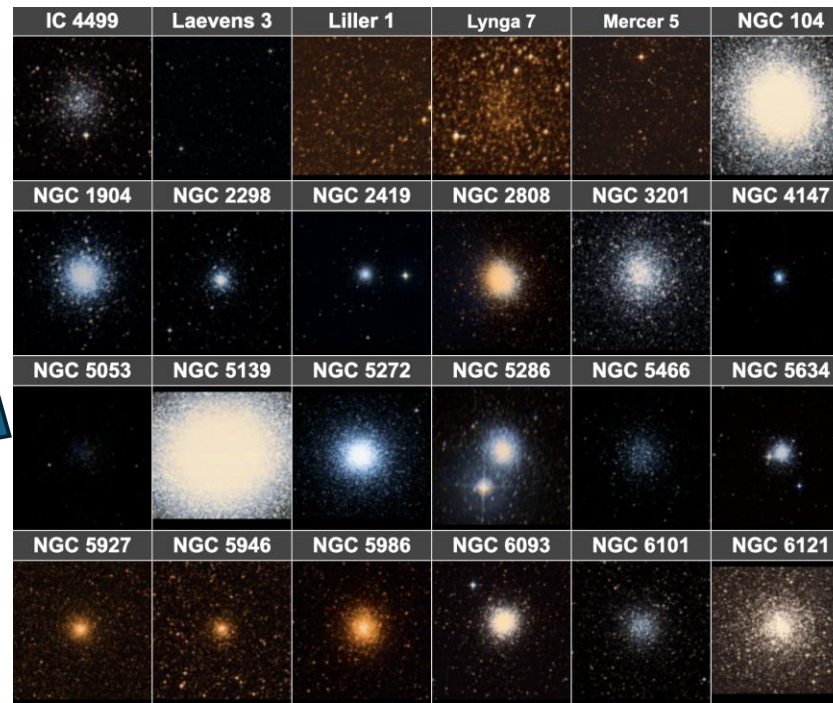


Motivation

GC dynamics

Binary stars

Galactic Context



Binaries in GCs: What do we know?

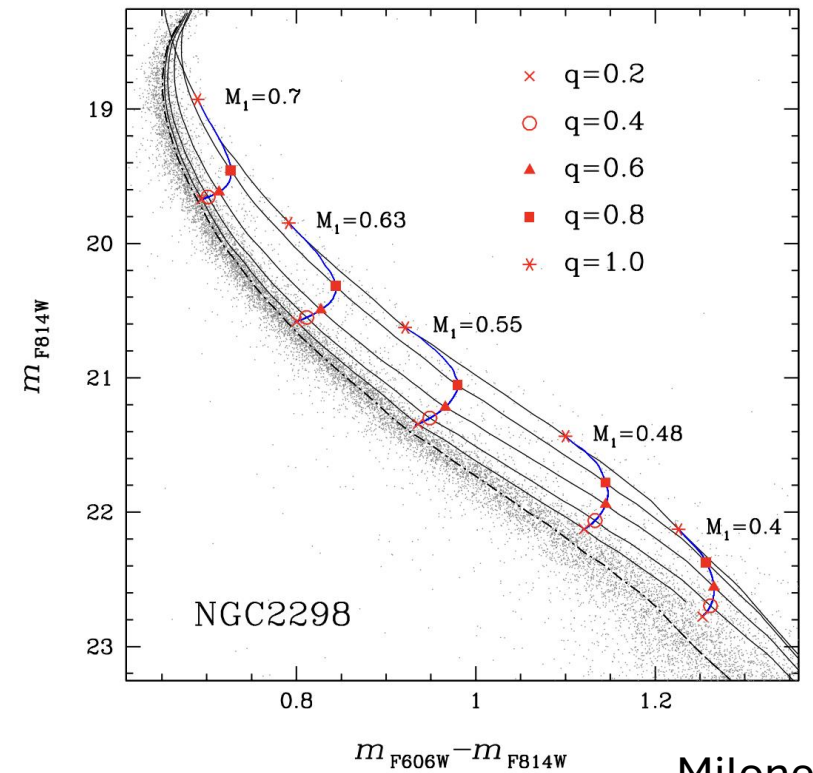
- **Spectroscopy**

- **Photometry**

Binaries in GCs: What do we know?

- Spectroscopy

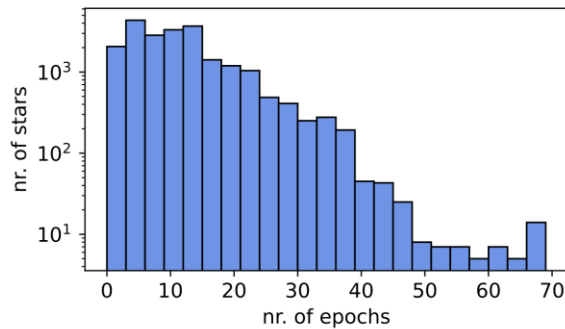
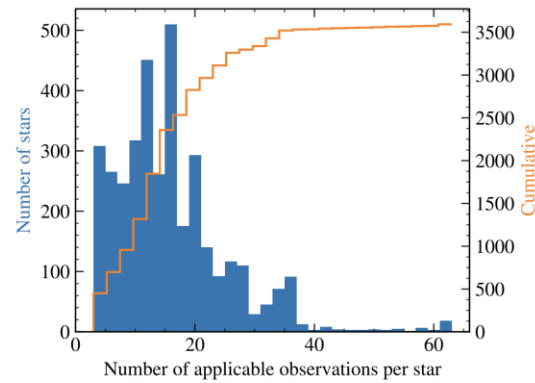
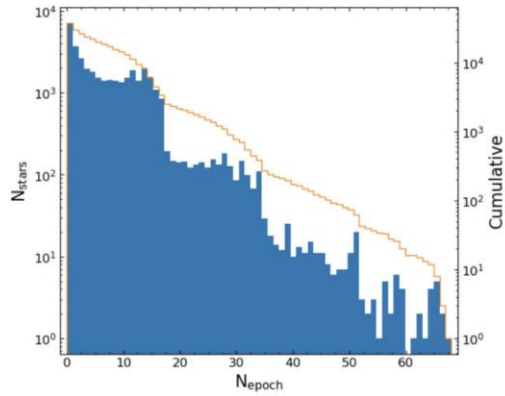
- Photometry



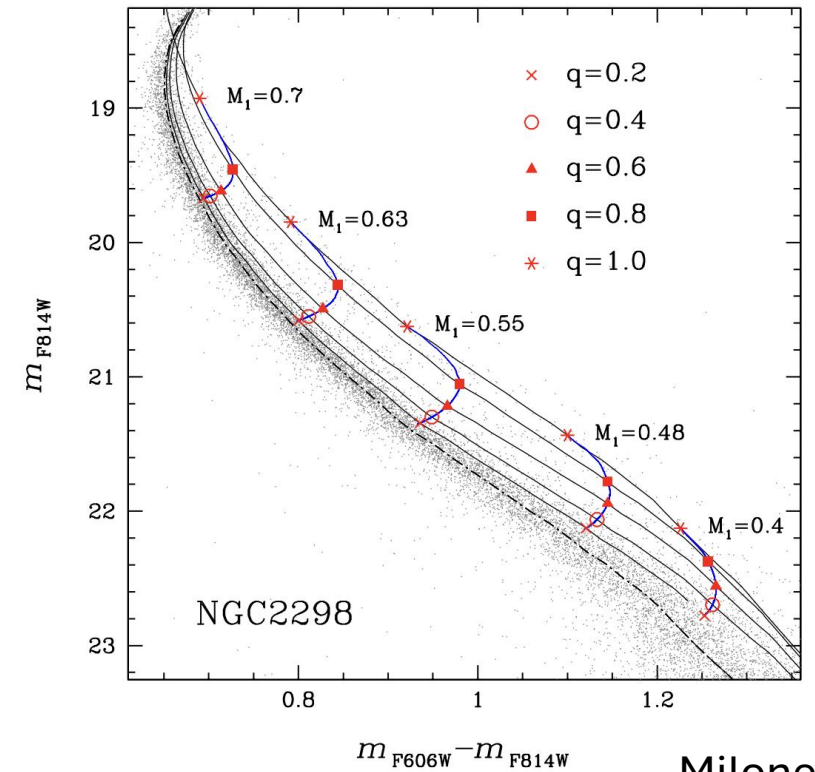
Milone+12

Binaries in GCs: What do we know?

- Spectroscopy



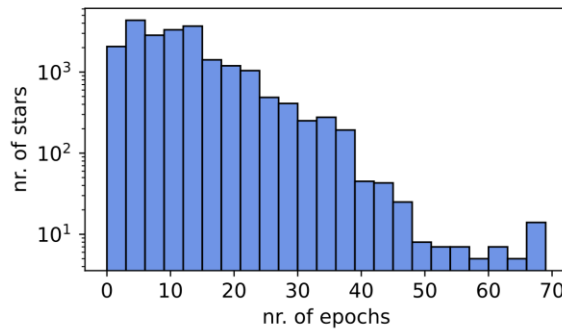
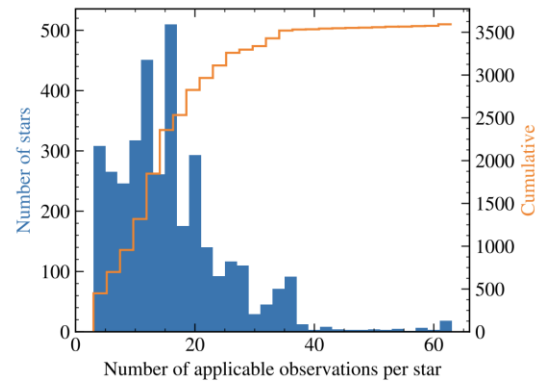
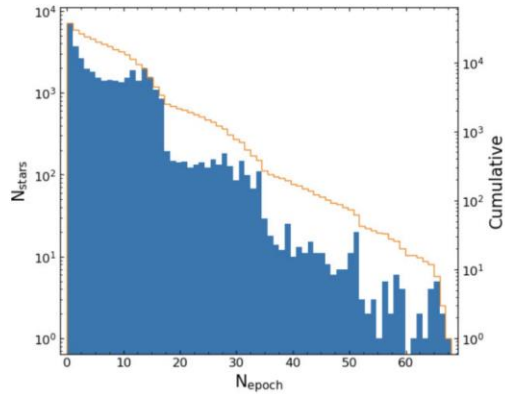
- Photometry



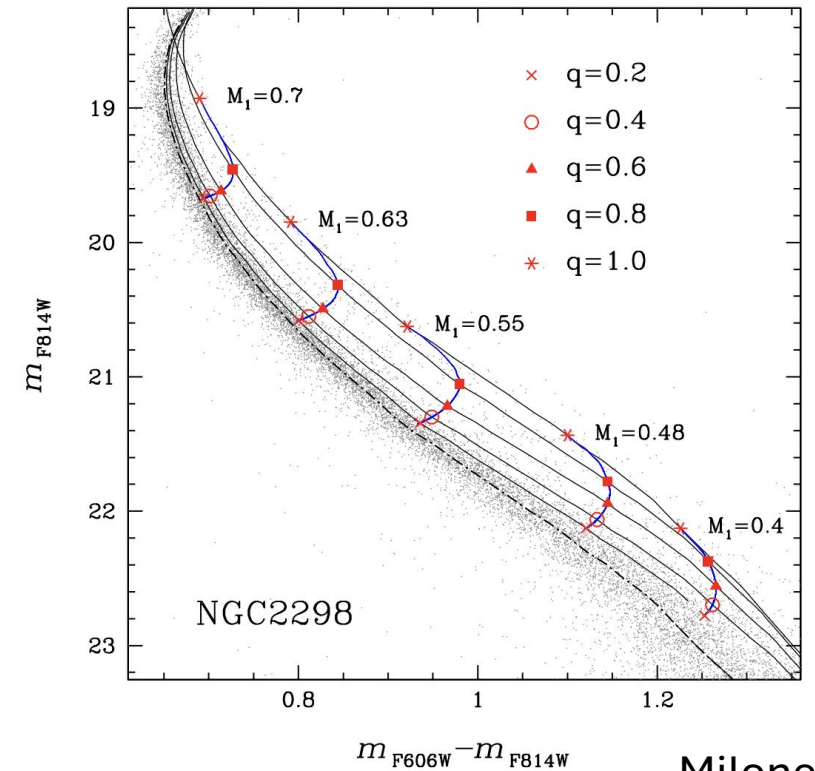
Milone+12

Binaries in GCs: What do we know?

- Spectroscopy



- Photometry

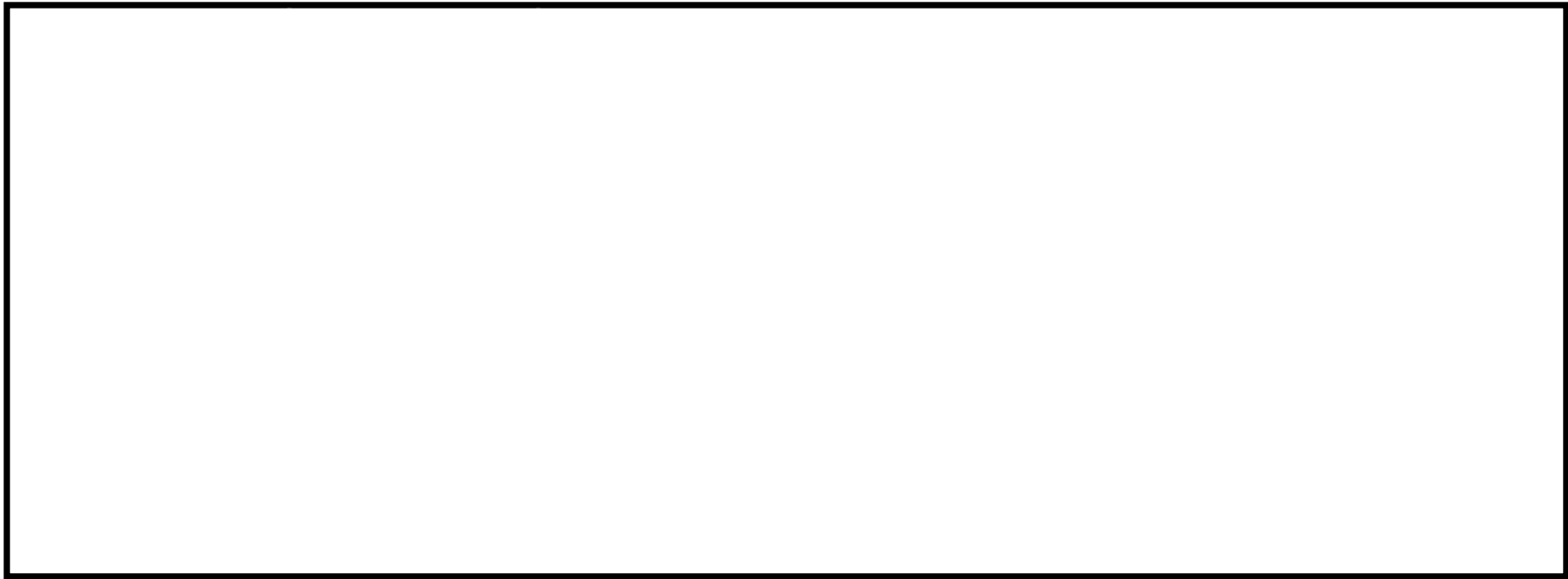


Milone+12

Binaries in GCs are rare!

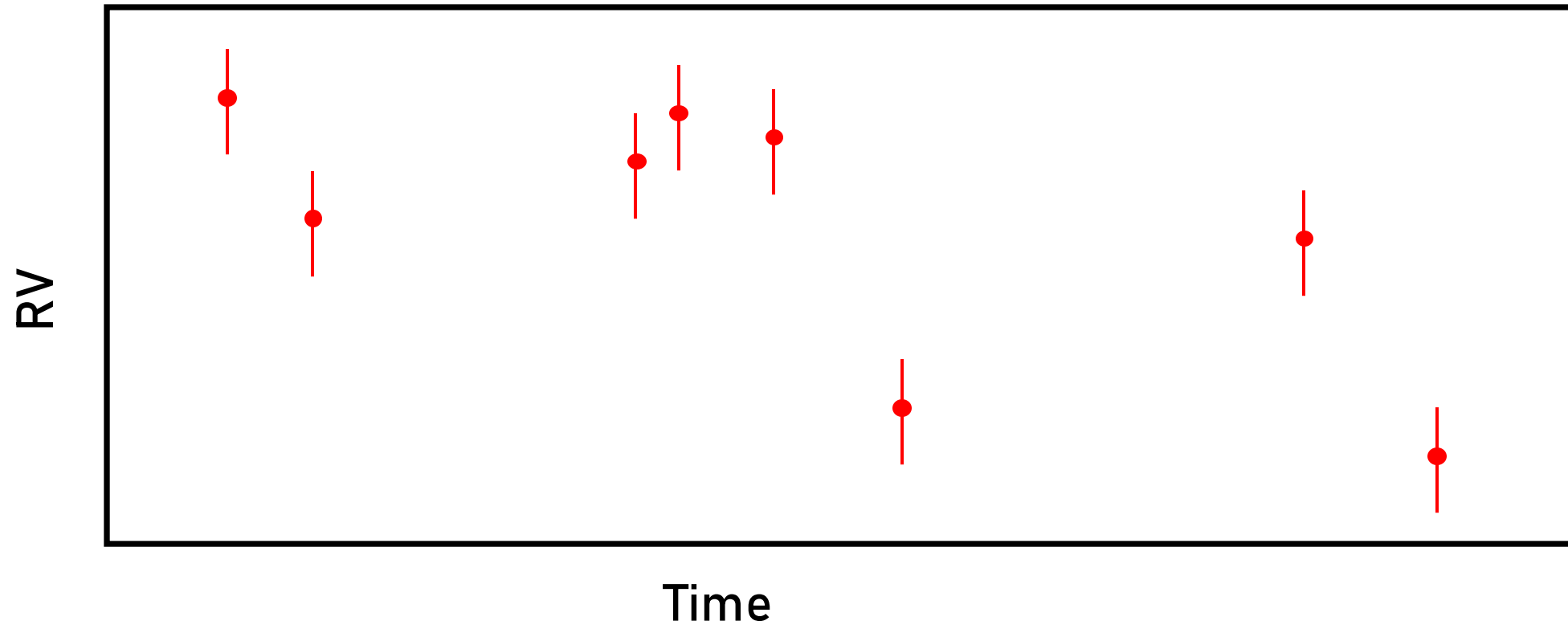
Binary fractions using *Gaia*: what we'd want

RV

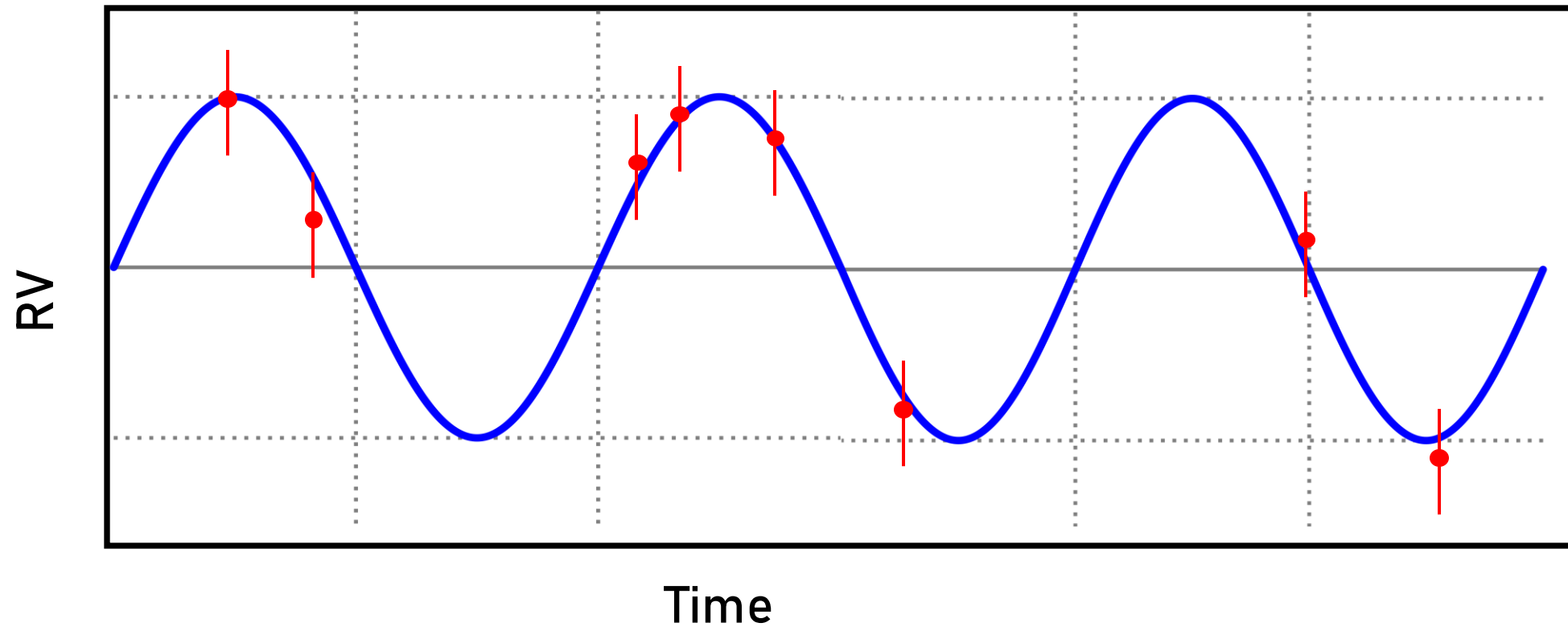


Time

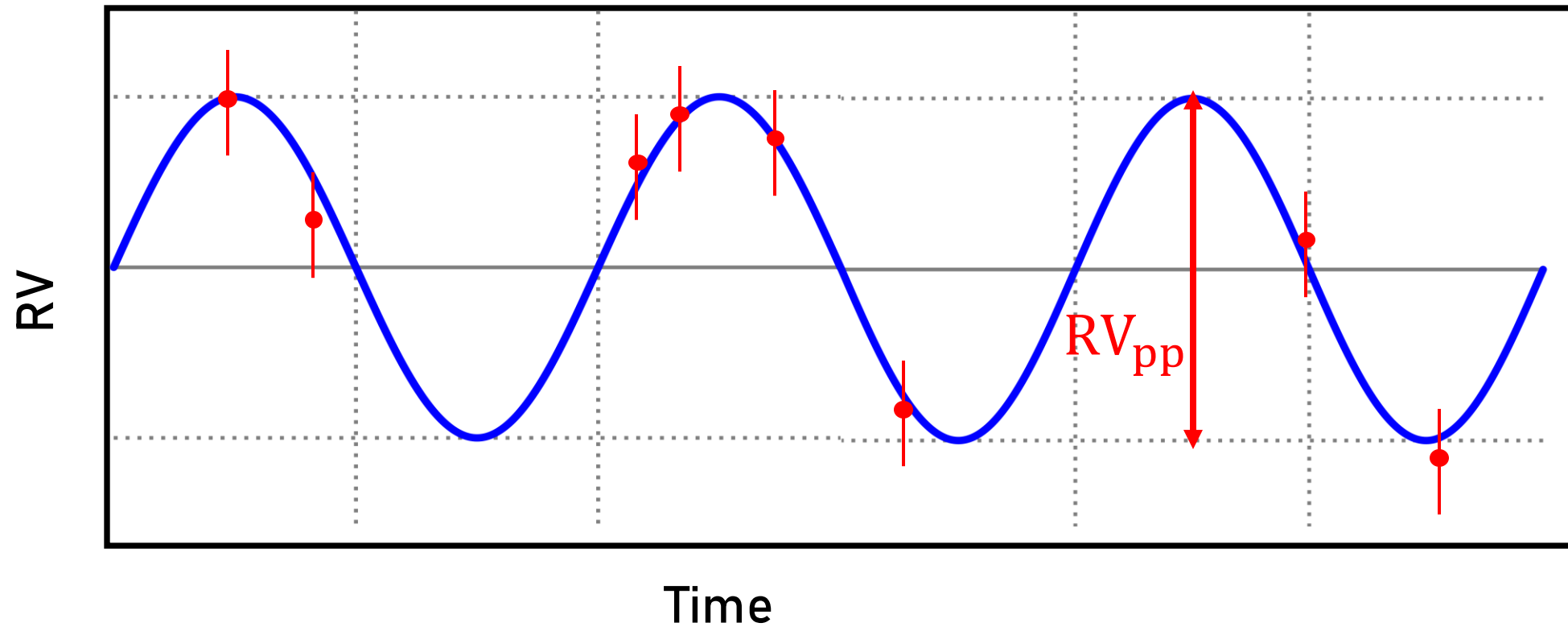
Binary fractions using *Gaia*: what we'd want



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Binary fractions using *Gaia*: what we'd want



Binary fractions using *Gaia*: what we have

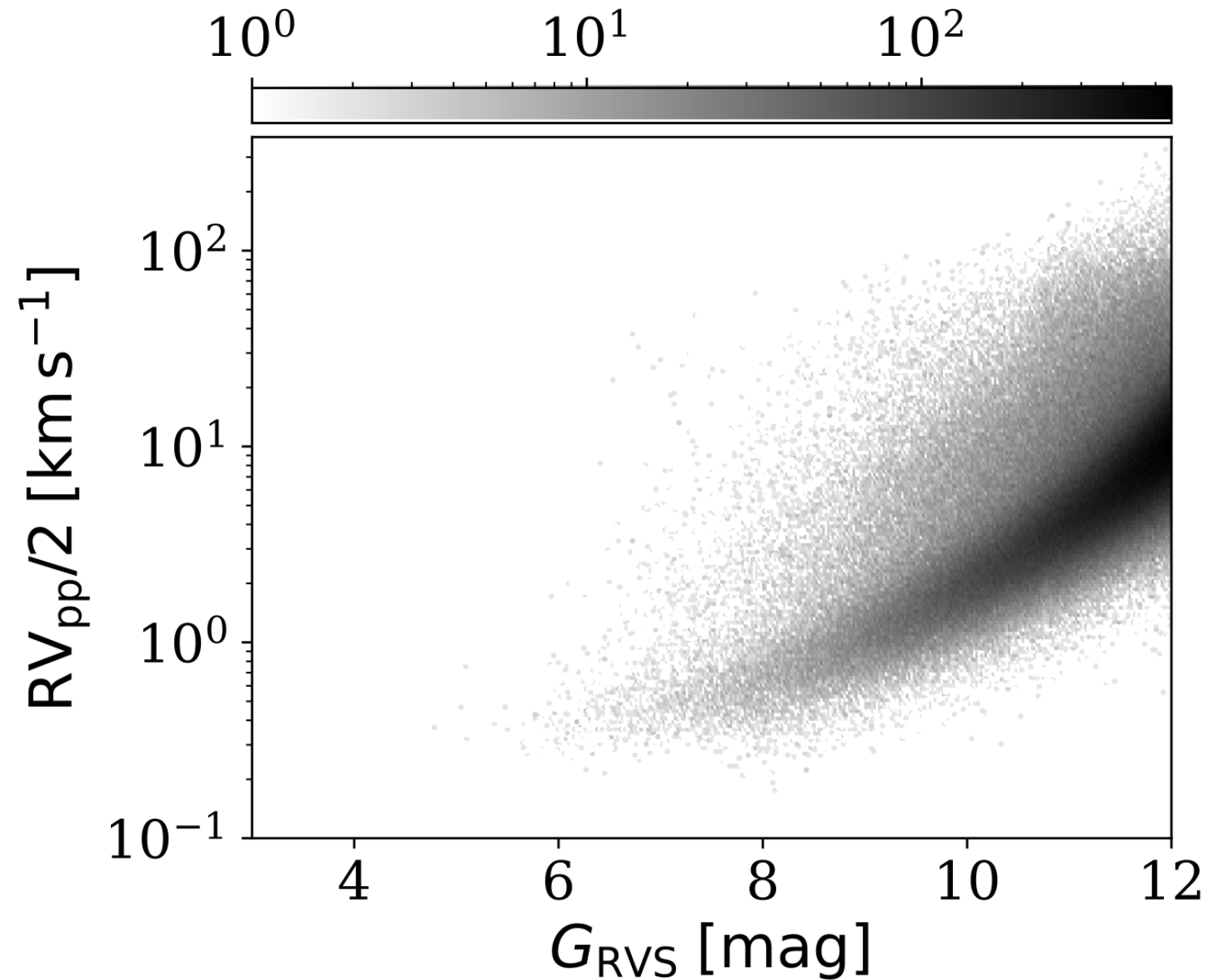
RV_{pp}



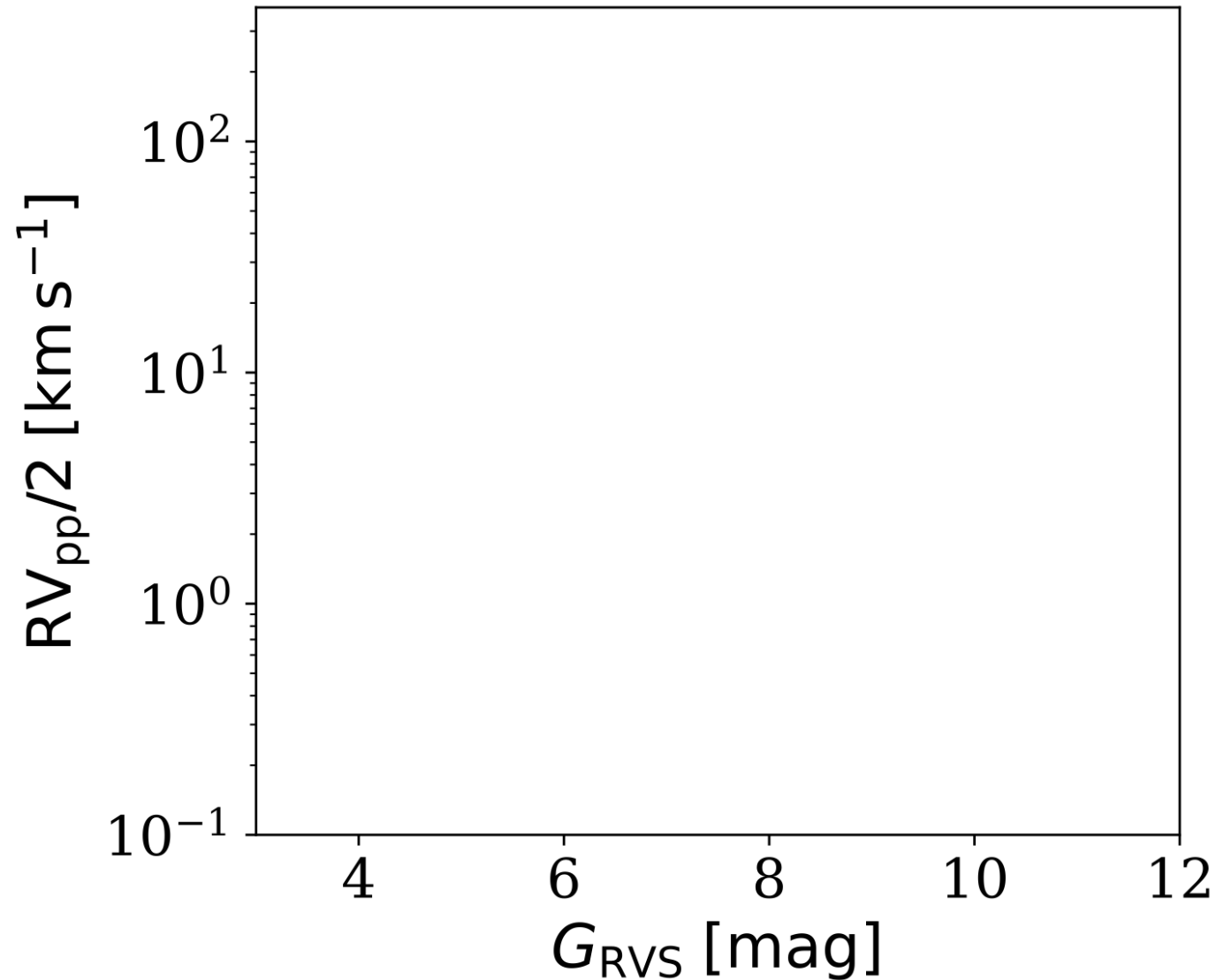
$G_{RVS} < 12; N \sim 5M$

Katz+23

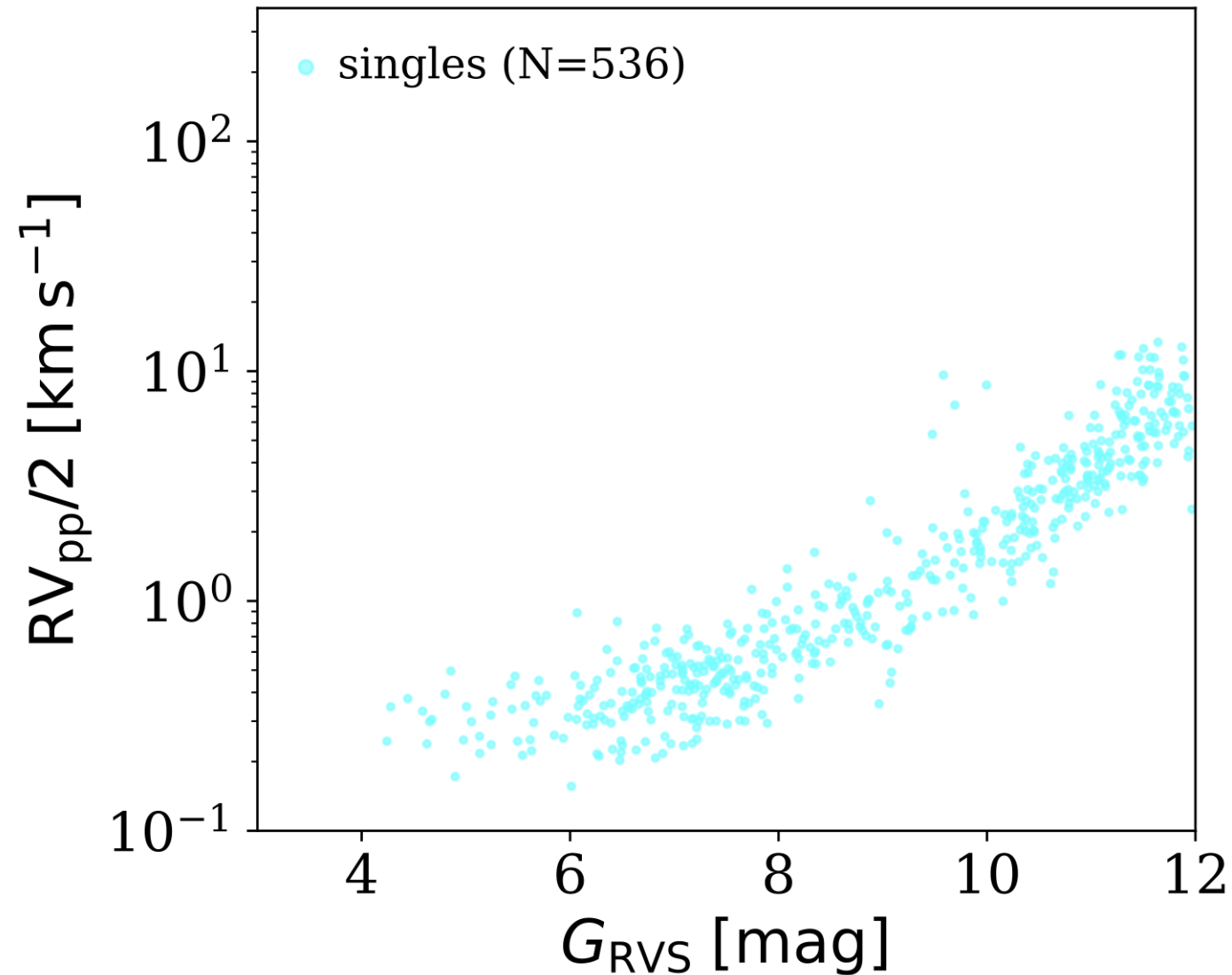
Binary fractions using *Gaia*: what we have



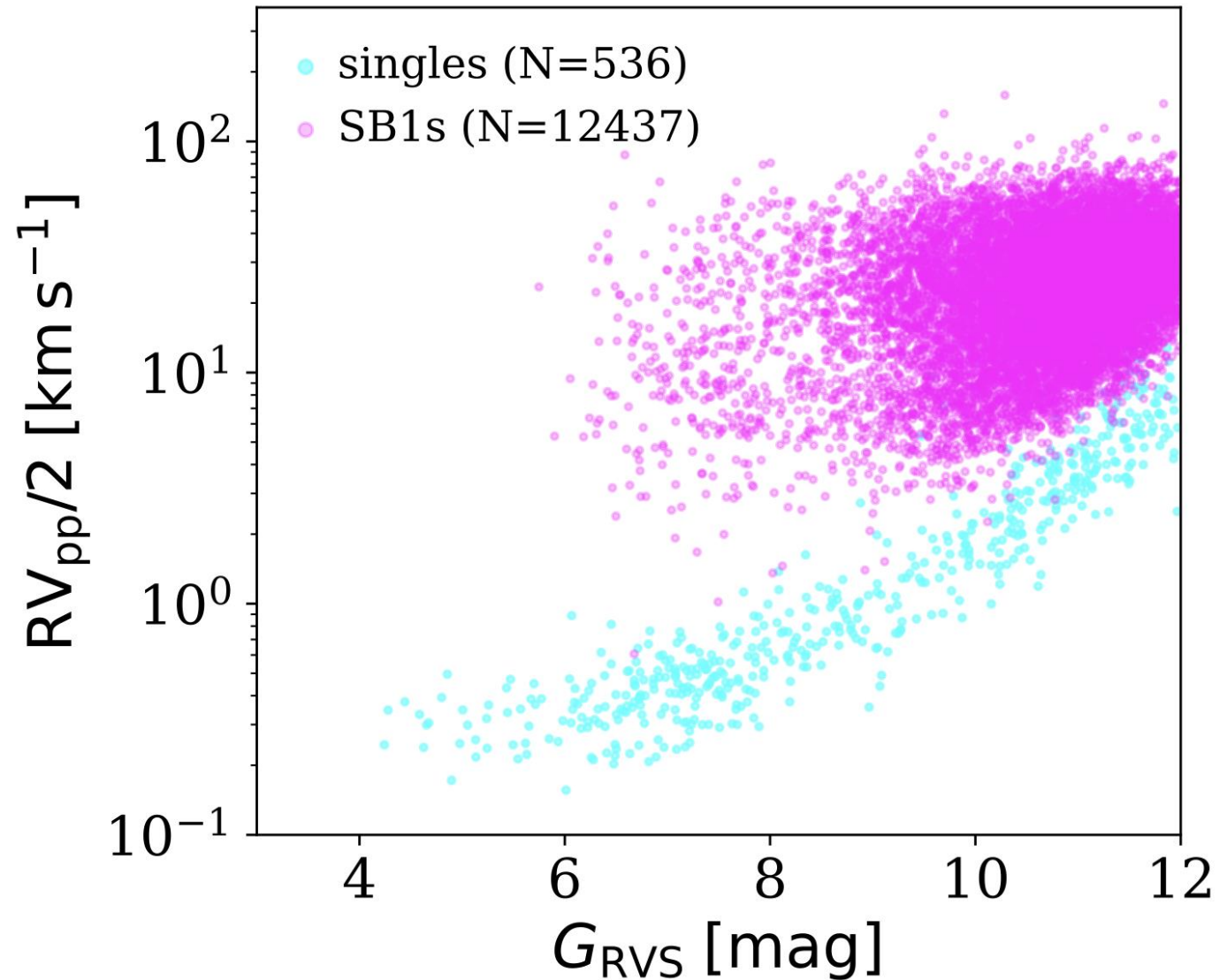
Binary fractions using *Gaia*: what we have



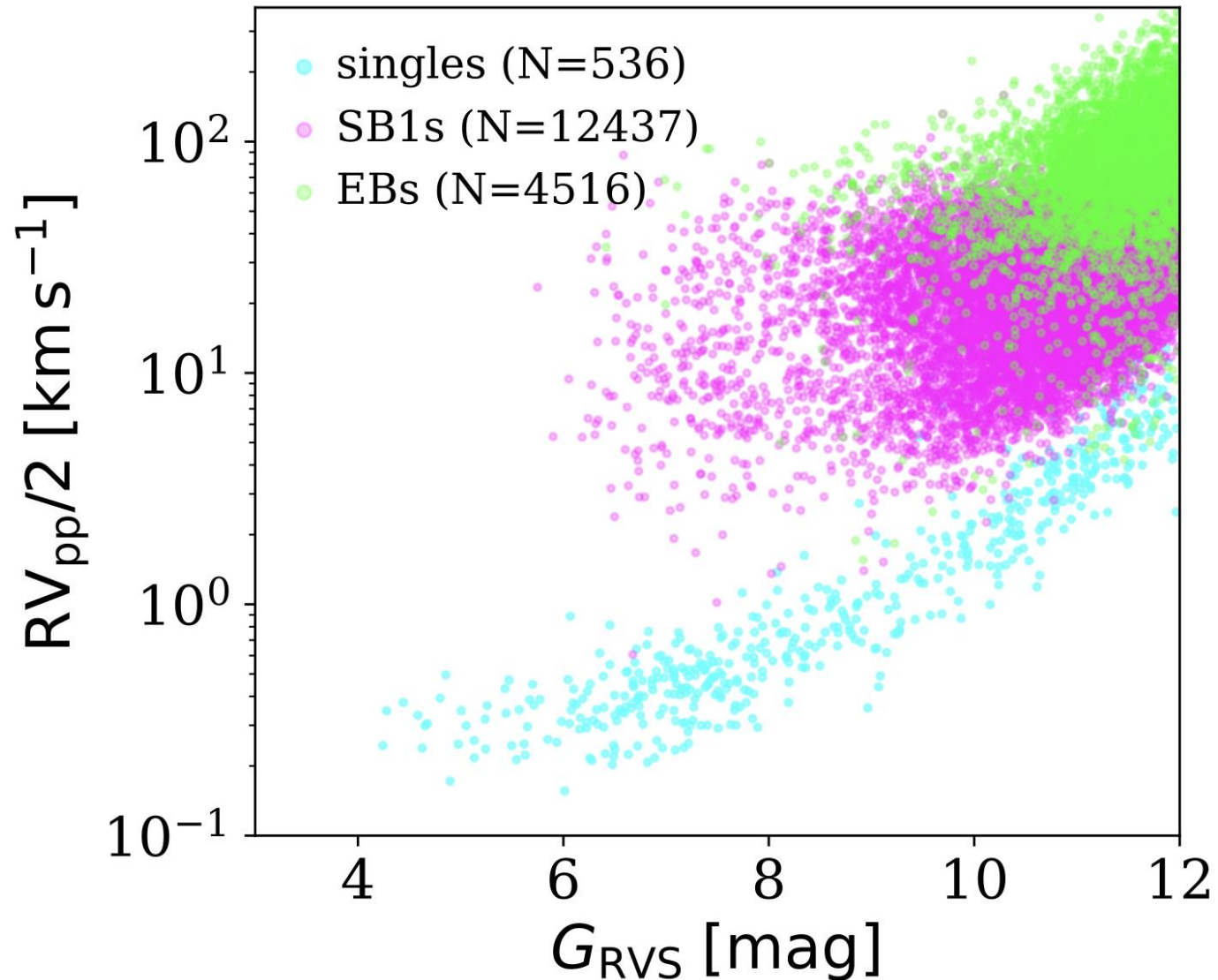
Binary fractions using *Gaia*: what we have



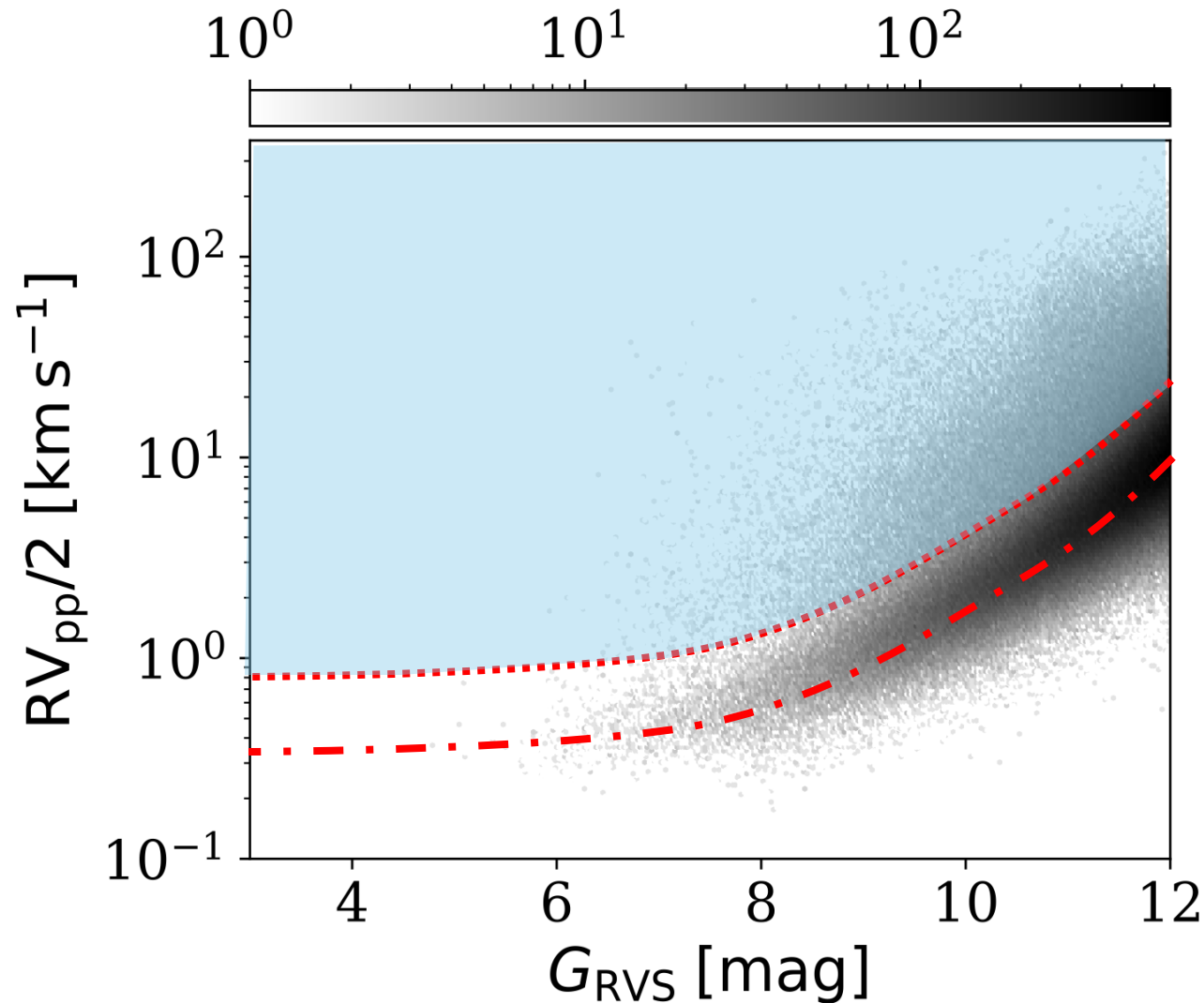
Binary fractions using *Gaia*: what we have



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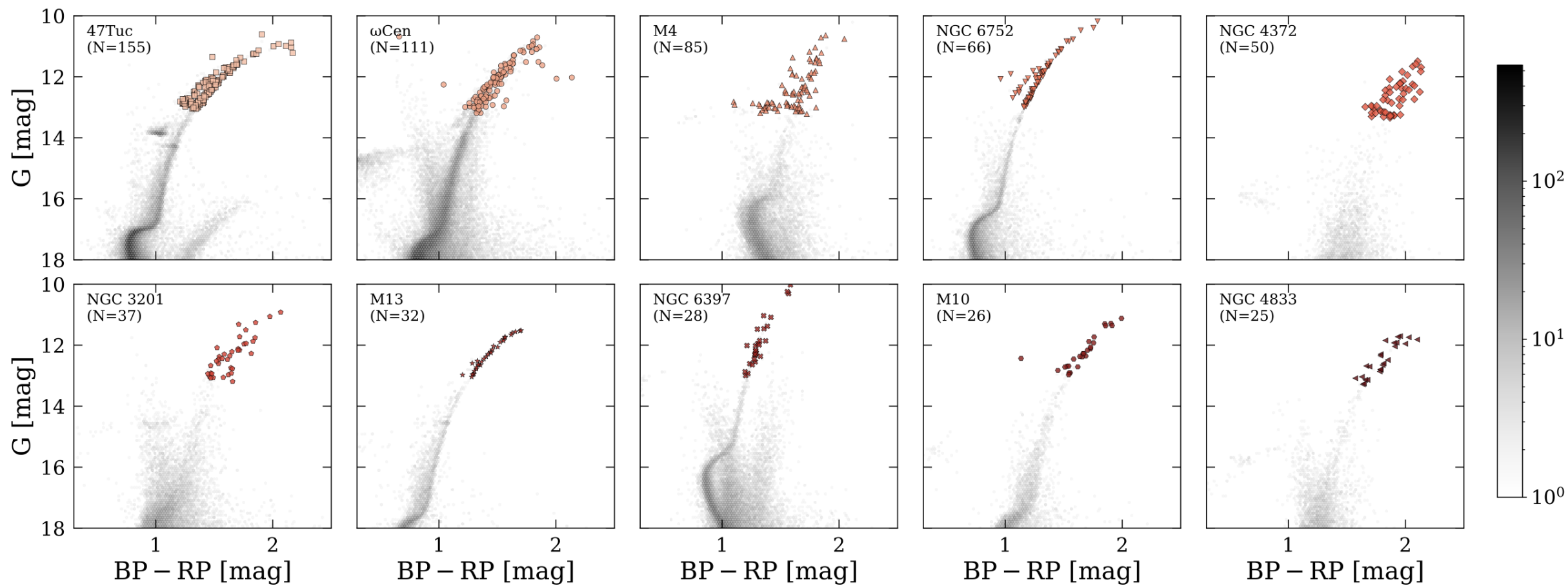
We define a sum of two Gaussian distributions: one for low-RV_{pp} (single stars) and one for high-RV_{pp} (binary stars):

$$f(x|G_{\text{RVS}}; \theta) = (1 - F) \cdot \mathcal{N}_s(x|\mu_s(G_{\text{RVS}}), \sigma_s^2) + F \cdot \mathcal{N}_b(x|\mu_b(G_{\text{RVS}}), \sigma_b^2),$$

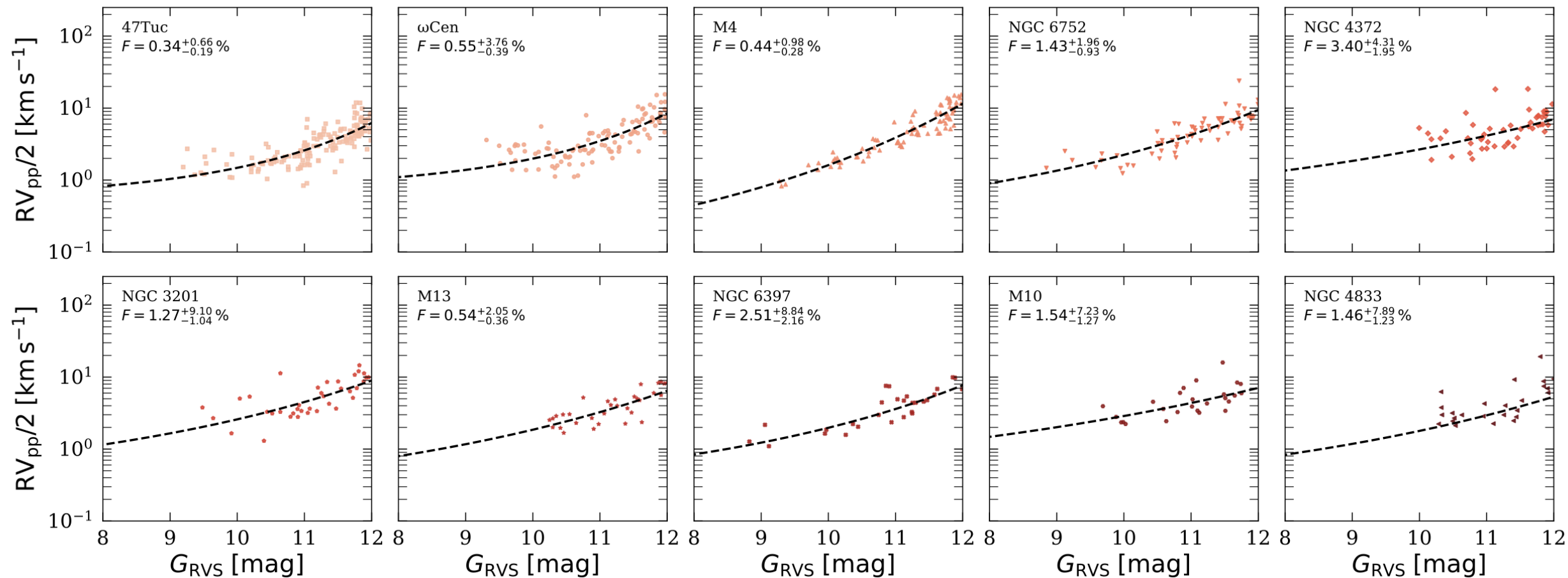
$$\theta = (F, a, b, G_{\text{min}}, d, \sigma_s, \sigma_b)$$

F — binary fraction

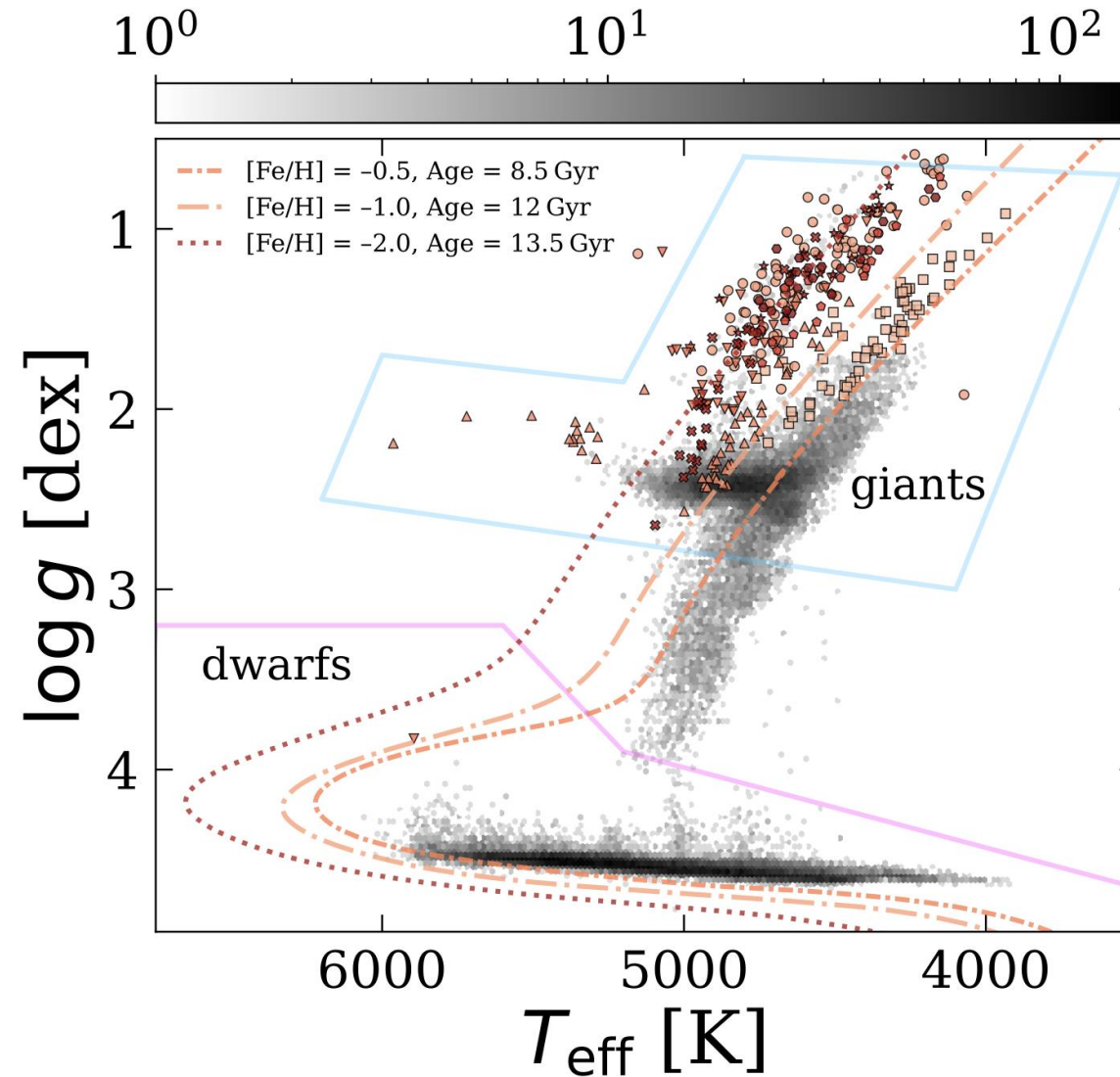
Gaia RVS view GCs



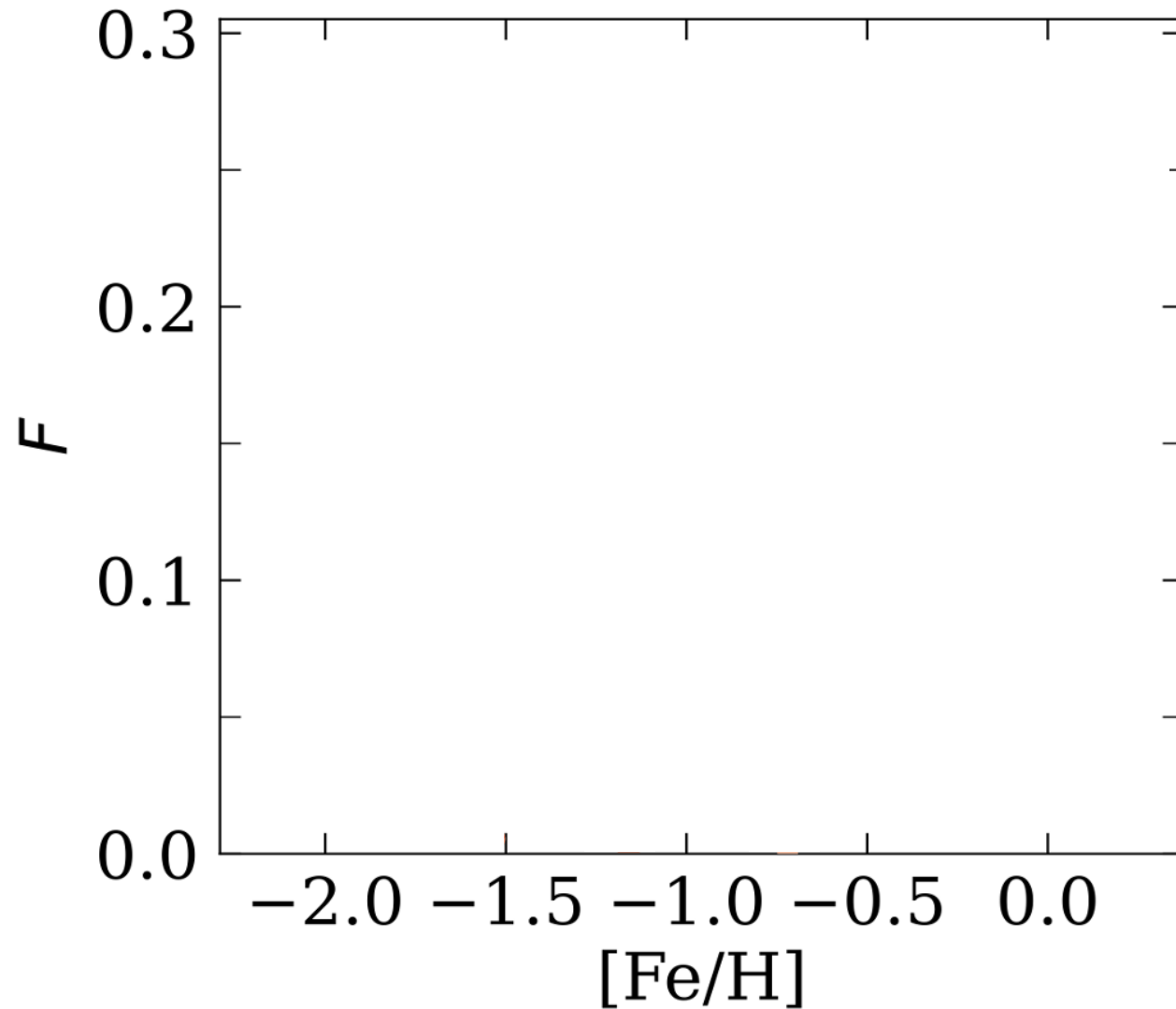
Binary fractions in GCs using *Gaia*



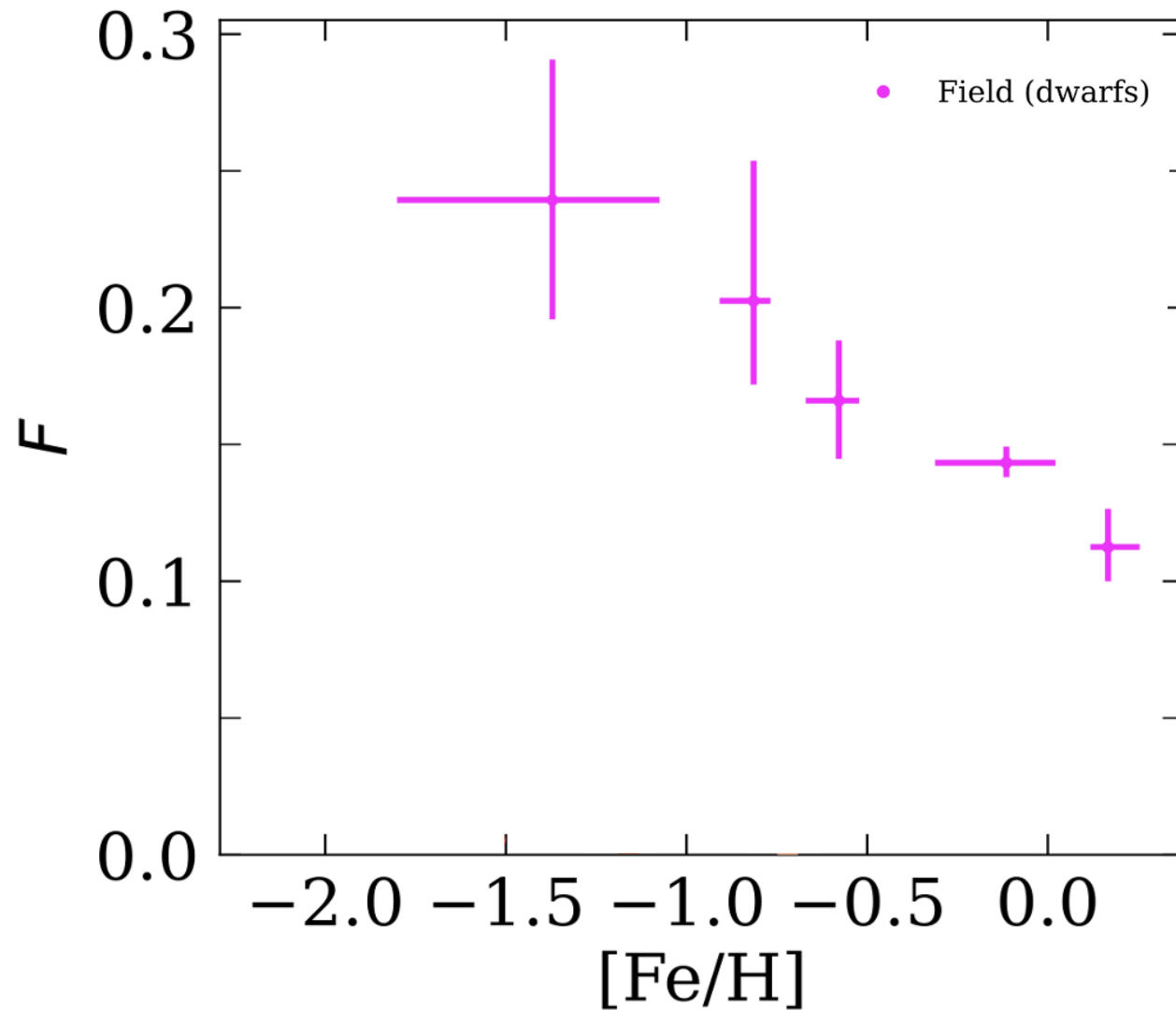
Binary fractions using *Gaia*: in GCs vs. Field



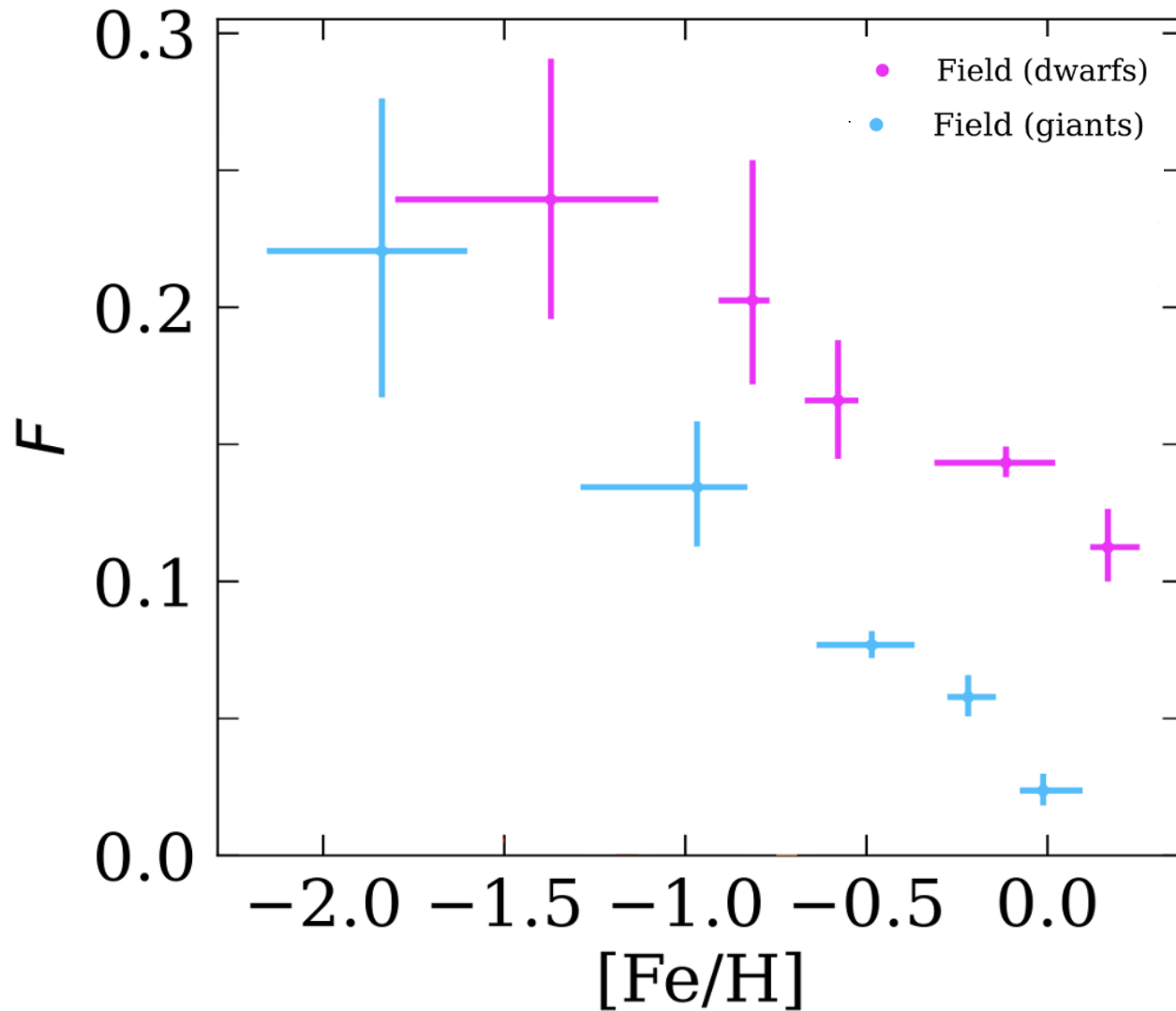
Binary fractions using *Gaia*: in GCs vs. Field



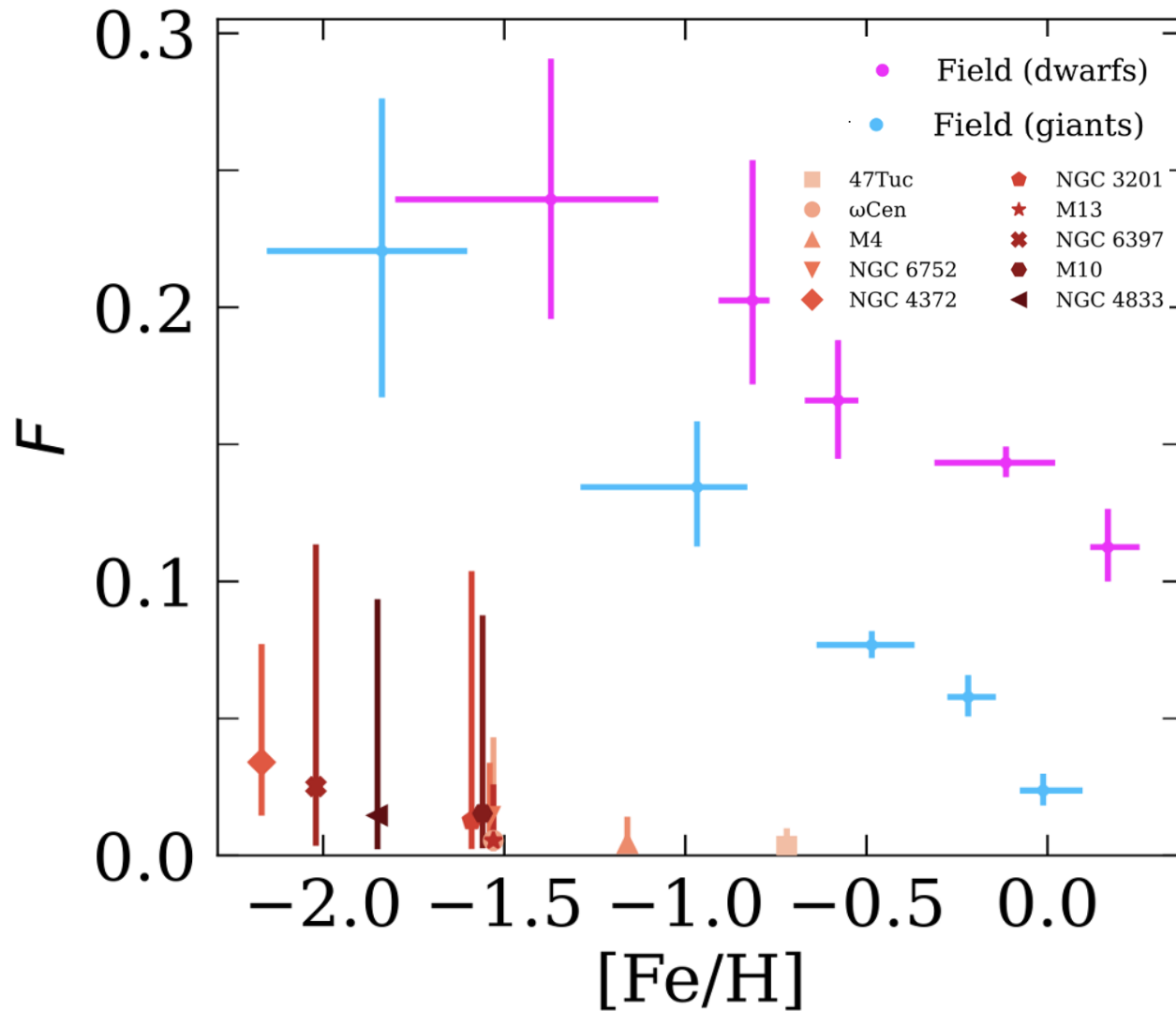
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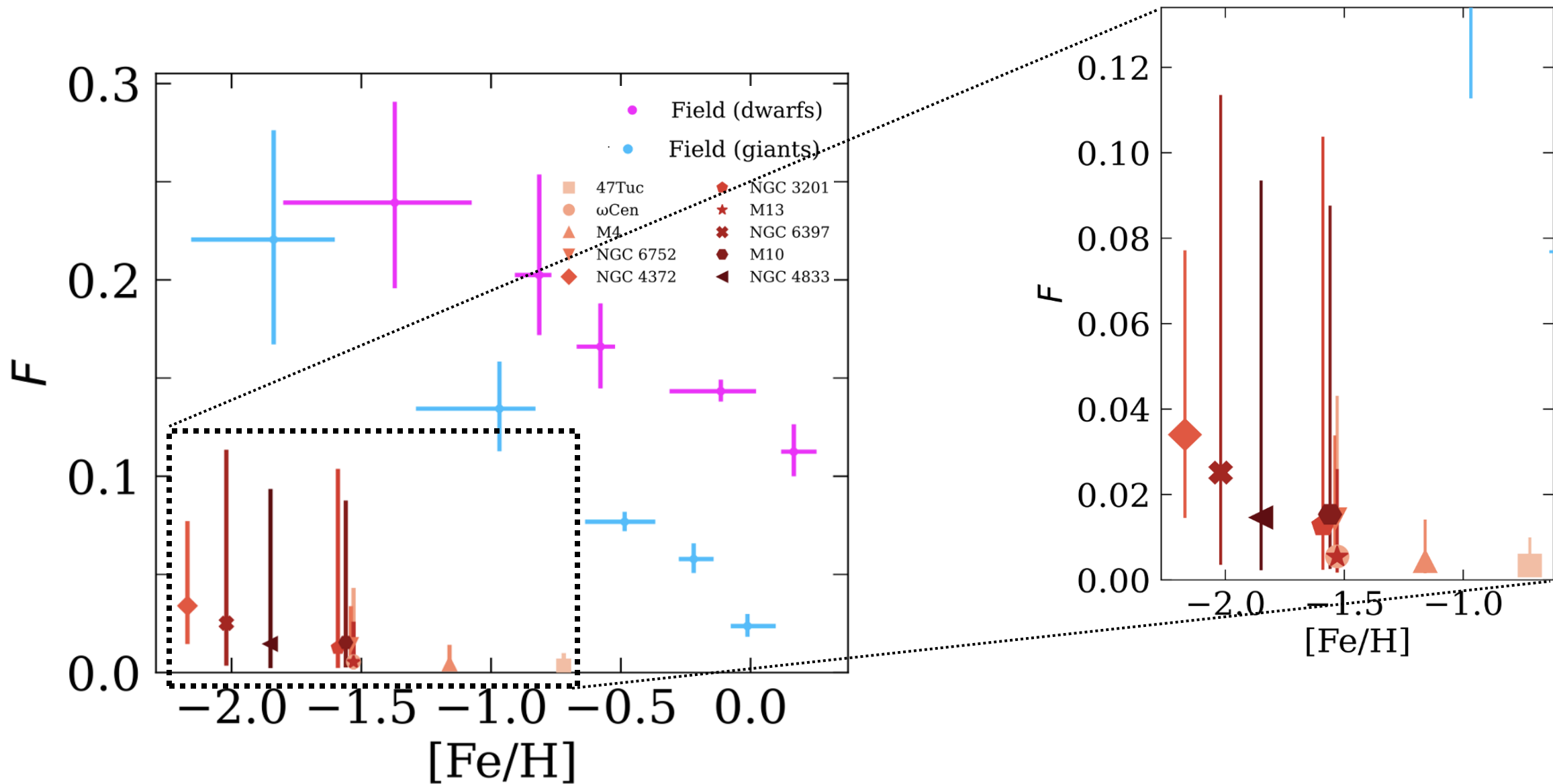
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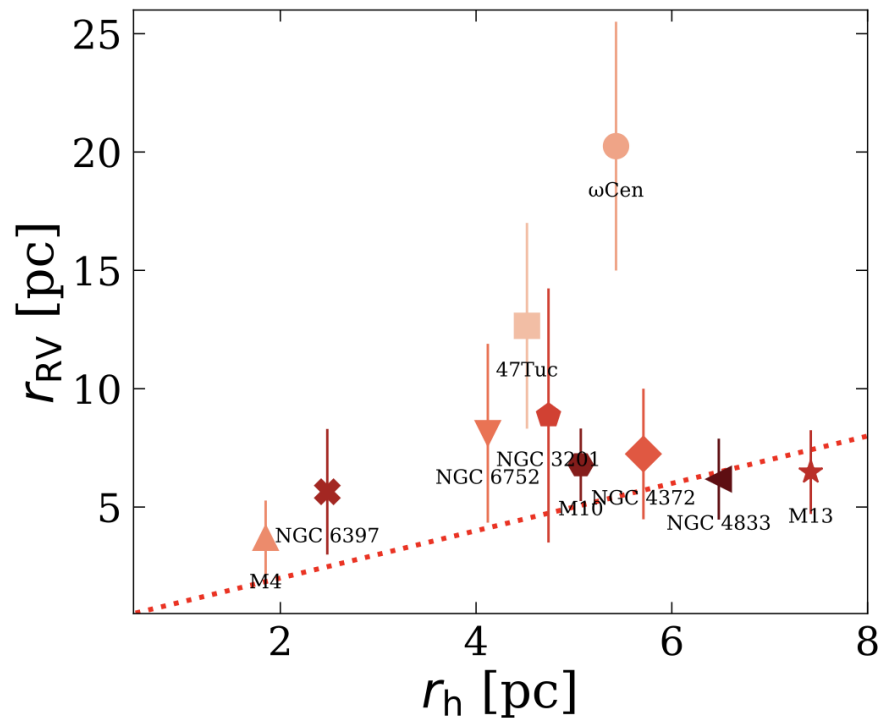
Binary fractions using *Gaia*: in GCs vs. Field



Why So Low?

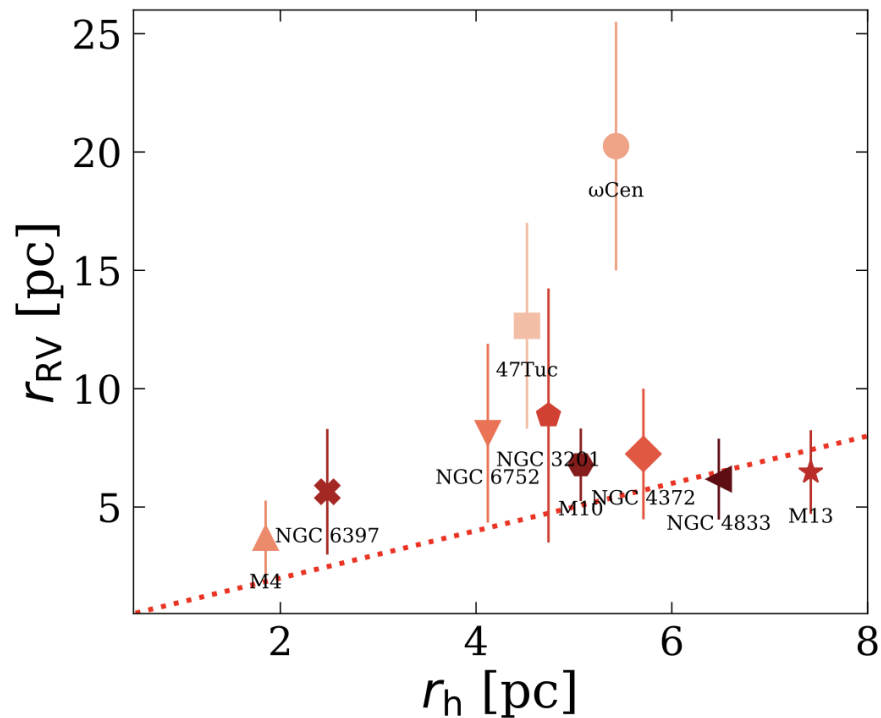
Why So Low?

- Dynamical disruption & hardening
- RV-selection bias

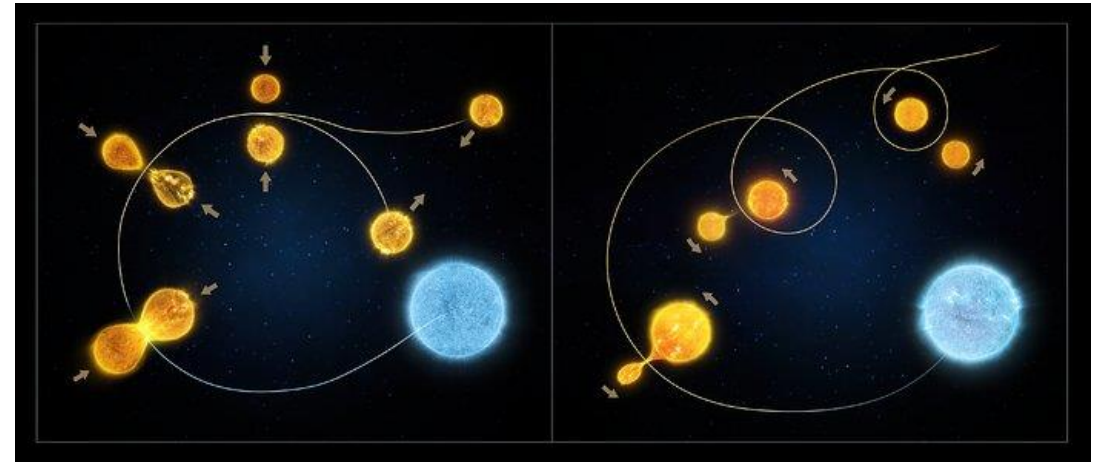


Why So Low?

- Dynamical disruption & hardening
- RV-selection bias



- Post-RGB stellar evolution



Take aways

- **A statistical framework to estimate binary fractions: provides a homogeneous benchmark in GCs and field stars**
- **Suggests that dynamical interactions dominate over metallicity in shaping binaries**
- **Prospect towards DR4: Gaia may ultimately enable orbital fitting of the detected binary stars**

Thanks for listening!