

The role of star formation-driven outflows in the baryon cycle

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And the DUVET team:

Deanne Fisher, John Chisholm, Glenn Kacprzak, Nikki Nielsen, Daniel McPherson, Magdalena Hamel Bravo, Alberto Bolatto, Danielle Berg, Alex Cameron, Karin Sandstrom, Anna McLeod, and others

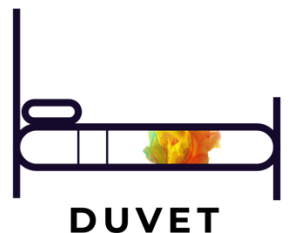


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Astrophysics and
Supercomputing



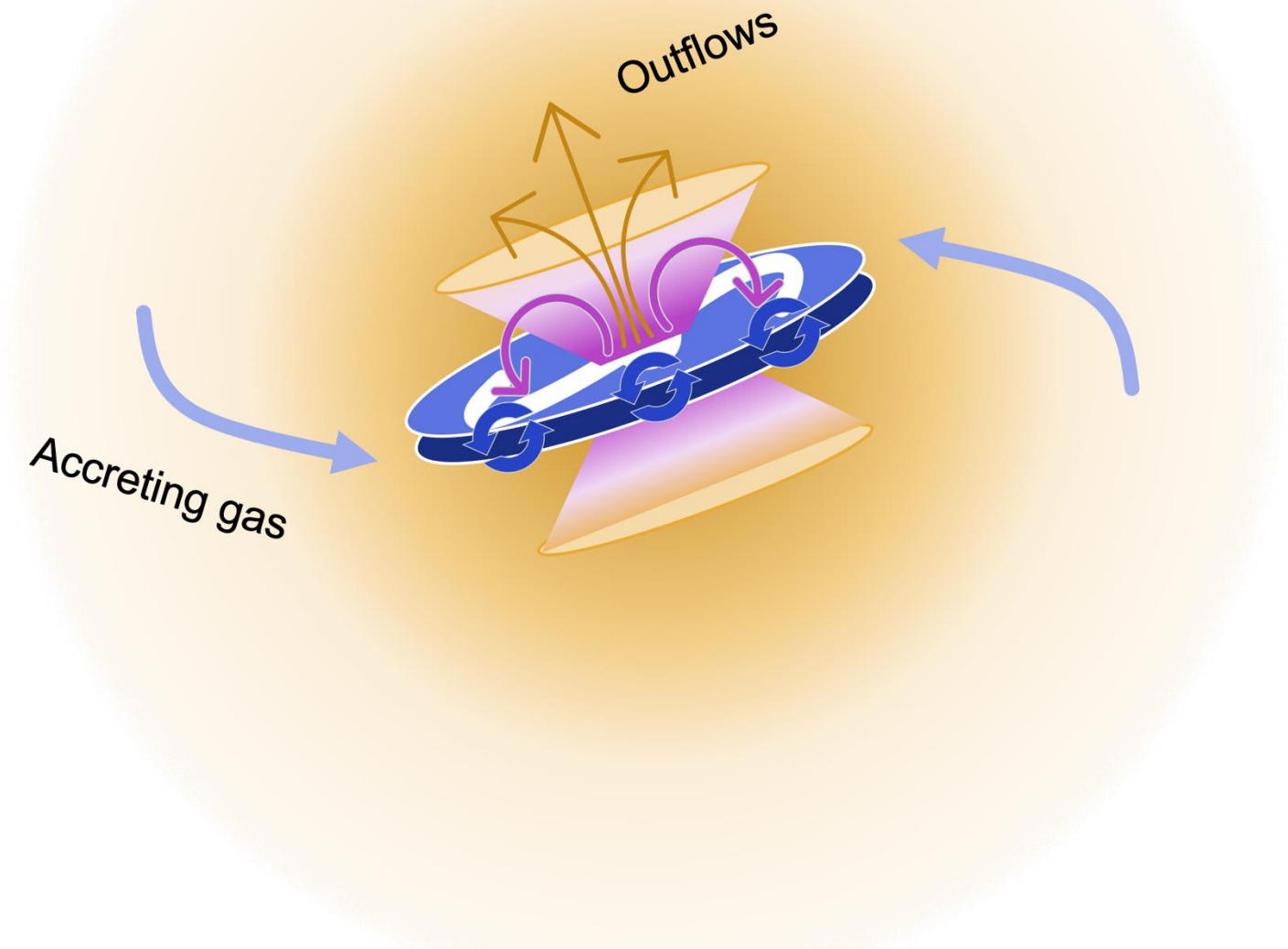
Durham
University

ASTRO3D





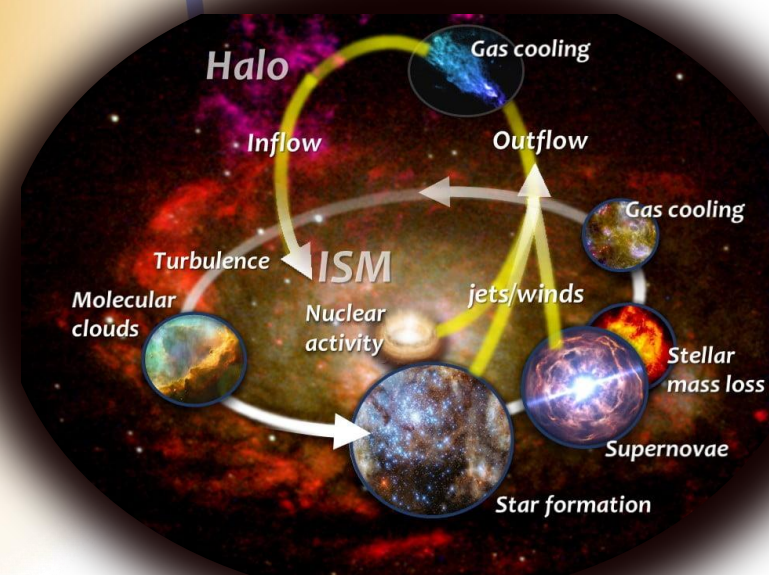
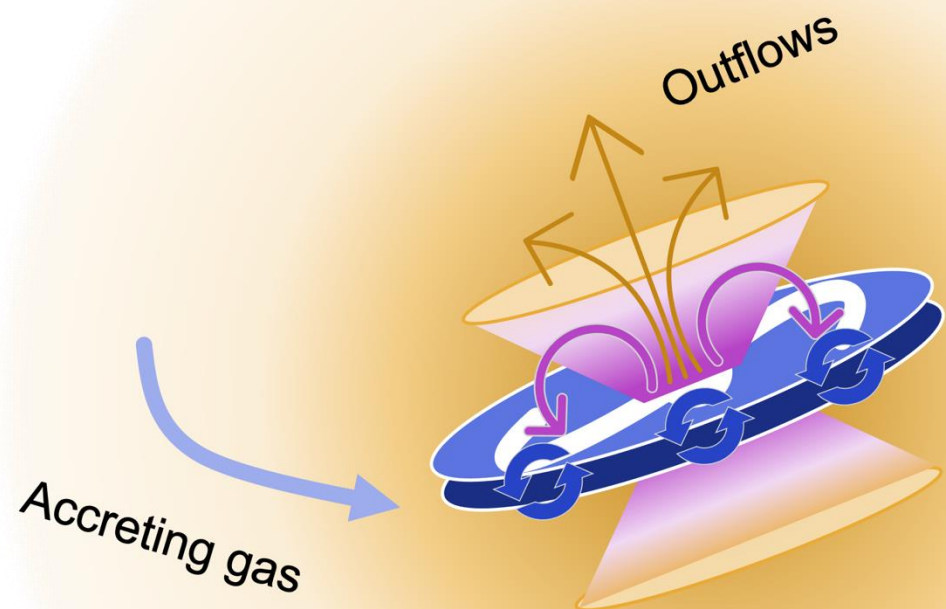
DUVET
Face-on





DUVET

Face-on





M82

Image Credit: X-ray: NASA/CXC/JHU/D. Strickland;
Optical: NASA/ESA/STScI/AURA/The Hubble Heritage Team;
Infrared: NASA/JPL-Caltech/Univ. of Arizona/C. Engelbracht

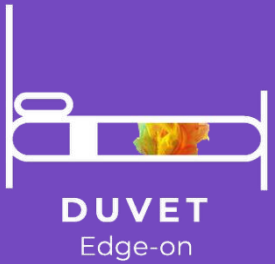


We need more resolved observations of local star formation-driven outflows

M82

Image Credit: X-ray: NASA/CXC/JHU/D. Strickland;
Optical: NASA/ESA/STScI/AURA/The Hubble Heritage Team;
Infrared: NASA/JPL-Caltech/Univ. of Arizona/C. Engelbracht

DUVET: We've got you covered!



PI : Deanne Fisher (Swinburne)

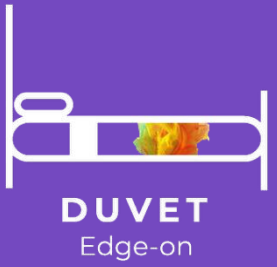
- Disks
- Starbursting: $SFR = 5-15 \times MS$
- Stellar Mass: $\log M_* = 9-11 M_\odot$
- Compact: $r_e = 1-3 \text{ kpc}$
- Nearby: $z = 0.015-0.03$
- Metallicity: $0.01-1 Z_\odot$

27 galaxies observed with KCWI/Keck II

Some overlap with HST observations



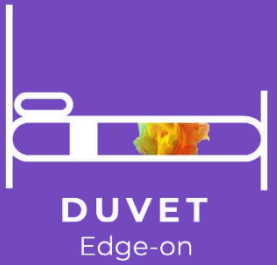
Observations of outflows using edge-on galaxies



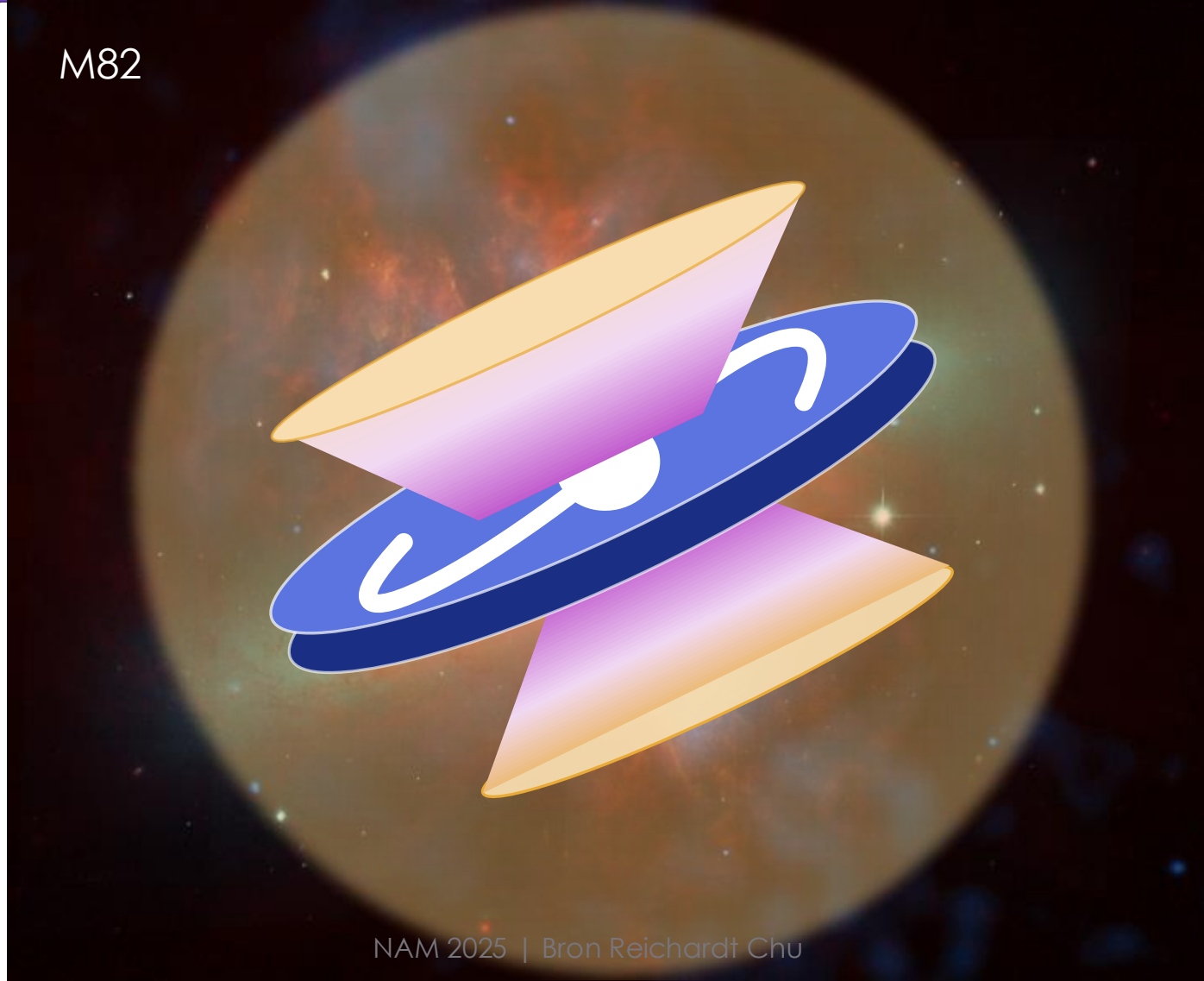
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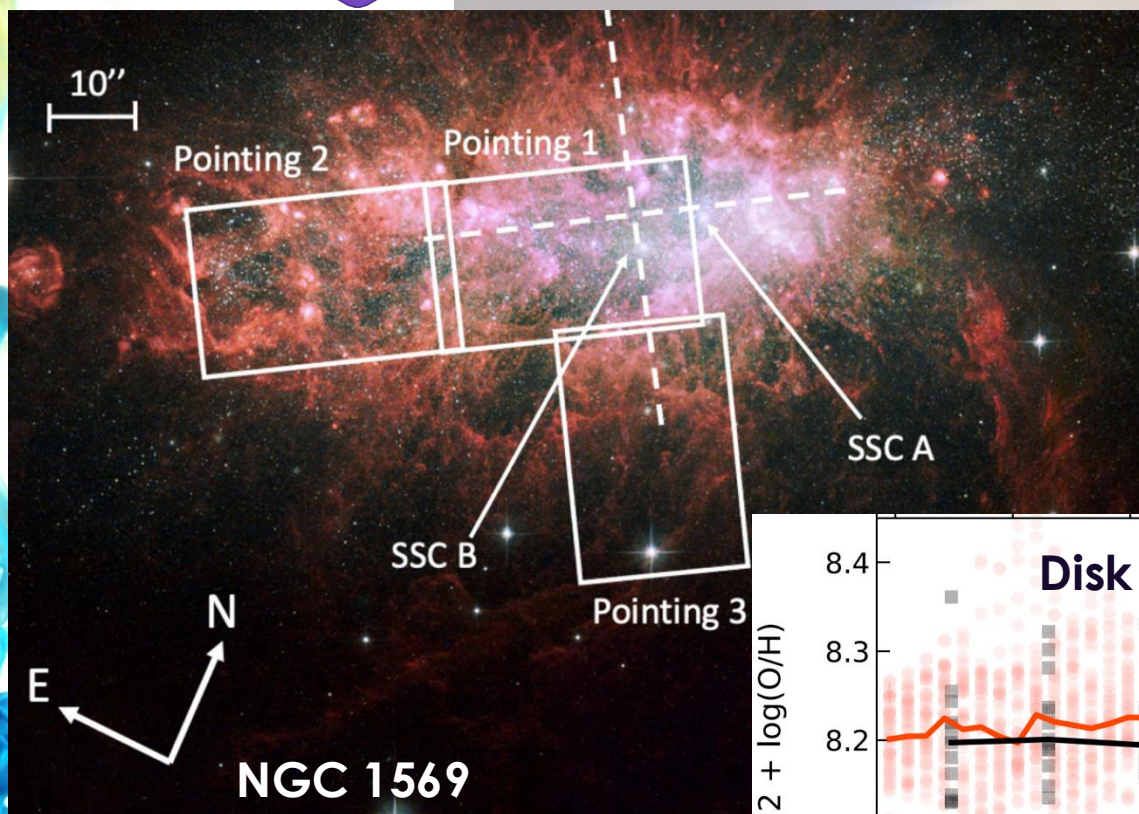
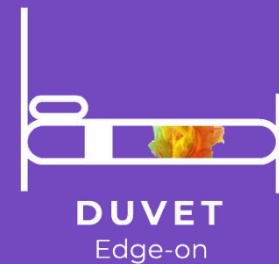
Observations of outflows using edge-on galaxies



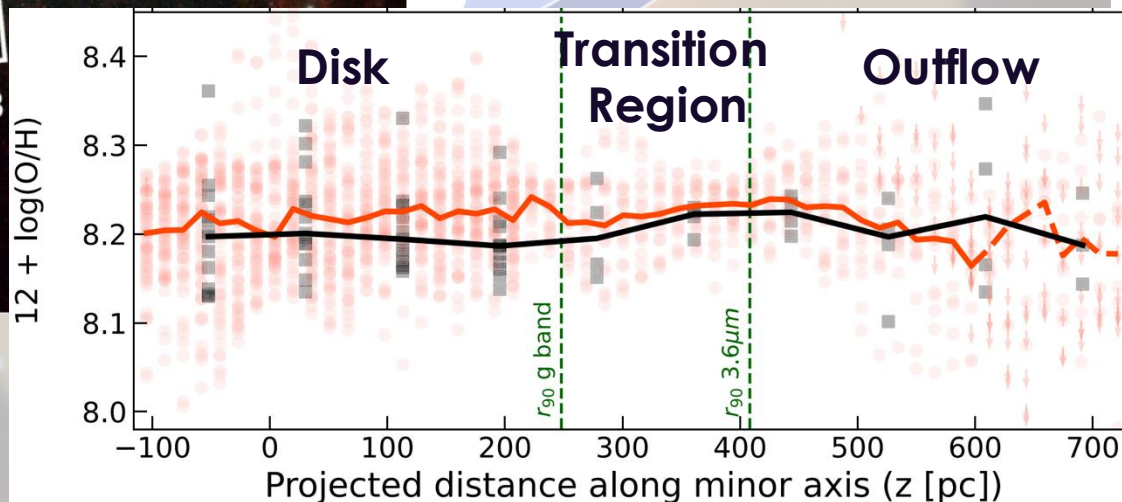
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Observations of outflows using edge-on galaxies gives us metal loading

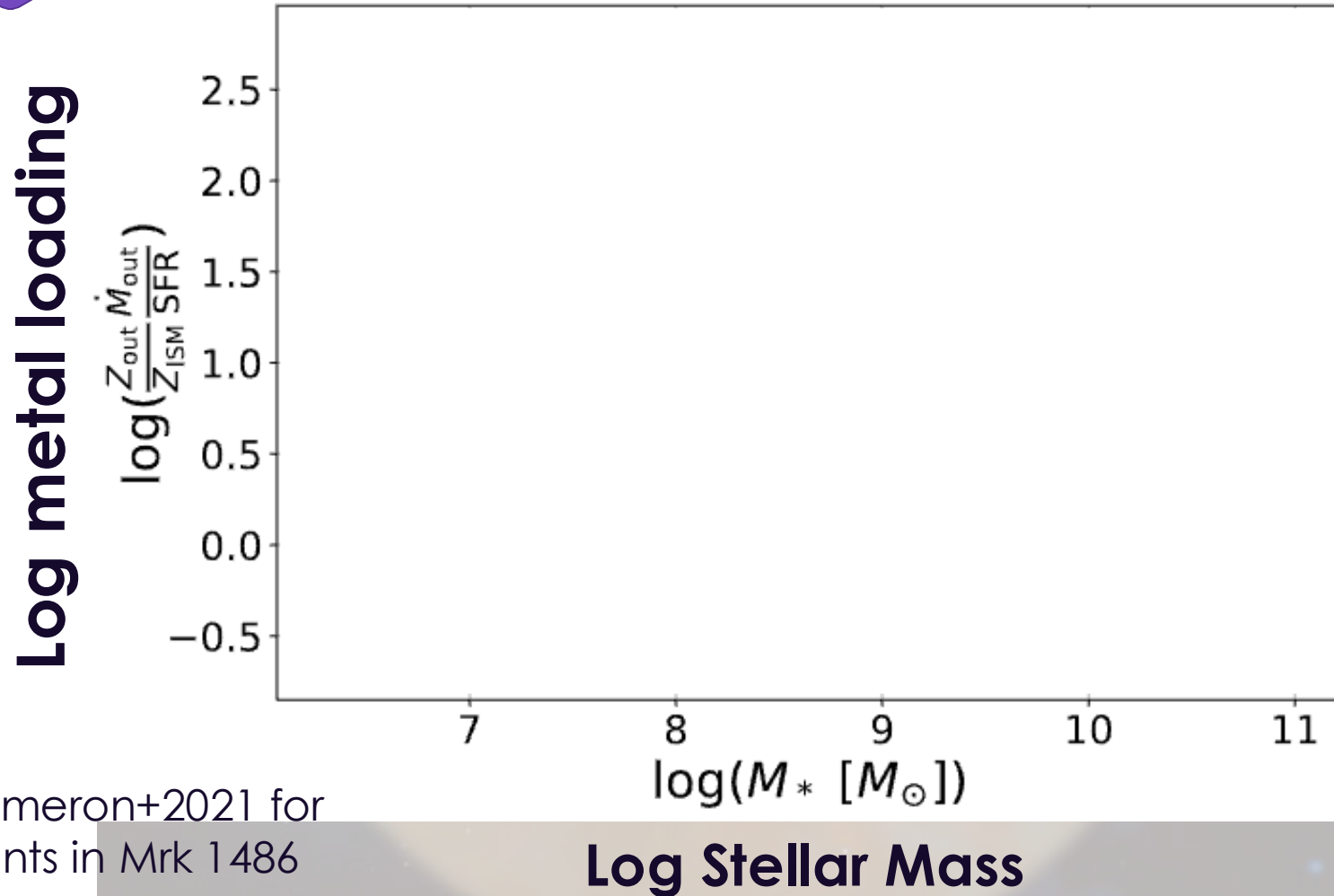
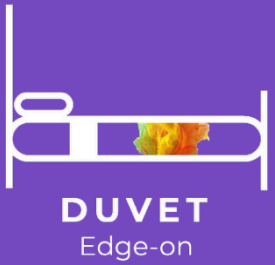


See also Cameron+2021 for measurements in Mrk 1486



Magdalena Hamel-Bravo
Hamel-Bravo+2024,
MNRAS 530, 3855

Observations of outflows using edge-on galaxies gives us metal loading



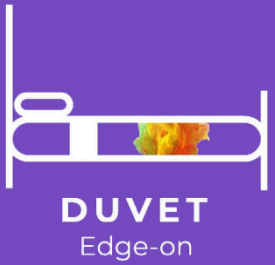
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Magdalena Hamel-Bravo
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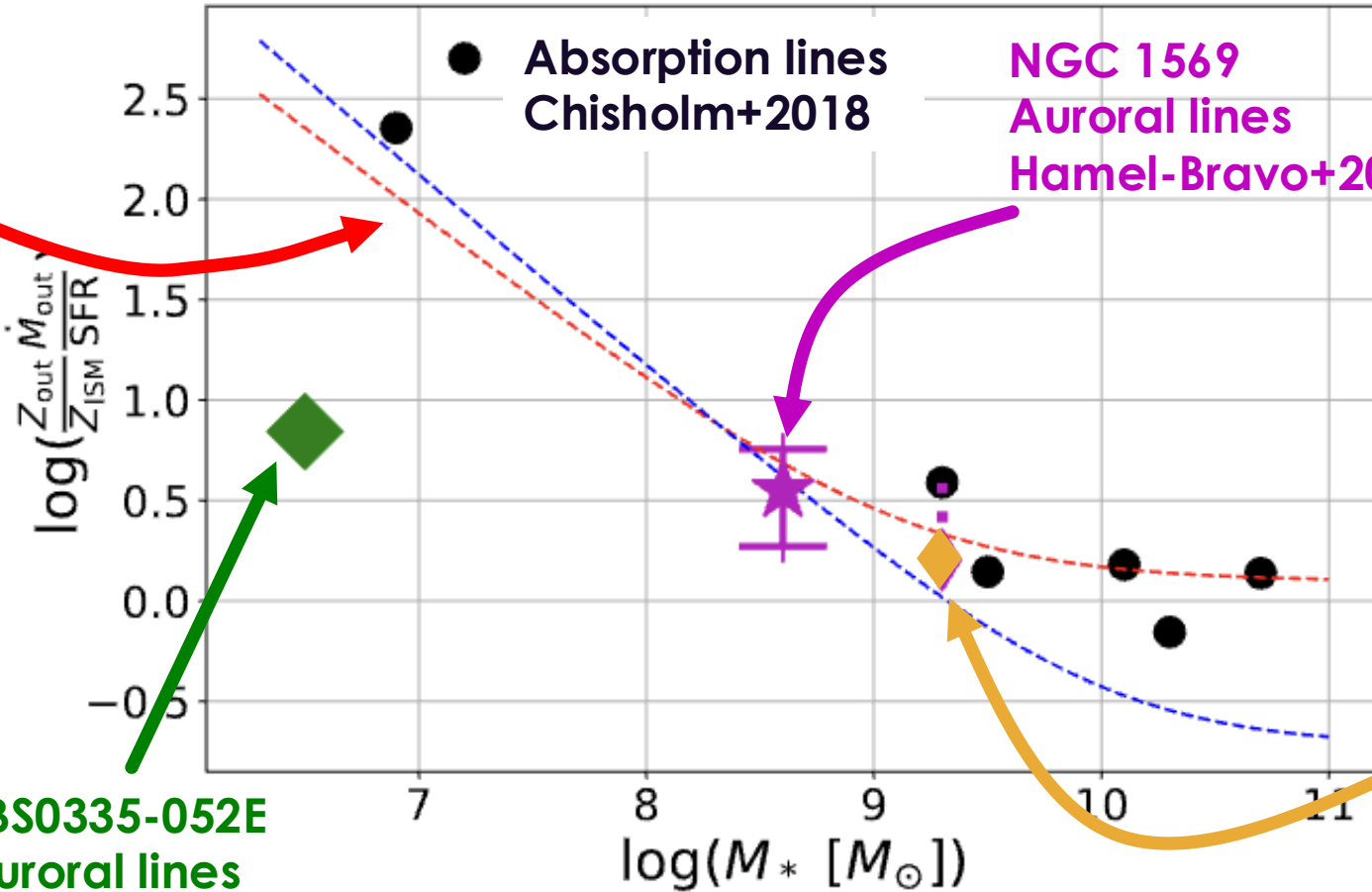
Hamel-Bravo et al.
(in prep)

Observations of outflows using edge-on galaxies gives us metal loading



Theory
necessary to
reproduce
the MZR

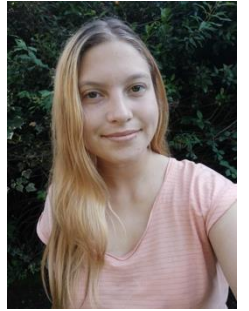
Log metal loading



SBS0335-052E
Auroral lines

See also Cameron+2021 for
measurements in Mrk 1486

Log Stellar Mass

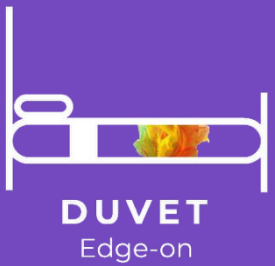


Magdalena Hamel-Bravo
Hamel-Bravo+2024,
MNRAS 530, 3855

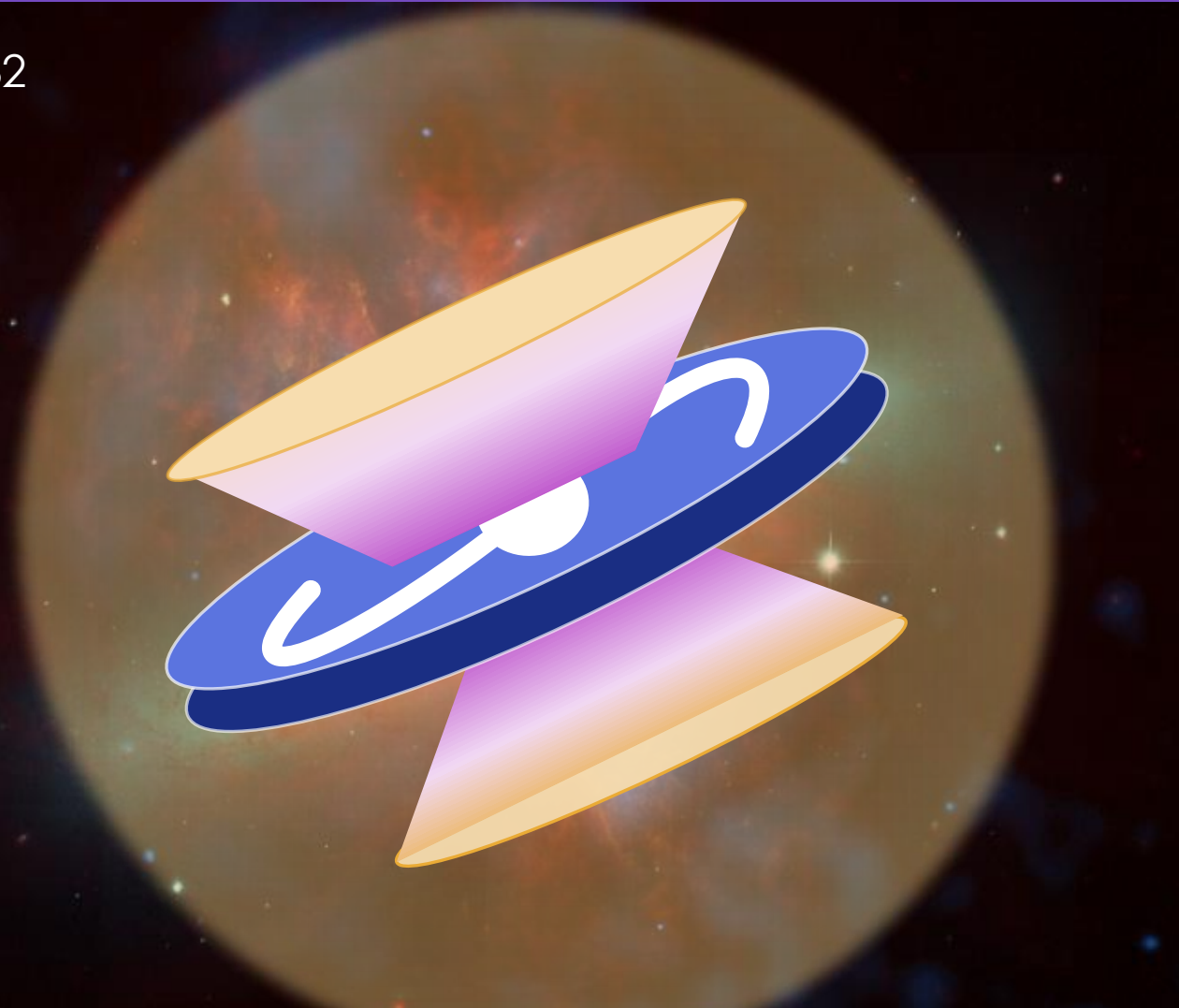
Hamel-Bravo et al.
(in prep)

Mrk 1486
Auroral lines
Cameron+2021

Observations of outflows using edge-on galaxies



M82



Measuring geometry
McPherson+2023,
MNRAS 525, 6170

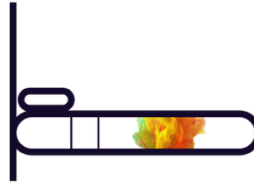
**Measuring
metallicities and
ionisation**
Hamel-Bravo+2024,
MNRAS 530, 3855

Cameron+2021,
ApJL 918, L16

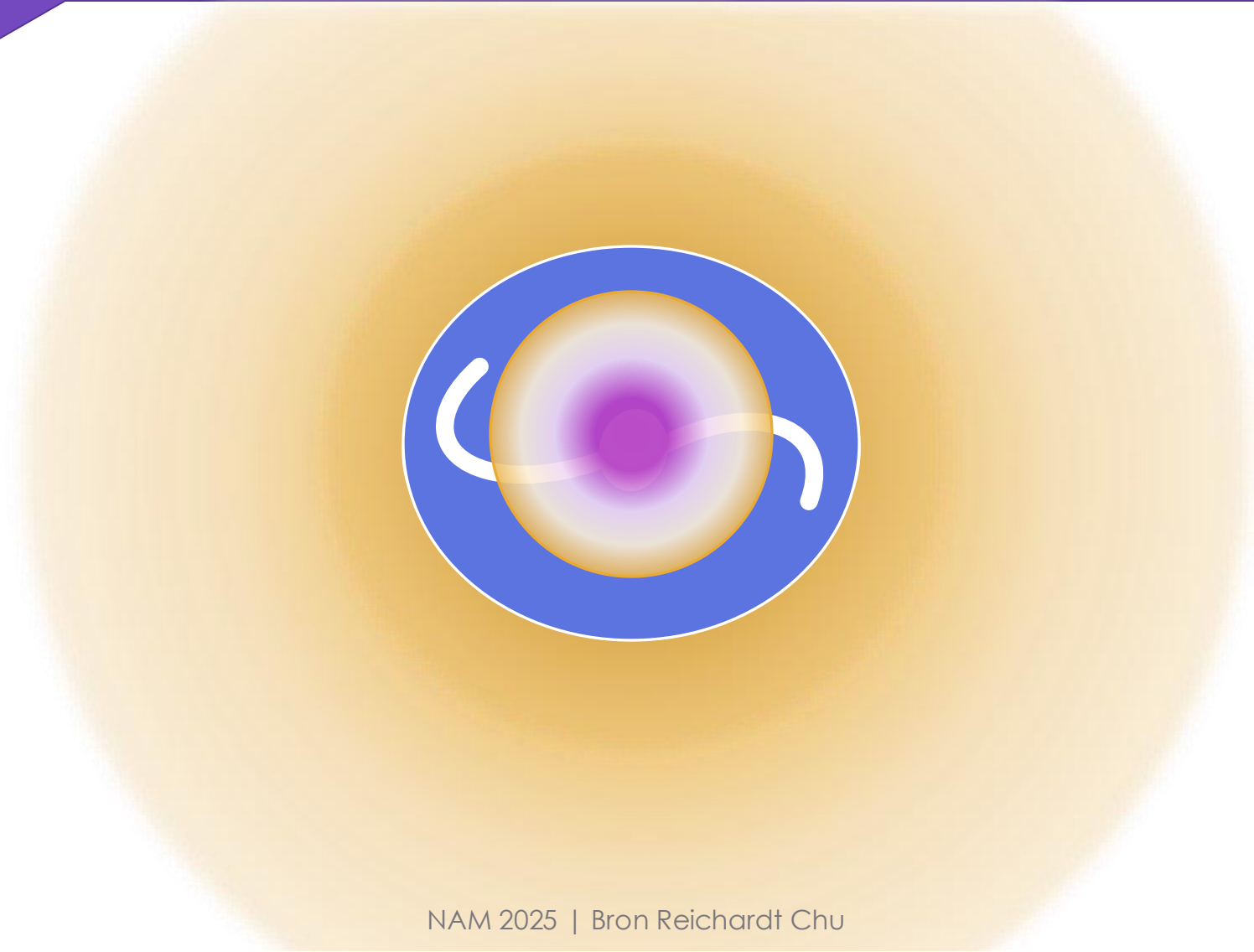
Observations of outflows using face-on galaxies



DUVET
Face-on

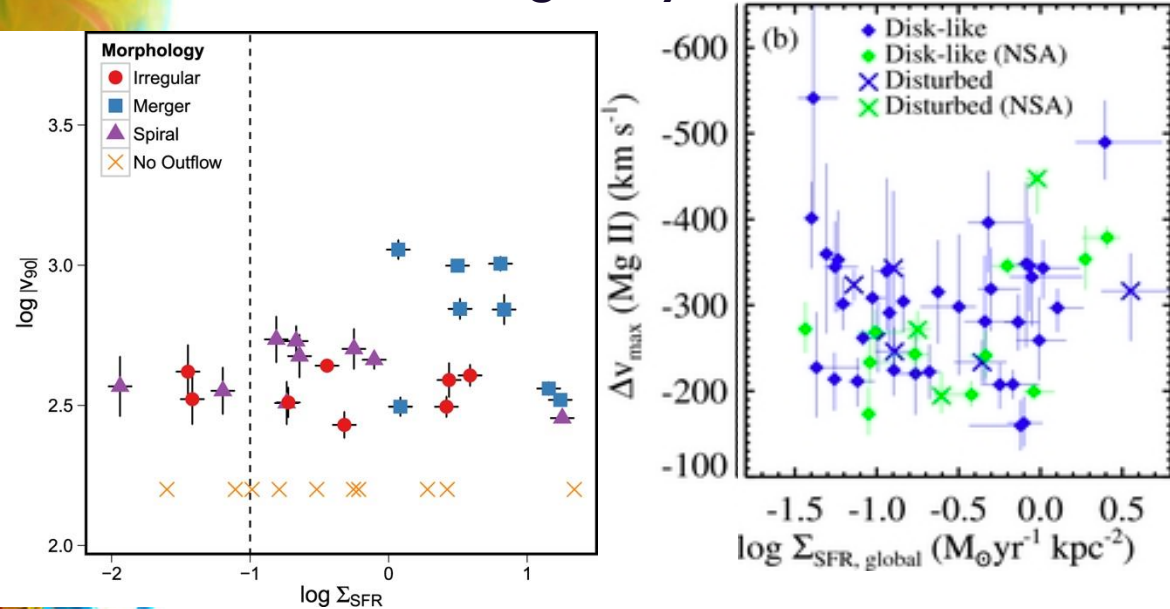


DUVET
Edge-on

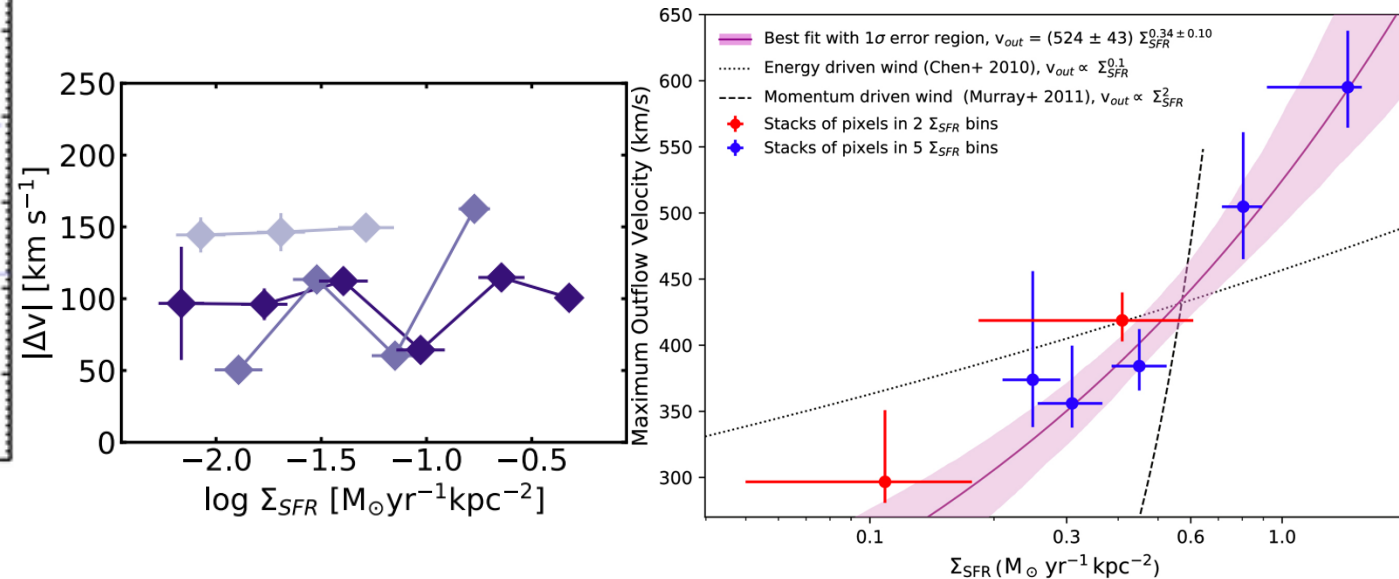


Outflow scaling relations

Global galaxy measurements



Stacked galaxy measurements



Chisholm et al. 2015
48 galaxies
 $z \sim 0$
Si II absorption

Rubin et al. 2014
105 galaxies
 $0.3 < z < 1.4$
Mg II absorption

Roberts-Borsani et al. 2020
405 galaxies
 $z \sim 0$
NaD absorption

Davies et al. 2019
28 galaxies
 $2 < z < 2.6$
H α emission

Testing sub-grid physical models

SNe-driven:

$$v_{out} \propto \Sigma_{SFR}^{0.1}$$

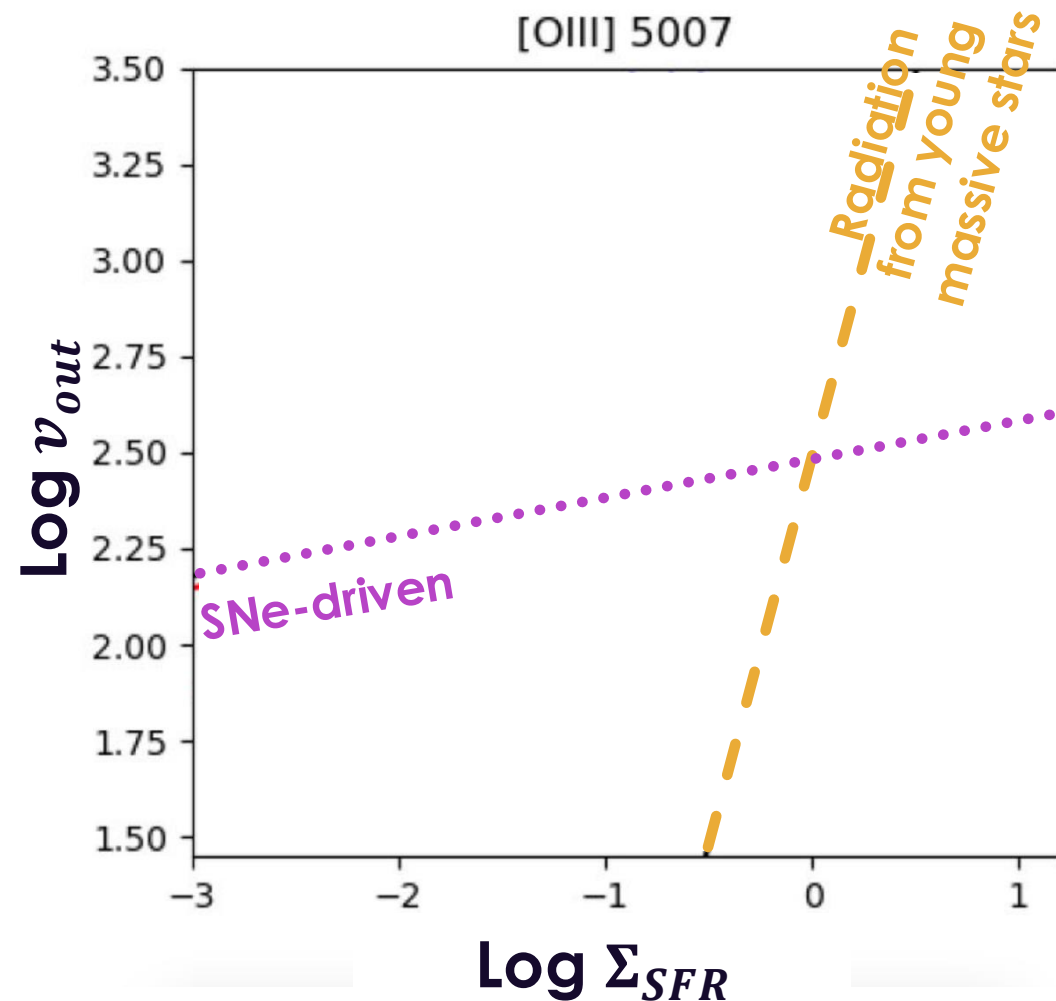
Chen et al. 2010, Li et al. 2017, Kim et al. 2020



Radiation from
young massive stars:

$$v_{out} \propto \Sigma_{SFR}^2$$

Murray et al. 2011



Supernovae are primarily driving the outflows in DUVET face-on galaxies



SNe-driven:

$$v_{out} \propto \Sigma_{SFR}^{0.1}$$

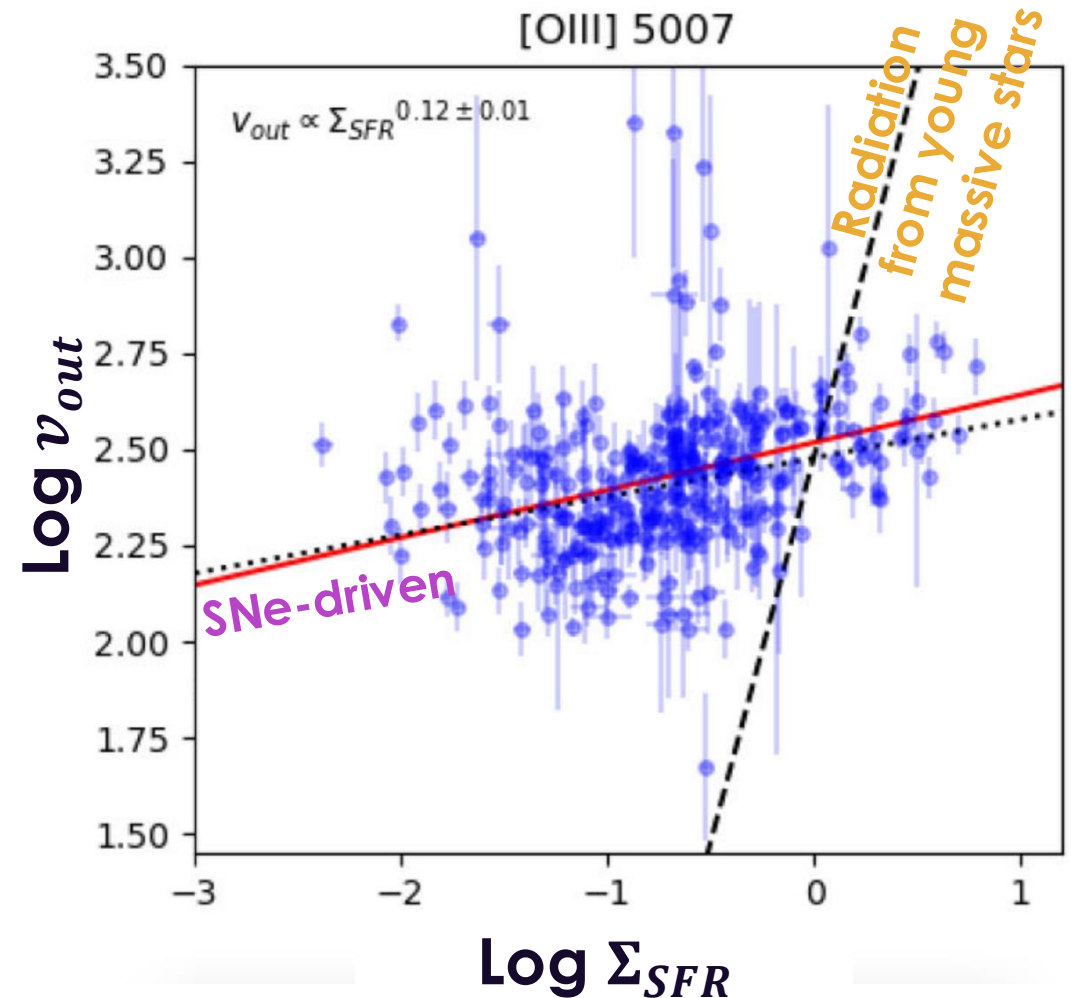
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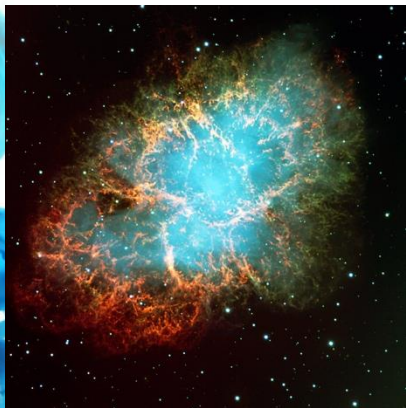
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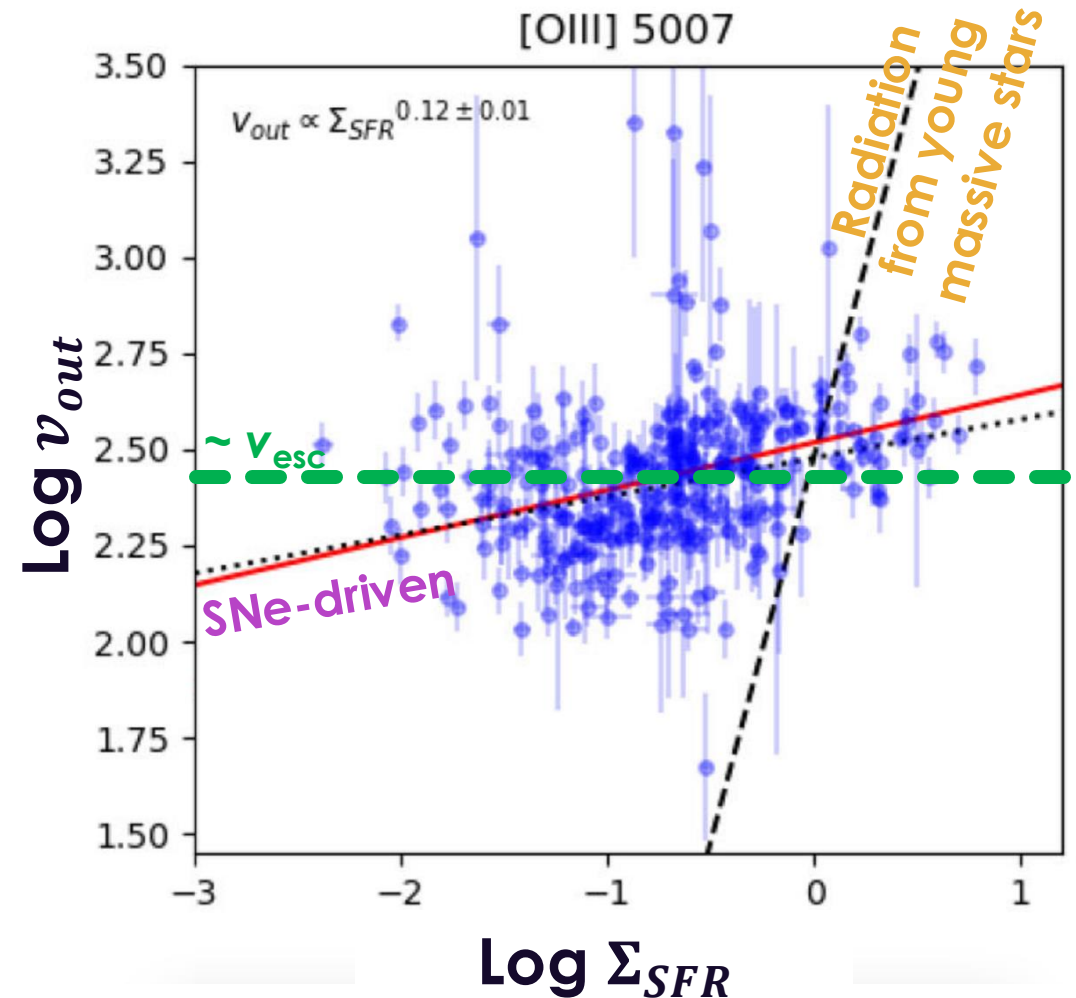
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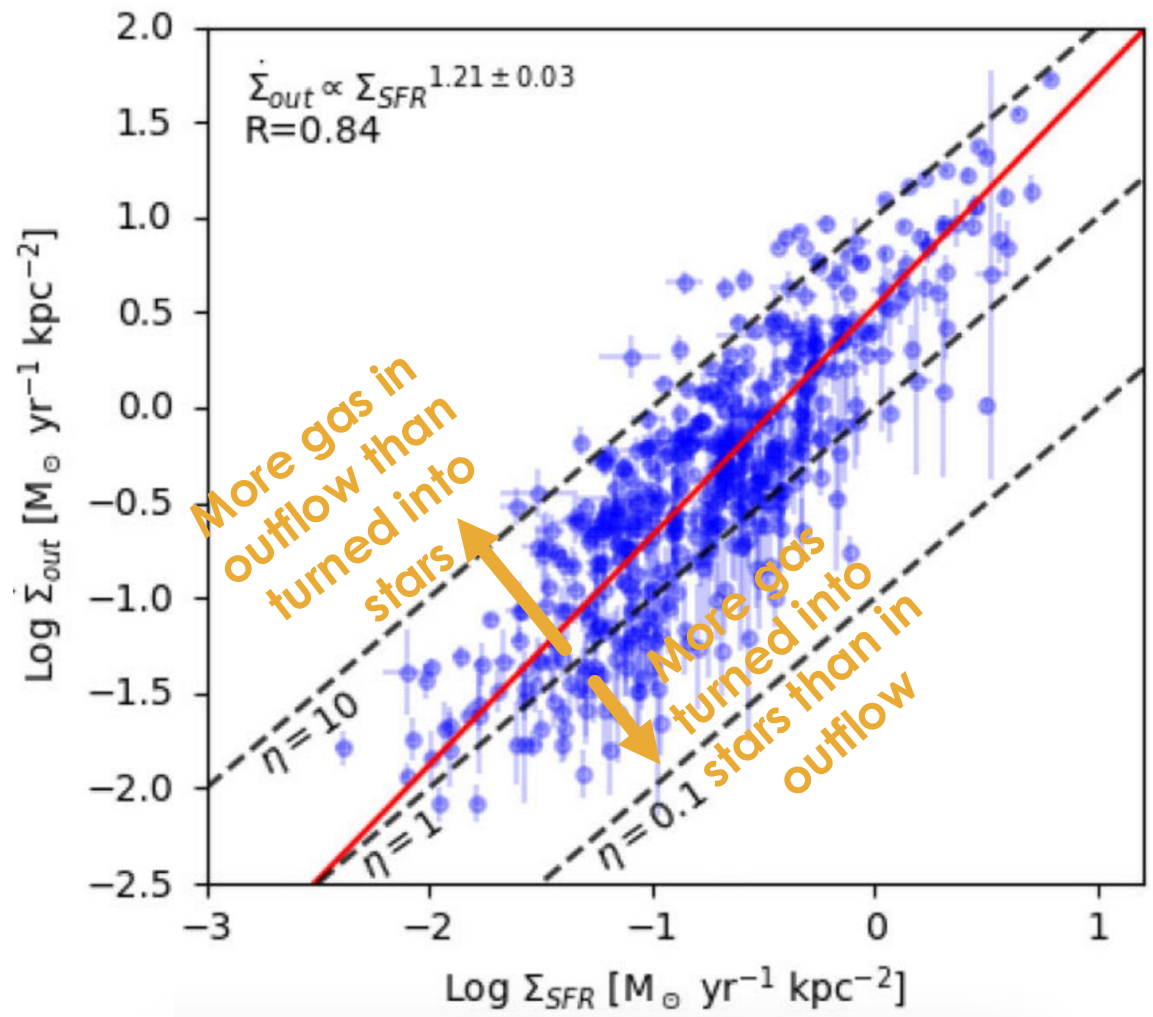


DUVET
Face-on

High Σ_{SFR} regions drive stronger outflows

$\dot{\Sigma}_{out}$
mass outflow
rate per area

Assumptions:
 $R_{out} \sim 500$ pc
 $n_e \sim 100$ cm⁻³



For a full table of
our results, see the
Supplementary
Files for the paper:



Reichardt Chu et al. (2025)

How do outflows affect the star formation?

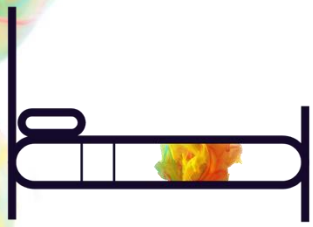
How do outflows affect the star formation?

Come to my talk tomorrow to find out!

Star formation across environments: From individual molecular clouds to entire galaxies

2:15pm – 3:45pm

TLC 113



DUVET
Edge-on



DUVET
Face-on



How are outflows distributed?

Outflows are integral to the metallicity cycle in galaxies

Hamel-Bravo+2024,
MNRAS 530, 3855

Cameron+2021,
ApJL 918, L16

Outflows in low-mass low-metallicity galaxies may not expel enough metals to follow MZR



How do outflows regulate star formation?



What mechanism is driving the outflow?

Outflow velocities consistent with SNe-driven models

Higher SFR surface density drives more gas out

Reichardt Chu et al. (2025)
MNRAS, 536, 1799

For a full table of our results, see the Supplementary Files for the paper



See my talk tomorrow!