

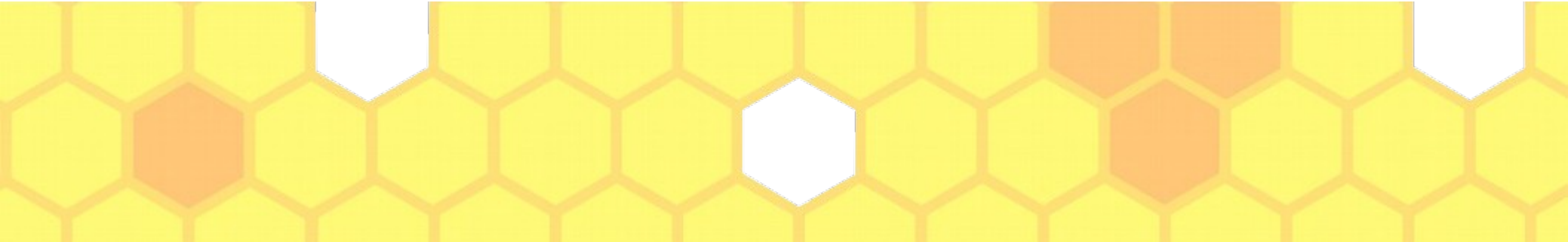
A decorative border at the top of the slide consisting of a grid of yellow hexagons with orange outlines. Some hexagons are solid orange, while others are white, creating a honeycomb-like pattern.

Dust in the Wind

circumnuclear dust
in nearby AGN imaged with JWST

David Rosario

Newcastle University

A decorative border at the bottom of the slide, similar to the top one, with a grid of yellow hexagons and orange outlines. It includes solid orange hexagons and white hexagons in a honeycomb pattern.

Rosario's JWST AGN team



Houda Haidar - 3rd year PhD



Steph Campbell - postdoc



Devang Liya – 3rd year PhD



**Manthos Charidis
MPhil (grad)**

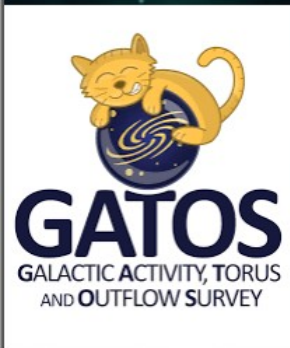


Daniel Hill – 1st year PhD

GATOS

Welcome News Science Collaborations
Members Contact

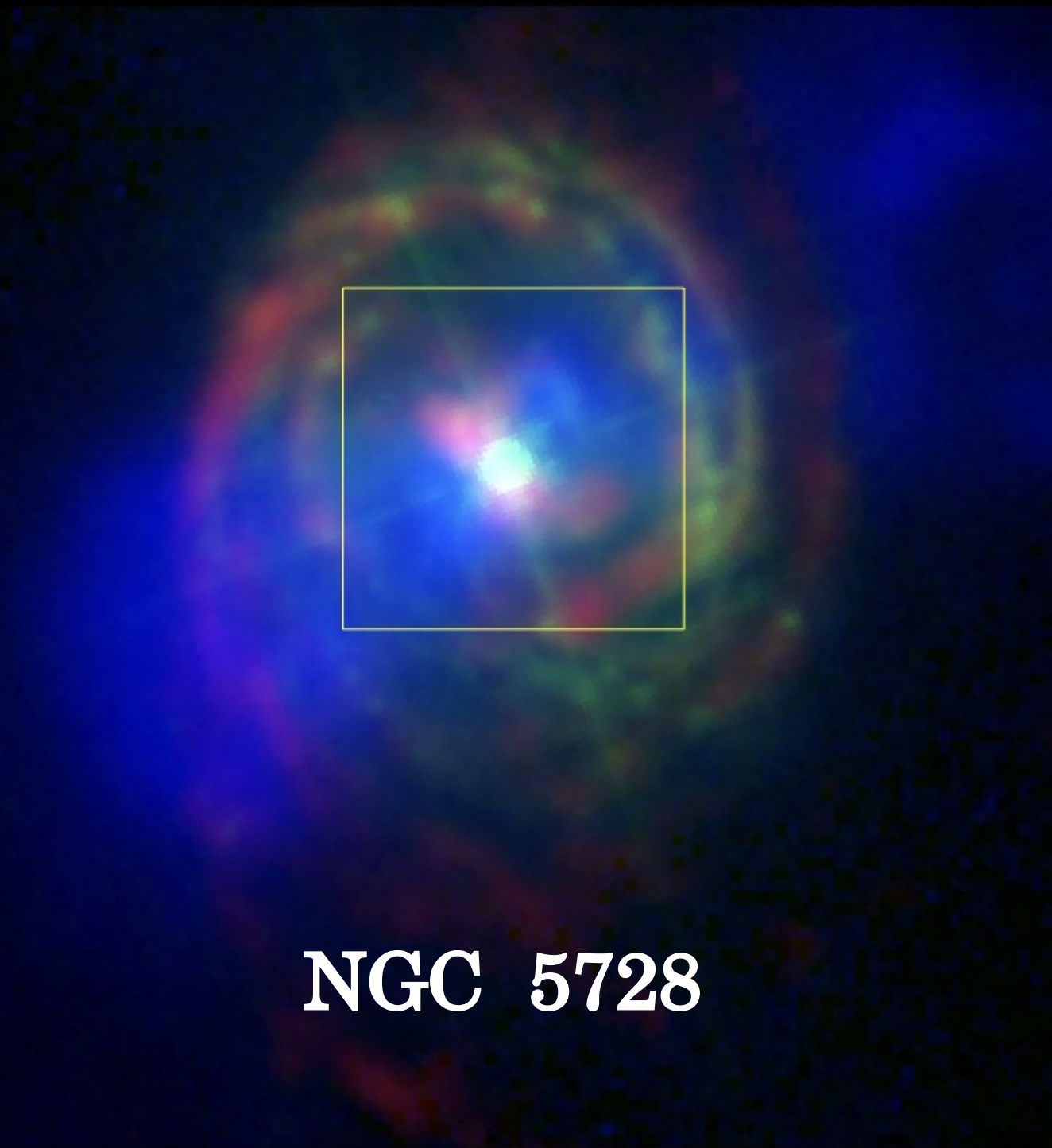
Image credit: ESA/Hubble, A. Riess et al., J. Greene



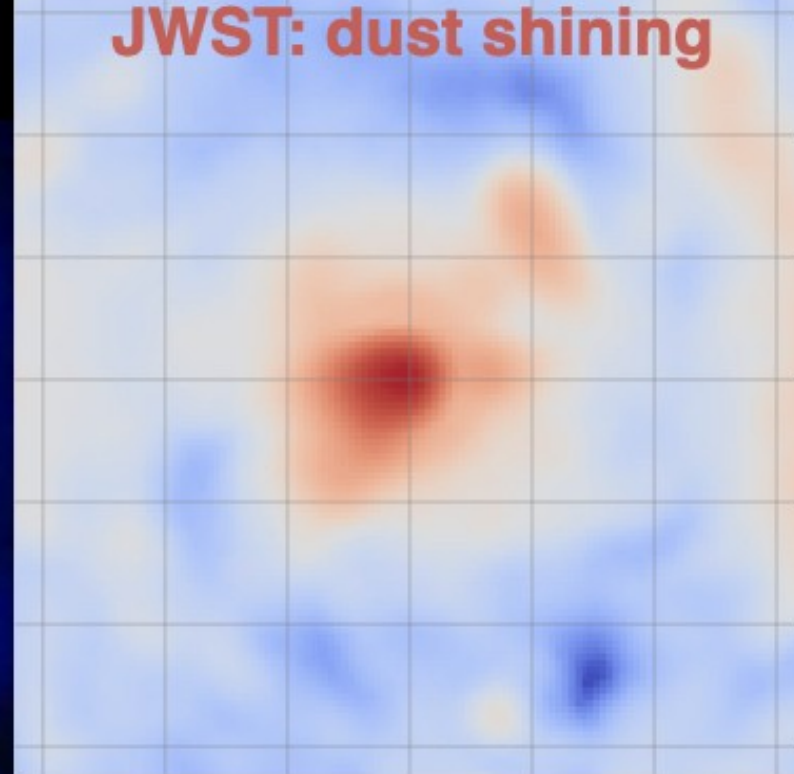
Galactic Activity, Torus, and Outflow Survey

We are an international team of observers, modellers and instrument specialists with the express aim of understanding the small-scale structure and dynamics of AGN, from the circumnuclear region (inner 100s of pc) to the accretion disk.

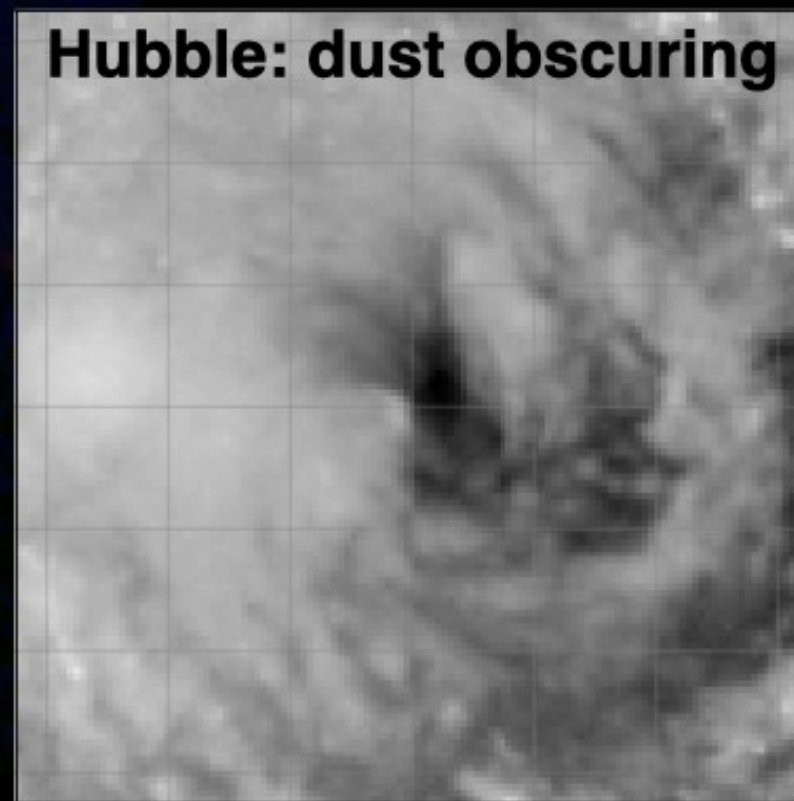
We use **ALMA, JWST, 8-10 mt class telescopes (VLT, GTC, Keck), SOFIA** and other facilities.



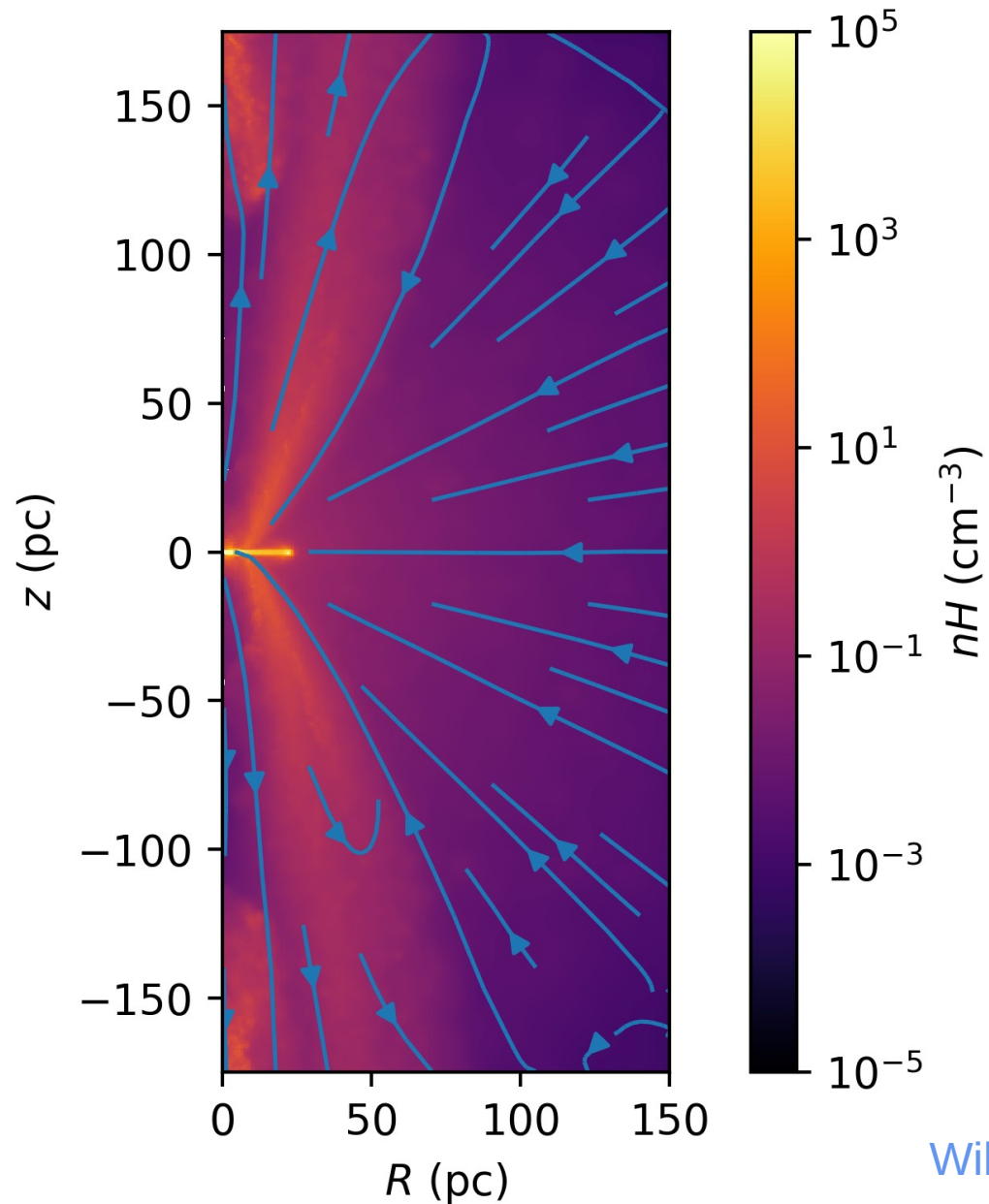
JWST: dust shining



Hubble: dust obscuring



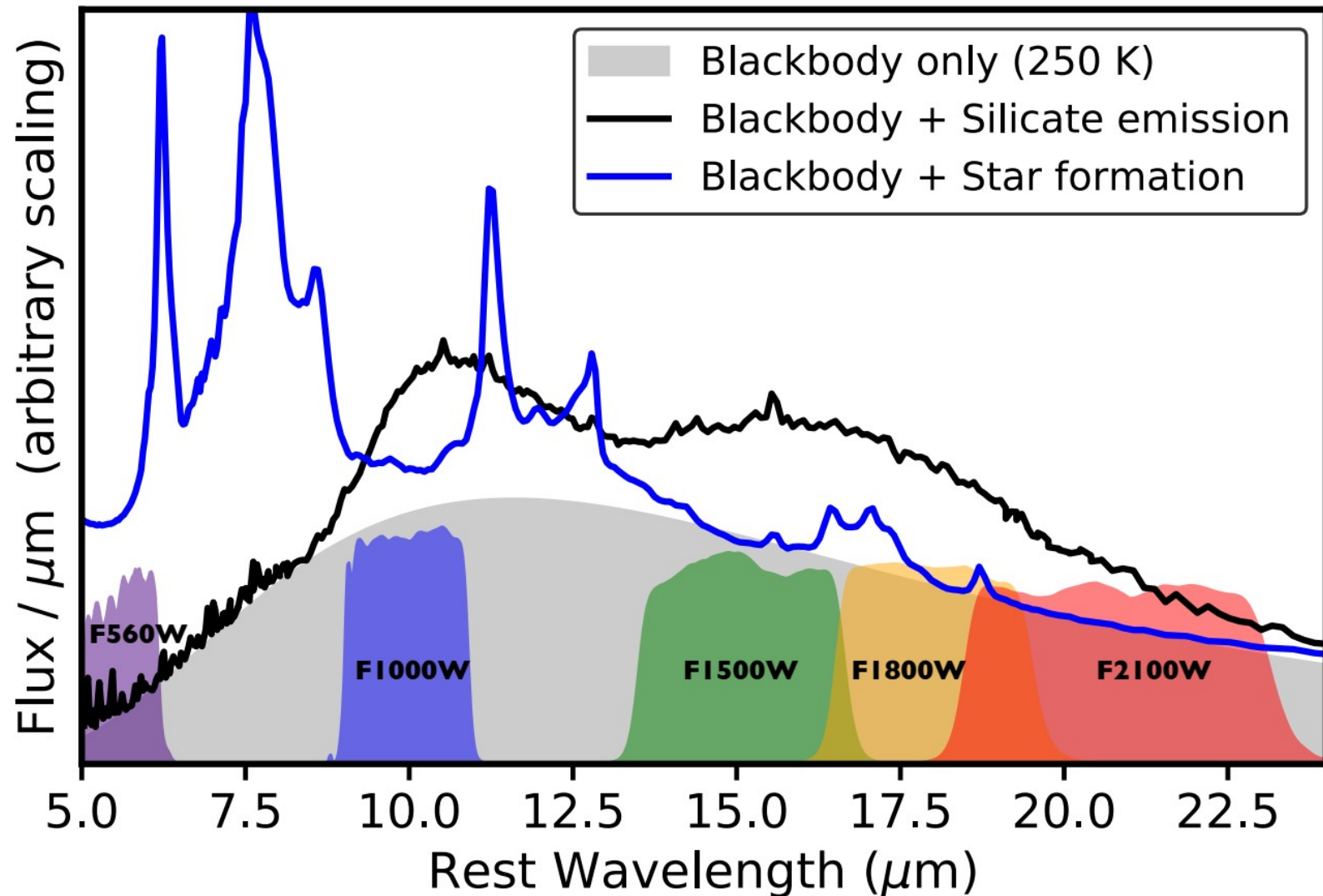
The Windy Torus



Williamson+ (2020)

Dust in the Wind

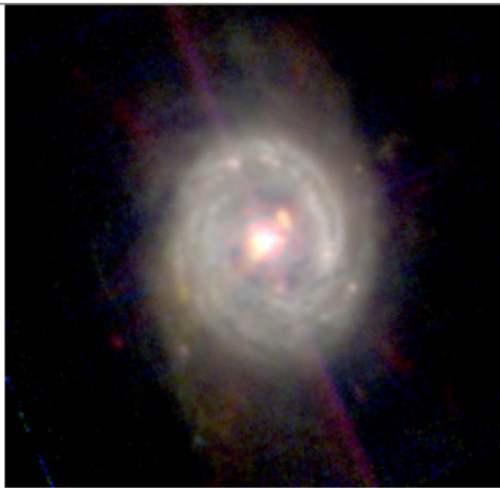
a JWST Cycle 1 program



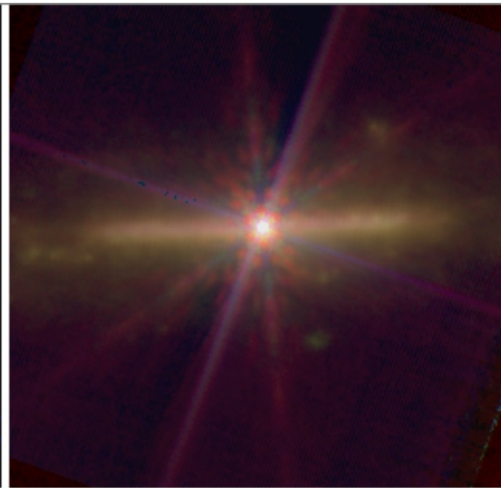
ESO 428-G14



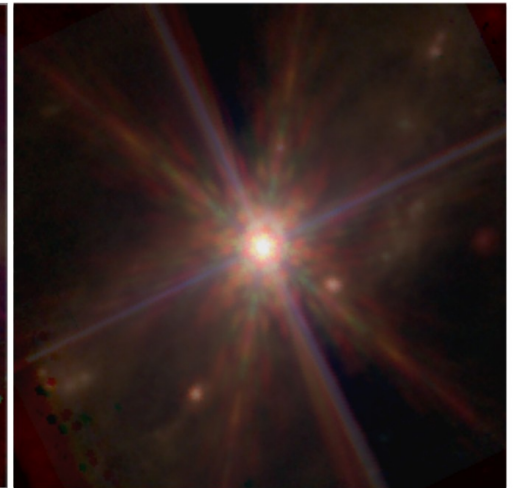
NGC 5728



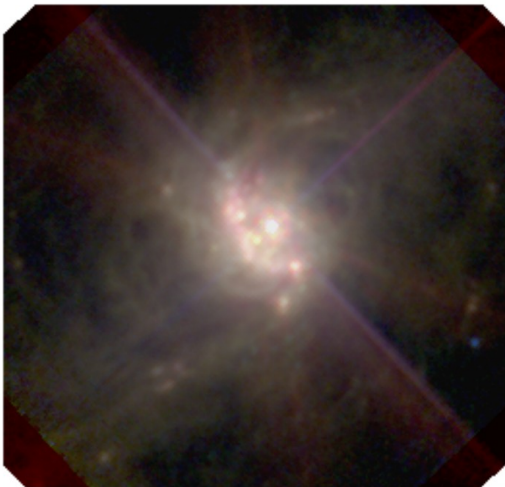
NGC 7172



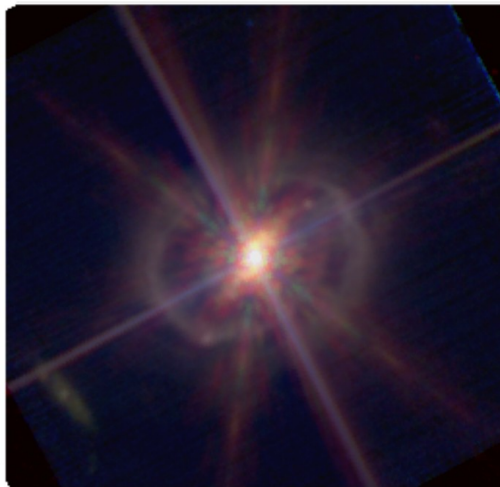
NGC 3227



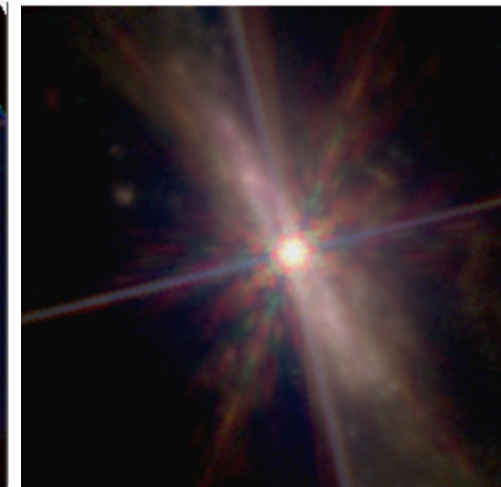
NGC 5135



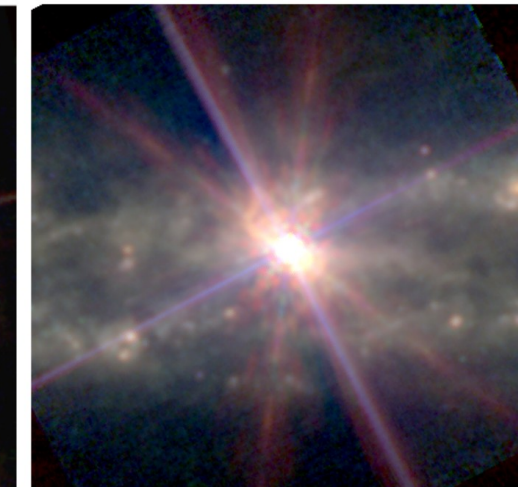
NGC 3081



NGC 2992



NGC 4388

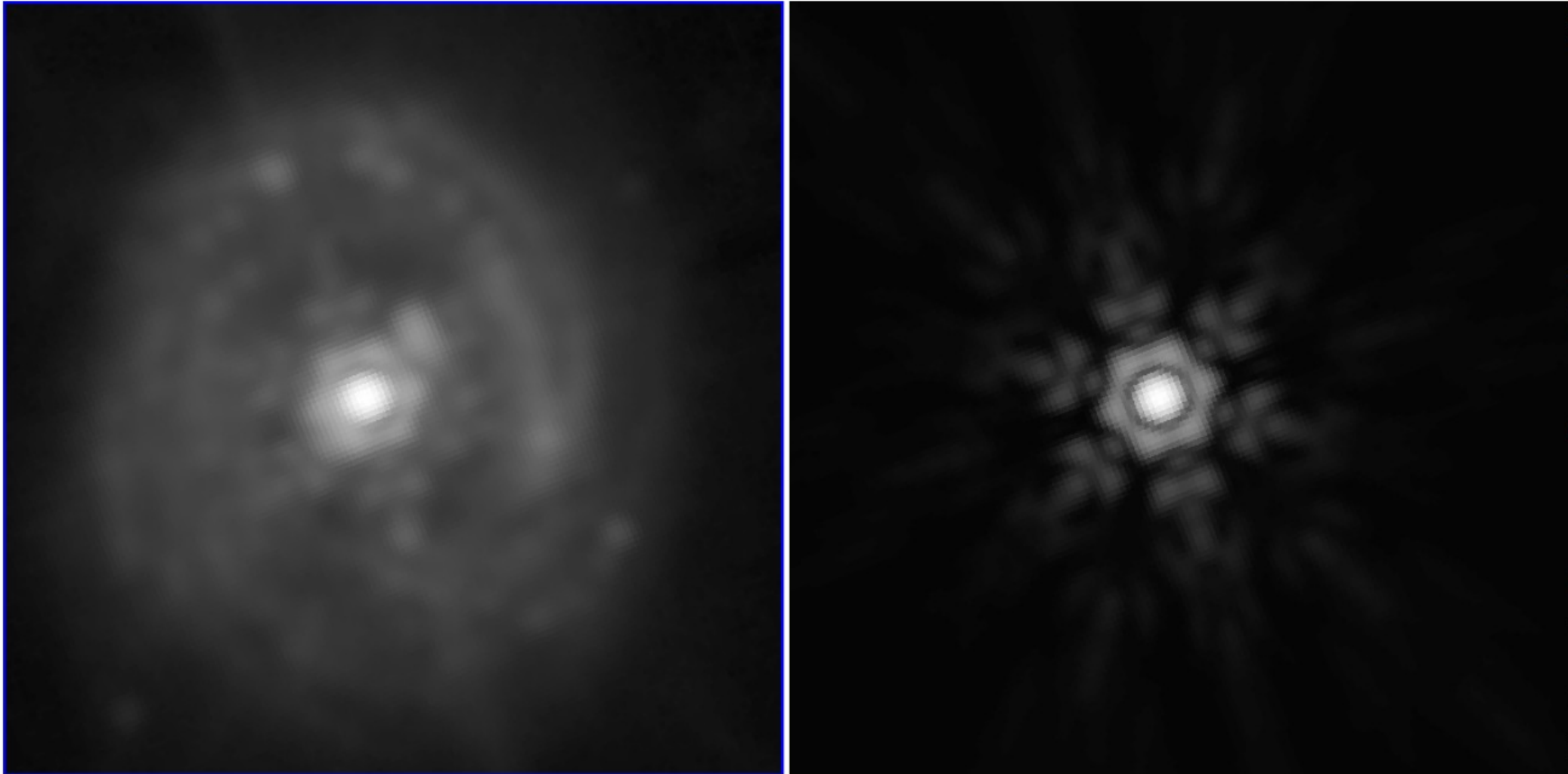


Three-colour full frame images (28" on a side)

15, 10, 5.6 microns

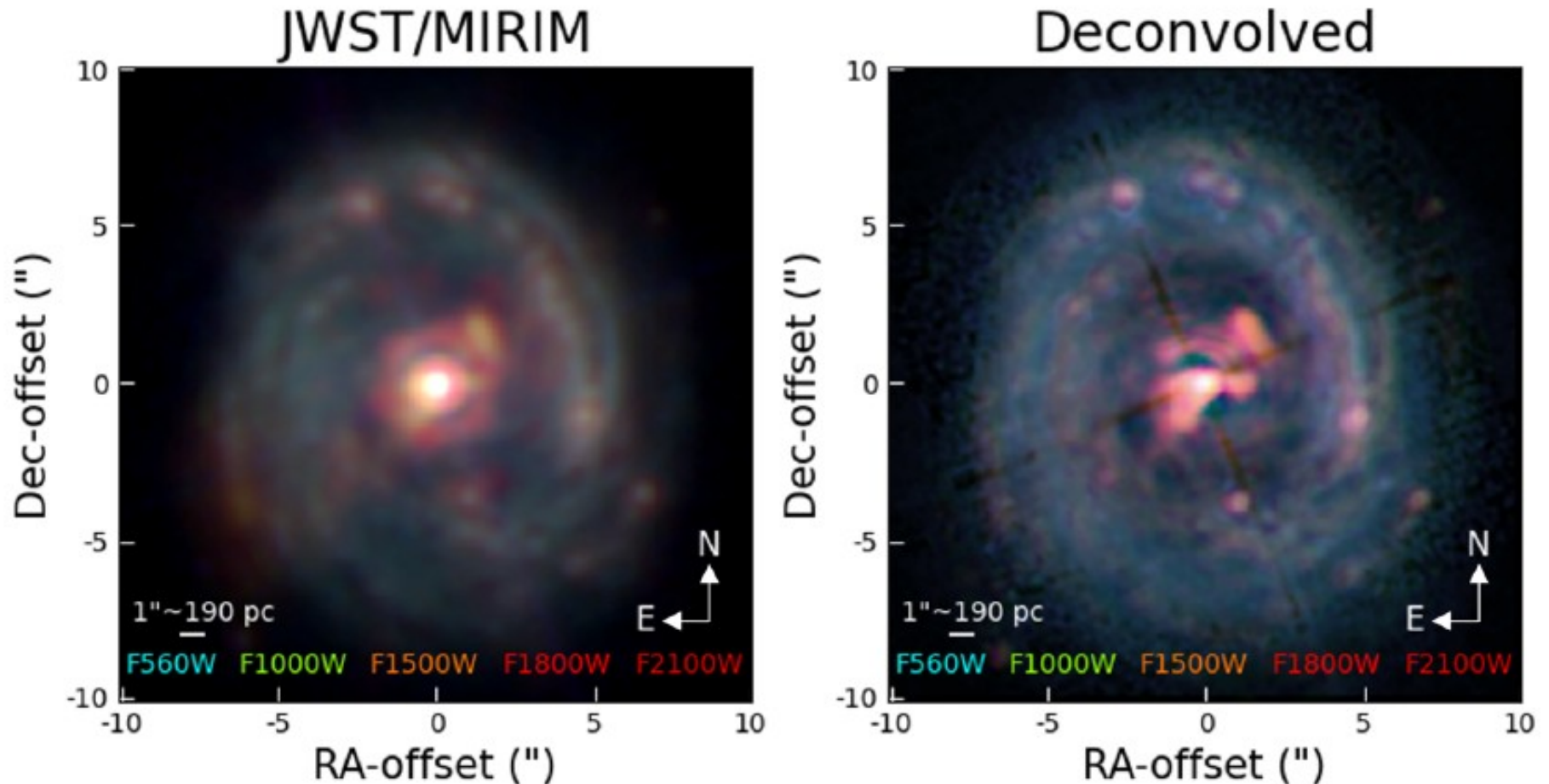
Model PSFs for handling the Point Source

Using the STPSF (formerly WebbPSF) package



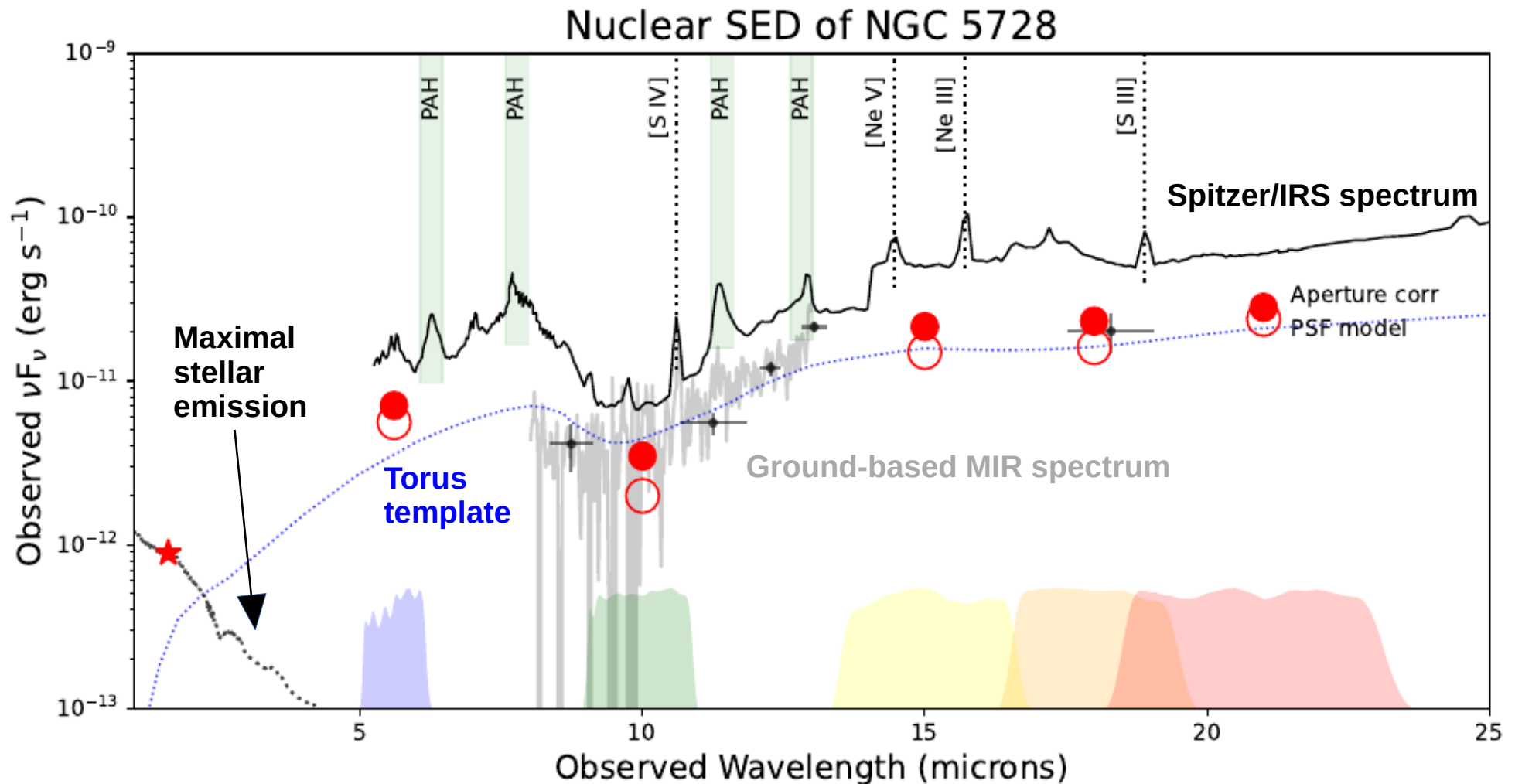
Deconvolution studies

Leist+ (2024)



A test of various modern deconvolution methods on our images. Currently, Kraken (a forward-modelling algorithm) and Richardson-Lucy work best.

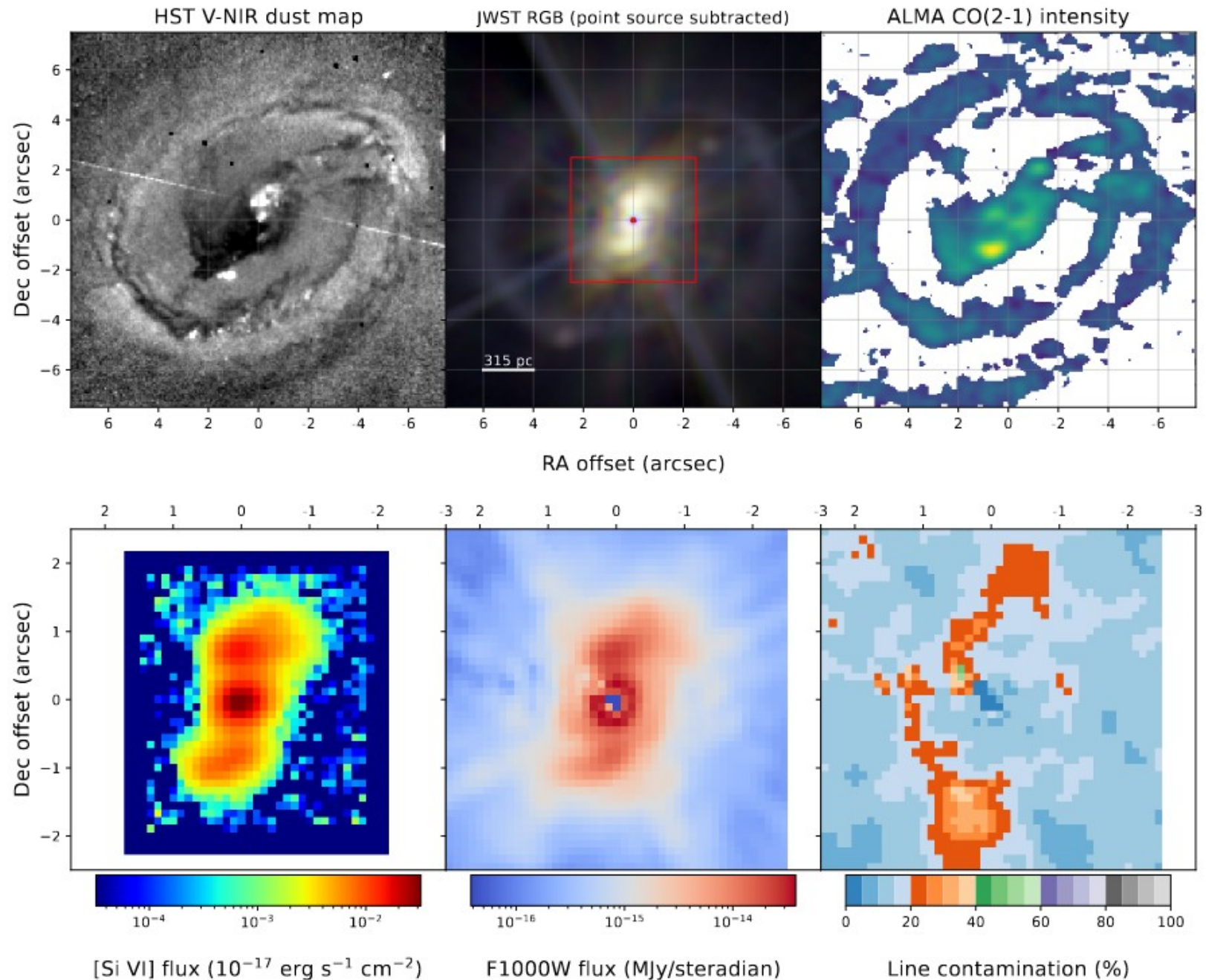
The nuclear point source

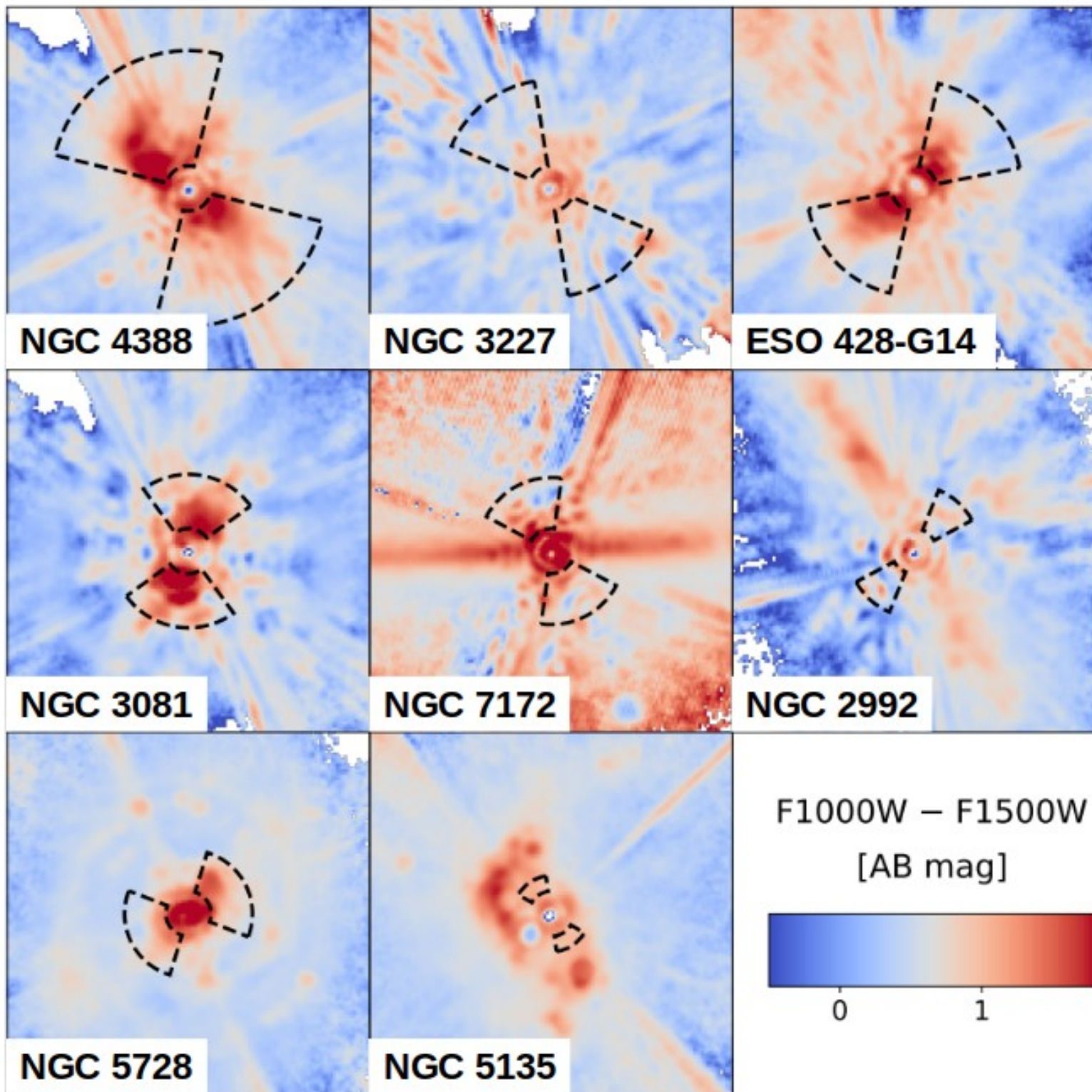


Red points: JWST (and HST) photometry

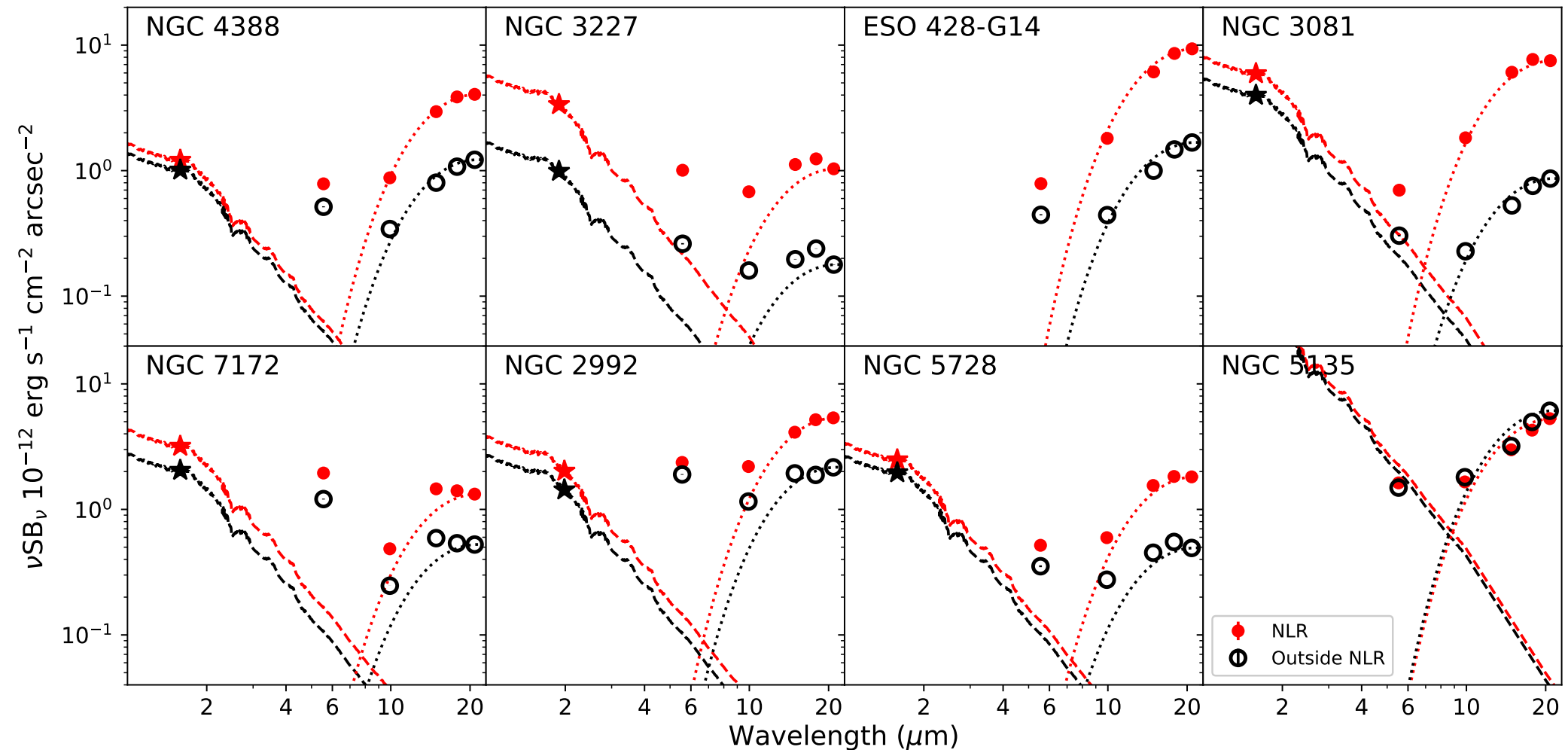
Black points: Ground-based MIR photometry

Multiphase gas and dust structure





The extended emission



AGN-illuminated regions have high surface brightnesses.

Stellar contamination can be important at wavelengths < 7 microns.

Summary

Our GATOS JWST/MIRI imaging program reveals that AGN-illuminated dust is widespread in nearby Seyferts, with good evidence for compact polar emission in most.

We've developed methods to recover and remove the bright nuclear point sources, which can serve as the basis for accurate nuclear photometry and deconvolution.

Analysis of the extended emission is on-going, but early results suggest normal ISM dust compositions and heating from sources that are not nuclear. Shocks?

Future GATOS data, both imaging and spectroscopy in the MIR, will allow a substantial sample study to place these dust structures in the context of the population of local Seyferts.